NB1: Preparing the data for analysis

Written by Maria Salem, in supplement to master thesis.

In this notebook we prepare all the datasets for analysis by reading in the data, interpolating values where necessary, redefining the data as probability distributions (KDE), and resample the data on a regular time grid.

At the end of the notebook, we prepare any relevant figures to include in the main text data chapter.

The records that will be used in this project are:

- 1. $\delta^{18}Q$
- Lisiecki & Raymo (2005) (denoted **LR04**). $\delta^{18}O$ global reference stack, spanning 5.3 Ma
- 1. Global sea-level (GSL)
- Spratt & Lisiecki (2016) (denoted SL). GSL stack, spanning 798 ka (PCA)
- Elderfield et al. (2012) (denoted E). GSL record, spanning 1.5 Ma (Mg/Ca temperature deconvolution of d180)
- Grant et al. (2014) (denoted G). Red Sea RSL record, spanning 500? ka.
- Rohling et al. (2014) (R). Mediterranean RSL record, spanning 5.3 Ma
- 1. insolation
- Laskar (2004) (denoted **La2004**). Numerical solution for northern hemisphere summer insolation (top of atmosphere solar flux for 21^{st} of June at 65°N). 50 Myr back and forth in time from present without significant uncertainty associated (uncertain beyond that).
- 1. pCO2
- Bereiter et al. (2015) (denoted B). pCO2 record from ice core at Epica Dome C, spanning ca 800 ka.
- Chalk et al. (2017) (denoted **C**). pCO2 record from $\delta^{11}B$ proxy, spanning ca 1.090-1.240 Ma.
- 1. Dust
- Lambert et al (2008) (denoted L). Dust record from dust concentration in ice core at Epica Dome C, spanning ca 800 ka. (Note, the data has been set to updated age model since publication.)
- Martinez-Garcia et al. (2011) (denoted **MG**). Record of iron mass accumulation rate (Fe MAR) in the Southern Ocean, spanning the last 4 Myr.

Outline for each dataset

- i) Read in data. Reference, read in and plot the data for a first overview.
- ii) Subtle redefinition from age to time. We reverse the rows rev to get arrays where time increases forward with index value (instead of age, increasing backwards). We then name the arrays, and make any necessary adjustments on units ets (f.ex. reformulation from years to kyrs (dividing by 1000). For one of the datasets (Lambert), we also redefine the age model.
- iii) Additional information on age model and uncertainties Since our method operates with a lead-lag definition of causality, it is very sensitive to age reversals, and it is crucial to include uncertainties in the time series' age models. For records where age model uncertainties are not included in the dataset, we discuss and include these. For the Lambert record, we also redefine the age model.
- iv) Discuss time series resolution, and interpolate data if needed. We will run our general analyses with a timestep of 1000 years (binsize = 1 kyr). Additionally, we discuss which time series may be suitable for higher resolution analyses. Since our method requires continuous records, we have to interpolate values for the datasets that have gaps in resolution that are larger than the resolution we will run our analysis on. This is to ensure we won't have any empty bins in the grid we will define later for our analysis (step vi). We use linear interpolation, to make as few assumptions as possible. The LinearInterpolation function from the *Interpolations* package makes a continuous function of the data where you can pick any time point to carry on. The arrays of interpolated values are marked with the prefix intp.
- v) Redefine the datasets as UncertainIndexValueDatasets (uivDs). This datatype from the UncertainData package carries on uncertainties in kernel density estimates (KDE). Resampling can then be done from the KDE, which is much faster and more accurate than running computationally heavy resampling loops.
- vi) **Binned resampling of the time series**. Having the time series on the exact same timesteps is crucial for our method to work. We therefore define a regular time grid onto which we bin the datapoints by resampling. This shuffles the uncertainty in age over to the value uncertainty / probability distribution. The <code>BinnedResampling</code> function resamples values from the KDE, and assigns one value in each bin (each timestep). We choose a binsize (time step) of 1000 years (1 kyr), in accordance with the Spratt & Lisiecki grid, and also this being roughly the order of resolution of out other records. For records with higher resolution, we may choose a finer binsize to run high resolution analyses. Which records are suitable for high resolution analysis and with what time step will be discussed below.

(Note: The *common time grid for analysis* is defined not only by the resolution but also by the common time interval of the time series pairs we want to analyze. Since the BinnedResampling function is computationally heavy and time consuming, we want to avoid running it again for each time series pair we want to analyse. We therefore decide on what will be the common *resolution* and do a binned resampling of the full length of each time series in this notebook. The common *time interval* will be selected later, and define the notebooks for analyis and results (*NBRs*). This saves us a lot of computational time, given that we have so many time series combinations and time intervals to run analyses over.

• vii) Save the wrangled data. The the binned resampled uivD time series are saved in a .jld2 file, so that we can smoothly load it in to the next notebooks where we run our analysis, without having to rerun the computations.

(Outline for our next notebooks)

For overview. Include here or in main text?

NB. OLD. Some repetition since I later moved BinnedResampling from NBRs to NB1.

The next steps for the wrangled data will be to get the time series on a common time grid. This is crucial for our method to work. We will therefore define a time grid onto which we bin the datapoints. The time grid is defined according to the length and resolution of the time series pairs we want to analyze. Once a common grid is defined, the time series will be binned on the grid by assigning a corresponding value to each bin. For data with associated uncertainty (uncertainties are carried in the uivD-objects), binning the data to the time grid is done by the function <code>BinnedResampling</code>. This function resamples the values within each bin. This generates a probaliblity distribution, allowing us to quantify the confidence for the mean value assigned to each bin. We will represent all our data with a 95% confidence envelope, for robust results.

Once the datasets are binned to a common time grid, we can run our analyses.

Note: The different time series combinations will require different common grids. We will therefore create different notebooks (NBRs) for the different common grids. These are the notebooks where we will compute our results (**NBR**, see below).

NB2

In notebook 2, we thoroughly go through all the steps in our analysis. We use the example of an autoregressive system of the first order, from which we create two time series realizations X and Y, and show how Predictive Asymmetry is estimated between timeseries data. The notebook walks through all the steps in the analysis, with abundant comments and explanatory figures. The aim is to give a general understanding of predictive asymmetry and of what is done in the following notebooks (NB3 and NBRs).

(After synthetic time series are represented, include note on reflections/challenges for using real-world empirical time series. -introduce binned resampling here instead?)

NB₃

In notebook 3, we summarize the the code in notebook 2 into functions. This allows us to run the analysis between the many time series more efficiently.

NBRs

The results notebooks (NBR) is where we run the analyses between our time series data and plot our results. We include the functions for analysis defined in NB3, and the time series prepared in this notebook (NB1). We make a separate NBR for every grid (time interval) we will run analysis over. Each time grid is constrained primarily by the timespan and resolution of the available records. (ice core records, for example, only cover the last 800 kyrs). We also define time span constraints on time grids that allow us to check if dynamical coupling differs over the hypothesized periods post-MPT (0-800 kyr BP), syn-MPT (800-1250 kyr BP) and pre_MPT ((1250-4000 kyr BP).

Loading in the time series data and wrangle data

First, we load in the necessary packages/libraries

```
In [2]:
```

```
# loading in the necessary packages/libraries
Pkg, # For adding the libraries we will use
Distributions, # for probability distributions and associated functions
DelimitedFiles, # For reading in dataset files that are .tab or .csv or .txt
Plots; # for making the figures
pyplot() # running pyplot in backend to plots (not as pretty as GR, but faster,
 and compatible with LaTeXStrings)
using
LaTeXStrings, # For writing special symbols and characters
JLD2, # For saving the arrays we want to carry on in different notebooks
DataInterpolations # For linear interpolations on time series with discontinuous
or insufficient resolution
using
UncertainData, # for carrying on uncertainties as kernel density estimates (type
UncertainIndexValueDatasets), and BinnedResampling
# maybe not used here?
CausalityTools, # ...???
DynamicalSystems # ...??? A Julia suite for chaos and nonlinear dynamics
# not sure if I use these at all?
using
StatsBase, # ???
Test, # ???
XLSX # For reading in dataset files written in excel
```

In [2]:

1. LR04

- · Lisiecki & Raymo (2005), denoted LR04
- Global $\delta^{18}O_b$ stack, spanning the last 5.32 Myrs.
- Both the LR04 age model and $\delta^{18}O$ serve as reference stacks in paleoclimatologic sciences.
- Data available from Pangaea, DOI: https://doi.org/10.1594/PANGAEA.704257)
- NB: the LR04 age model is constructed using orbital tuning. This may create bias in our results in analyses between this record and insolation time series.

i) Read in the data

In [33]:

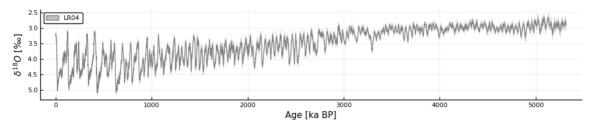
```
# Read in dataset from Lisiecki & Raymo
filepath_LR04 = "../../MASTER_2.0/data/sea-level/LR04/Lisiecki_Raymo_2005/datase
ts/Global_stack_d180.tab"
rawD_LR04 = readdlm(filepath_LR04, skipstart = 58)
;
```

In [34]:

```
# name columns
age_LR04_ = rawD_LR04[:,1]  # Age [ka BP]
d180_LR04_ = rawD_LR04[:, 2]  # \delta180 [%]
d180_l\u03c3_LR04_ = rawD_LR04[:,3]  # standard deviation [±]
;
```

In [36]:

```
# plot LR04 record
plot_LR04_raw_age =
plot(age_LR04_, d180_LR04_,
    label = "LR04", color = :gray,
    ribbon = 2*d180_1o_LR04_,
    yflip = true, # by convention, we plot d180 on a reversed axis #(mirror read
ability of GSL, which constitutes on average 70% of the signal)
    xlabel = "Age [ka BP]",
    ylabel = L"\delta^{18}0 \ [\perthousand]",
    size = (1000,200))
savefig("../../MASTER_2.0/figurar/RawData/LR04_age.pdf")
```



ii) **Subtle redefinition of record as time series**. We reverse dataset to go from age, increasing backwards, to time, increasing forwards. To keep sensical index values, e define present as 0 kyrs BP.

In [37]:

```
# Reverse dataset to go from age increasing backwards to time increasing forward s.

revD_LR04 = reverse(rawD_LR04, dims = 1)

# name columns

t_LR04 = -revD_LR04[:,1]  # Age [ka BP] # negative, to define present as 0 k

yrs

d180_LR04 = revD_LR04[:, 2]  # &180 [%]

d180_1\sigma_LR04 = revD_LR04[:,3] # standard deviation [±]

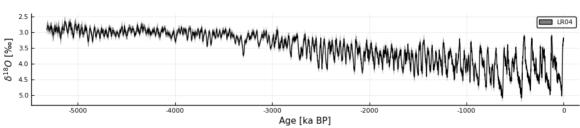
;
```

In [38]:

```
# plot LR04 time series
plot_LR04_raw_time =
plot(t_LR04, d180_LR04,
    label = "LR04",
    color = :black,
    ribbon = 2*d180_1\sigma_LR04,
    xlabel = "Age [ka BP]",
    ylabel = L"\delta^{18}0 \ [\perthousand]",
    yflip = true, # by convention, we plot d180 on a reversed axis #(mirror read ability of GSL, which constitutes on average 70% of the signal)
    size = (1000,200)
    )

#savefig("../figurar/RawData/LR04_time.pdf")
```

Out[38]:



iii) Potential systematic uncertainties in age model

Since our method operates with a lead-lag definition of causality, it is very sensitive to age reversals, and it is crucial to include uncertainties in the time series' age models.

Lisiecki & Raymo (2005) report the LR04 age model uncertainties as following:

"Including all sources of error, we estimate the uncertainty in the LR04 age model to be 40 ky from 5.3-5 Ma, 30 ky from 5-4 Ma, 15 ky from 4-3 Ma, 6 ky from 3-1 Ma, and 4 ky from 1-0 Ma".

Kommentar frå Bjarte: kanskje overdreven usikkerhet. Men me bruker som dei oppgir i første omgang (konservativt), og kan evt kjøre sensitivitetstest seinare med mindre usikkerhet.

We interpret this as a 95% confidence interval $(\pm 2\sigma)$, and define an array to carry on the age uncertainty:

In [39]:

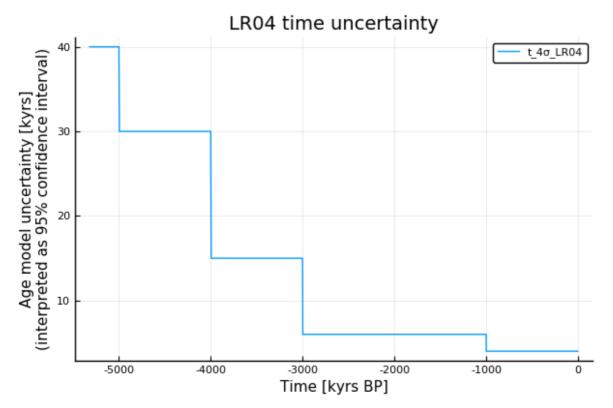
In [40]:

```
# we now have all the arrays we need to produce the raw plot, lets save them in a .jld2 file for easy access
@save "../Koding/WrangledDataFiles/BasicArrays/LR04.jld2" t_LR04 t_1\sigma_LR04 d180_LR04 d180_1\sigma_LR04
```

In [45]:

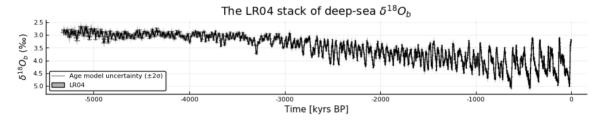
```
plot(title = "LR04 time uncertainty",
    t_LR04, t_40_LR04, # Showing stepwise definition of age uncertainty, where u
ncertainty increases back in time. All good.
    xlabel = "Time [kyrs BP]",
    ylabel = "Age model uncertainty [kyrs]
(interpreted as 95% confidence interval)",
    label = "t_40_LR04")
```

Out[45]:



In [51]:

```
# plot the LR04 time series
plot LR04 raw timeunc =
plot(title = string("The LR04 stack of deep-sea ", L"$\delta^{18}0 {b}$"),
    label = "LR04",
    size = (1000, 200),
    xlabel = "Time [kyrs BP]",
    ylabel = L"\delta^{18}0 {b} \ (\perthousand)",
    yflip = true # by convention, we plot d180 on a reversed axis #(mirror reada
bility of GSL, which constitutes on average 70% of the signal)
plot!(t_LR04, d180 LR04,
    xerr = (2*t 1σ LR04, 2*t 1σ LR04), # age uncertainty, what we have interpret
ed as 95% confidence interval
    ms=1, color = :grey,
    label = "Age model uncertainty (\pm 2\sigma)"
plot!(t LR04, d180 LR04,
    ribbon = (2*d180 1σ LR04, 2*d180 1σ LR04), # plotting with 2σ (95% confidence
interval)
    color = :black, fillalpha = 0.3,
    label = "LR04"
savefig("../../Master 2.0/figurar/RawData/LR04 timeunc.pdf") # NB currently save
d with the plot!(xerr)
```



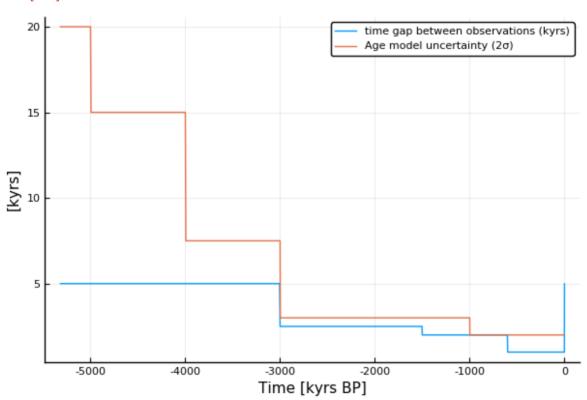
Interpolation. In the final analysis we use a 1000 year resolution grid (defined by SprattLisiecki grid), and we cannot have any empty bins in our final analysis. We therefore need to check that our data has at least 1000 years resolution, and if not, we need to interpolate values where there is missing information.

In [54]:

```
# Check if need for interpolation?
print("Smallest time gap between observations is ", minimum(diff(t_LR04)), " kyr s.
")
print("Mean time gap between observations is ", mean(diff(t_LR04)), " kyrs.
")
print("Largest time gap between observations is ", maximum(diff(t_LR04)), " kyr s.
")
plot(t_LR04, diff(t_LR04),
xlabel = "Time [kyrs BP]",
ylabel = "[kyrs]",
label = "time gap between observations (kyrs)")
plot!(t_LR04, 2*t_1\subsetcharpoolume{"LR04, label = "Age model uncertainty (2\sigma)")
```

Smallest time gap between observations is $1.0~\rm kyrs$. Mean time gap between observations is $2.5165562913907285~\rm kyrs$. Largest time gap between observations is $5.0~\rm kyrs$.

Out[54]:



WARNING: both StateSpaceReconstruction and DynamicalSystems export "dimension"; uses of it in module CausalityTools must be qualified

t_LR04 # ends at present (0 yrs BP), with 1 kyr between datapoints. So why the big gap at the end here?

```
In [103]:
```

```
\# Question about the plot t_LR04 ;\# ends at present (0 yrs BP), with 1 kyr . So why the big gap at the end here?
```

Based on the ensemble of records we want to analyse, we have chosen to use millennial resolution as default for our analysis.

As visualized in the plot above, the LR04 record has segments with varying resolutions (each segment on different regular grid). In about half the record, there is up to 5 kyrs between each datapoint in the LR04 record. Only in the verymost recent part of the record do we have millenial resolution.

By including the age model uncertainty (which is larger than the time gaps of the record), the binned resampling will still assign one value to every bin of the 1 kyr grid.

If we choose to not include the age model uncertaint, we will need to interpolate values for the LR04 record. To make the least possible assumptions, we use linear interpolation between the datapoints in the record.

In [9]:

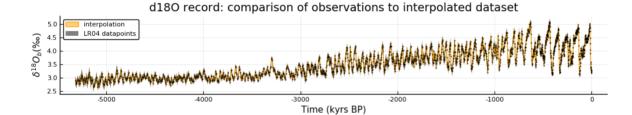
```
# interpolation
# create a continuous function with linear interpolation between every datapoint
in the array.
# This allows us to choose the interpolated value for any point in time.
                                                                     # time array
interpolate t LR04 = LinearInterpolation(t LR04, t LR04)
interpolate mean LR04 = LinearInterpolation(d180 LR04, t LR04)
                                                                    # mean d180 v
alue
interpolate 1\sigma LR04 = LinearInterpolation(d180 1\sigma LR04, t LR04)
                                                                    # d180 value
uncertainties
interpolate t 1σ LR04 = LinearInterpolation(t 1σ LR04, t LR04)
                                                                    # age model u
ncertainty (full)
#interpolate_t_1\sigma_LR04_trunc = LinearInterpolation(t_1\sigma_LR04_trunc, t_LR04) # ag
e model uncertainty, truncated at 1 sigma
# make a fine grained grid to contain the interpolated values
binsize intp = 0.1
fine grid LR04 = ceil(minimum(t LR04)) : binsize intp : floor(maximum(t LR04)) #
One bin for every 100 years (0.1 kyrs)
print(fine grid LR04)
# make new arrays for interpolated data
# give a value from the interpolate function to every bin in the fine grid
intpD t LR04 = [interpolate_t_LR04(i) for i in fine_grid_LR04]
intpD mean LR04 = [interpolate mean LR04(i) for i in fine grid LR04]
intpD 1\sigma LR04 = [interpolate 1\sigma LR04(i) for i in fine grid LR04]
intpD t 1σ LR04 = [interpolate t 1σ LR04(i) for i in fine grid LR04]
#intpD t 1\sigma LR04 trunc = [interpolate t 1\sigma LR04 trunc(i) for i in fine grid LR0
4] # Error here, so we truncate below
#intpD t 10 LR04 trunc = Truncated.(Normal.(intpD t LR04, intpD t 10 LR04), intp
D t LR04 .- intpD t 10 LR04, intpD t LR04 .+ intpD t 10 LR04)
;
```

-5320.0:0.1:0.0

In [105]:

```
# plot to check
plot(title = "d180 record: comparison of observations to interpolated dataset",
    size = (1000, 200),
    xlabel = "Time (kyrs BP)",
    ylabel = L"\delta^{18}0_{b} (\perthousand)")
plot!(intpD t LR04, intpD mean LR04,
    ribbon = (2 * intpD_1\sigma_LR04),
    color = :orange,
    ms = 0.1,
    label = "interpolation")
scatter!(t_LR04, d180_LR04,
    ribbon = (2 * d180_1\sigma_LRO4),
    color = :black,
    ms = 0.5,
    label = "LR04 datapoints")
# We see that the original time series and the interpolated time series overlap,
so all good
```

Out[105]:

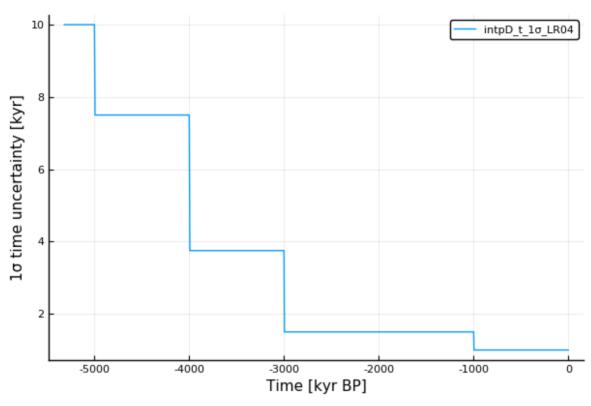


In [84]:

```
plot(#title = "LR04 time uncertainty - interpolated array",
    intpD_t_LR04, intpD_t_1\sigma_LR04,
    xlabel = "Time [kyr BP]",
    ylabel = "1\sigma time uncertainty [kyr]",
    label = "intpD_t_1\sigma_LR04")

# We see from the figure that the interpolated time uncertainty array is the sam
e as the original one we made above, so all good.
```

Out[84]:



Redefine as an UncertainIndexValueDataset

The type object <code>UncertainIndexValueDataset</code> from the <code>UncertainData</code> package is a format to carry on data with associated uncertainties in one single array. The uncertainties are carried on as kernel density estimates (KDE). This is less computationally heavy than the traditional for-loop resampling. This datatype is seamlessly integrated in the <code>DynamicalSystems</code> and <code>CausalityTools</code> packages, which we will use onwards in the NBRs.

In [55]:

```
# LR04 with maximum age uncertainty

d180_uiv_LR04 = [UncertainValue(Normal, d180_LR04[i], d180_1\sigma_LR04[i]) for i in
1:length(d180_LR04)];
t_uiv_LR04_fullageunc = [UncertainValue(Normal, t_LR04[i], t_1\sigma_LR04[i]) for i i
n 1:length(t_LR04)]
uivD_LR04_fullageunc = UncertainIndexValueDataset(t_uiv_LR04_fullageunc, d180_ui
v_LR04)
```

Out[55]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 2115 uncertain values coupled with 2115 uncertain ind
ices

In [85]:

```
# LR04 with maximum age uncertainty

d180_uiv_LR04_intp = [UncertainValue(Normal, intpD_mean_LR04[i], intpD_10_LR04[i
]) for i in 1:length(intpD_mean_LR04)];
t_uiv_LR04_fullageunc_intp = [UncertainValue(Normal, intpD_t_LR04[i], intpD_t_10
_LR04[i]) for i in 1:length(intpD_t_LR04)]
uivD_LR04_fullageunc_intp = UncertainIndexValueDataset(t_uiv_LR04_fullageunc, d1
80_uiv_LR04)
```

Out[85]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 53201 uncertain values coupled with 53201 uncertain i
ndices

#plot(uivD LR04 fullageunc)

InterruptException:

Stacktrace:

- [1] _update_series_attributes!(::Dict{Symbol,Any}, ::Plots.Plot{Plots.PyPlotBackend}, ::Plots.Subplot{Plots.PyPlotBackend}) at /Users/maria/.julia/packages/Plots/qZHsp/src/args.jl:1598
- [2] _process_seriesrecipe(::Plots.Plot{Plots.PyPlotBackend}, ::Dict {Symbol,Any}) at /Users/maria/.julia/packages/Plots/qZHsp/src/pipeline.jl:403
- [3] _process_seriesrecipe(::Plots.Plot{Plots.PyPlotBackend}, ::Dict {Symbol,Any}) at /Users/maria/.julia/packages/Plots/qZHsp/src/pipeline.jl:417
- [4] _plot!(::Plots.Plot{Plots.PyPlotBackend}, ::Dict{Symbol,Any},
 ::Tuple{UncertainIndexValueDataset{UncertainIndexDataset,UncertainVa
 lueDataset}}) at /Users/maria/.julia/packages/Plots/qZHsp/src/plot.j
 1:234
- [5] #plot#138(::Base.Iterators.Pairs{Union{},Union{},Tuple{}},NamedT
 uple{(),Tuple{}}},::typeof(plot),::UncertainIndexValueDataset{Unce
 rtainIndexDataset,UncertainValueDataset}) at /Users/maria/.julia/pac
 kages/Plots/qZHsp/src/plot.jl:57
- [6] plot(::UncertainIndexValueDataset{UncertainIndexDataset,Uncerta
 inValueDataset}) at /Users/maria/.julia/packages/Plots/qZHsp/src/plo
 t.jl:51
 - [7] top-level scope at In[109]:1

The Age model uncertainties reported (from tie points, not included in the original dataset) were quite substantial, and may overestimate the uncertainty (...). Therefore, let's make three versions of the record:

- one without age uncertainty,
- one including the 95% confidence interval,
- and one truncated at 1σ.

As a starting point, we will use (truncated? full? none?)

We can later do a sensitivity analysis, to check how the different uncertainty levels (?) impact the causality results.

In [56]:

```
# LR04 without age model uncertainty
t_uiv_LR04_noageunc = [UncertainValue(Normal, intpD_t_LR04[i], 0) for i in 1:len
gth(intpD_t_LR04)]
uivD_LR04_noageunc = UncertainIndexValueDataset(t_uiv_LR04_noageunc, d180_uiv_LR
04)
```

UndefVarError: intpD_t_LR04 not defined

Stacktrace:

[1] top-level scope at In[56]:1

Should we run our analysis with full uncertainty? Or truncate?

Useful to do both, with sensitivity analysis, to check what impact it would have on the results

Eksempel frå Kristian på korleis ta med aldersusikkerhet. normalfordeling = Truncated(Normal(2, 0.2), 2-0.2, 2+0.2) # (normalfordeling(mean, 1σ), trunkert slik at nedre grense, øvre grense = +-1sigma) # rand(normalfordeling) # sjekk at trunkering blir som forventa (2 +/- 0.2)# LR04 with age uncertainty truncated at +/-1σ intpD_t_1σ_LR04_trunc = Truncated.(Normal.(intpD_t_LR04, intpD_t_1σ_LR04), # Trunker(normalfordeling(mean, 1σ) intpD_t_LR04 .- intpD_t_1σ_LR04, # trunkert slik at nedre grense = mean - 1σ intpD_t_LR04 .+ intpD_t_1σ_LR04) # trunkert slik at øvre grense = mean + 1σ length(intpD_t_1σ_LR04_trunc) # 53201 length(intpD_t_LR04) # 53201 typeof(intpD_t_1σ_LR04_trunc) # ...So I don't get why I can't reformulate as uivD below....# Truncated age model uncertainty redefined as an uivD # I wanted to use truncated age uncertainty (truncated at 1σ) for the first round of analysis, but cannot define truncated values as uivD t_uiv_LR04_truncated = [UncertainValue(Normal, intpD_t_LR04[i], intpD_t_1σ_LR04_trunc[i]) for i in 1:length(intpD_t_LR04)] #uivD_LR04_truncated = UncertainIndexValueDataset(t_uiv_LR04_truncated, d18O uiv LR04)

I wanted to use truncated age uncertainty (truncated at 1 σ) for the first round of analysis, but DOESN't WORK to define truncated values as uivD

We now have the object in the right format for analysis, an UncertainIndexValueDataset that carries on uncertainties in KDE.

In the meanwhile, we make a .jld2 file to contain the wrangeld data of the LR04 record:

```
In [88]:
```

```
# Save the relevant arrays of the LR04 record in a .jld2 file

uivD_LR04 = uivD_LR04_noageunc

# MERK: uivD_LR04 er no lagra utan aldersusikkerhet, for det gav heilt bananas k
onfidensintervall.

# Har sidan funne ut (Jo sjekka) at aldersusikkerheten ikkje var inkludert, så d
et blir rett å bruke fullageunc

@save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/LR04.jld2" uivD_LR04_noag
eunc uivD_LR04_fullageunc #uivD_LR04_truncated
```

```
In [89]:
```

```
@load "../Koding/WrangledDataFiles/uivDs/LR04.jld2"
```

```
Out[89]:
```

```
2-element Array{Symbol,1}:
   :uivD_LR04_noageunc
   :uivD_LR04_fullageunc
```

```
In [ ]:
```

```
#plot(uivD_LR04)
```

vi) Binned Resampling

The next step will be to perform BinnedResampling to get on a common time grid to run our analyses. What the common grid is will vary, depending on the length and resolution of the time series it is paired with for analysis. The definition of the common grid and ensuing BinnedResampling of the time series will therefore be done in the notebooks for analyses and results (NBRs), where every every NBR is defined by a different grid.

Define a grid for BinnedResampling method to allow one value for every 1000 years. It is important in analyses that all values are on the same time grid. For the initial analysis we will use a time step of 1000 years (following the example of the Spratt&Lisiecki sea level record). We may later experiment with higher and lower resolution (smaller and larger binsizes) in sensitivity analyses.

```
In [81]:
binsize = 1 # time step
Out[81]:
1
```

Intuitively, the grid would run between the first to the last point in the time series. However, since we want the datapoints (bin median) to fall on every discrete 1000 years, like in the Spratt Lisiecki grid, we must ensure that the bin midpoint falls on an integer. we therefore let the bin edges be midways between the intergers (every 0.5, 1.5, 2.5 etc kyrs). To avoid skewedness in the outermost bins or need for extrapolation, we sacrifice the time series length by a datapoint in each end (we start the grid half a binsize *over* the first point, rather than below, and end the grid hald a binsize *before* the last datapoint).

```
In [83]:
```

```
#= Define a grid for binned resampling,
with approx the length of LR04 record, and one bin for every 1000 years.
Formulate as to assign one value at every discrete 1 kyr (bin midpoint at Intege
rs)) =#

tmin_LR04 = ceil(minimum(t_LR04)) # -5320 kyrs BP
tmax_LR04 = floor(maximum(t_LR04)) # 0 kyrs BP (present)
binmidpoints_LR04 = tmin_LR04 : binsize : tmax_LR04
grid_LR04 = tmin_LR04 - binsize/2 : binsize : tmax_LR04 + binsize/2
```

```
Out[83]:
```

```
-5320.5:1.0:0.5
```

• Defining the BinnedResampling method:

In [59]:

```
#= Define the resampling method:
in each bin of the grid_LR04, draw 1000 resamples (with substitution) =#
resampling_method_LR04 = BinnedResampling(grid_LR04, 1000)
```

Out[59]:

BinnedResampling{StepRangeLen{Float64, Base.TwicePrecision{Float64}, B
ase.TwicePrecision{Float64}}}(-5320.5:1.0:0.5, 1000)

 Compute a binned resampling of the LR04 record on the defined 1 kyr grid, drawing a 1000 resamples in each bin.

In [68]:

```
77.944174 seconds (45.99 M allocations: 48.826 GiB, 39.04% gc time)
```

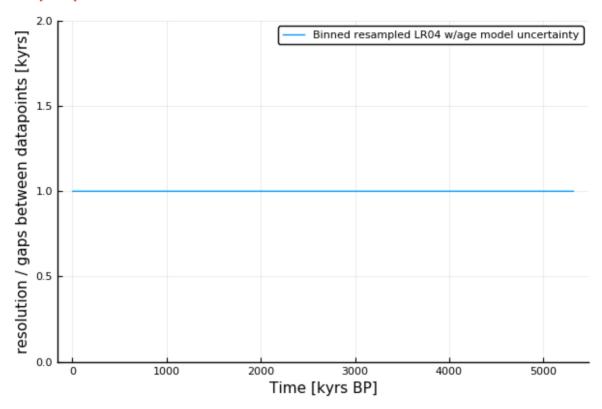
Out[68]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 5321 uncertain values coupled with 5321 uncertain ind
ices

· Check that we have a value in all bins

In [107]:

Out[107]:



```
In [93]:
```

```
LR04_binned_fullength_noageunc_intp = resample(uivD_LR04_noageunc, resampling_me
thod_LR04)
#LR04_binned_truncatedageunc = resample(uivD_LR04_truncated, resampling_method_L
R04)
```

```
UndefVarError: uivD_LR04_noageunc not defined
Stacktrace:
  [1] top-level scope at In[93]:1
In [69]:
@save "../Koding/WrangledDataFiles/Binned ts fullength/LR04.jld2" LR04 binned f
```

Plots

Plot the binned resampled uivD LR04 time series. We will plot the 95% confidence interval, for a solid communication of the uncertainty.

ullength fullageunc LR04 binned fullength noageunc intp

```
In [67]:
```

```
@load "../Koding/WrangledDataFiles/Binned_ts_fullength/LR04.jld2"
```

Out[67]:

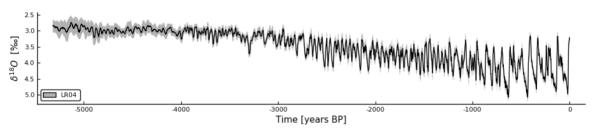
```
2-element Array{Symbol,1}:
  :LR04_binned_fullength_fullageunc
  :LR04_binned_fullength_noageunc_intp
```

1. with age model uncertainty:

In [78]:

```
### Plot the binned resampled uivD LR04 time series with the 95% confidence inte
rval
LR04 = LR04 binned fullength fullageunc
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(LR04.values, 0.5)
bin upper = quantile.(LR04.values, 0.975) .- bin median
bin lower = bin median .- quantile.(LR04.values, 0.025)
plot LR04 fullageunc =
plot(binmidpoints_LR04, bin_median,
   ribbon = (bin lower, bin upper),
   color = :black, fillalpha = 0.3,
   label = "LR04",
   xlabel = "Time [years BP]",
   ylabel = L"$\delta^{18}0$ $[\perthousand]$",
   grid = false, yflip = true,
   size = (1000,200), legend = :bottomleft
    )
```

Out[78]:

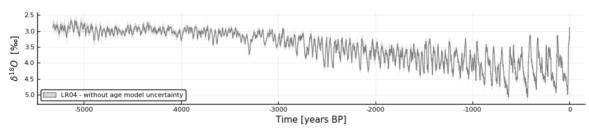


1. without taking into account the age model uncertainty:

In [80]:

```
### Plot the binned resampled uivD LR04 time series with the 95% confidence inte
rval
LR04 = LR04 binned fullength noageunc intp
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(LR04.values, 0.5)
bin upper = quantile.(LR04.values, 0.975) .- bin median
bin lower = bin median .- quantile.(LR04.values, 0.025)
plot LR04 noageunc =
plot(binmidpoints LR04, bin median,
    ribbon = (bin lower, bin upper),
    color = :grey, fillalpha = 0.3,
    label = "LR04 - without age model uncertainty",
    xlabel = "Time [years BP]",
    ylabel = L"$\delta^{18}0$ $[\perthousand]$",
    yflip = true, #grid = false,
    size = (1000, 200), legend = :bottomleft
    )
```

Out[80]:



2. Global sea level (GSL)

Due to different uncertainties/methodologies and limited overlap between the sea level time series, we operate with 4 different sea-level records to increase the robustness of our analysis:

- SL = Spratt & Lisiecki (2016) PCA of 5 different sea-level records spanning the last 800 ka
- E = Elderfield et al. (2012) Temperature deconvolution of $\delta^{18}O_b$ spanning 1.5 Ma, tuned to LR04
- G = Grant et al. (2014) Red Sea $\delta^{18}O_b$ stack, tuned to Sanbao speleothem by record of common monsoon-signal.
- R = Rohling et al. (2014) Mediterranean $\delta^{18}O_b$ stack, tuned to Sanbao speleothem by record of common monsoon-signal.

2.1 - Spratt-Lisiecki GSL stack

- GSL stack from Spratt & Lisiecki (2016), denoted SL
- Principal component analysis (PCA) of a stack of 5 GLS records, spanning the last 798 kyr.
- Note 1: the SL stack is tuned to the LR04, which is orbitally tuned. This may create bias in our results in analyses between this record and insolation time series.
- Note 2: no uncertainty communicated in Pangaea dataset. Have therefore used dataset from Jo.

Check with Jo where data comes from. (Article reads 1σ "Bootstrapping and random sampling yield mean uncertainty estimates of 9-12 m (1σ) ").

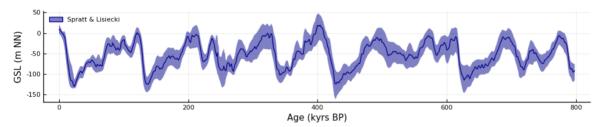
i) Read in dataset from Spratt & Lisiecki

In [51]:

```
# Read in dataset from Spratt & Lisiecki
filepath SL = "../../MASTER 2.0/data/sea-level/SprattLisiecki/SL v4(Jo).txt"
rawD SL = DelimitedFiles.readdlm(filepath SL,'\t', Float64, '\r', dims = (799, 7
), skipstart = 1)
# Naming the arrays
t SL = rawD SL[:,1]/1000 # age (kyrs), (converted from years to kyrs by div
iding with 1000.)
GSL mean SL = rawD SL[:,2] # SeaLev longPC1 # mean sea level (m)
GSL 1\sigma SL = rawD SL[:,3]
                             # sea level uncertainty 1\sigma
GSL SL err lo = rawD SL[:,4] # lower 95% confidence interval quantile from Mont
e Carlo analysis
GSL_SL_err_up_ = rawD_SL[:,5] # upper 95% CI quantile from Monte Carlo analysis
# SprattLisiecki is tuned to LR04 (meaning it has relicts of orbital tuning)
# Jo has tuned the stack to the Sanbao speleothem, to have the GSL stack on the
same age model as the Mediterranean record (Rohling)
## SHOULD WE USE Jo's AGE MODEL?
SL t speleotuning = rawD SL[:,6]
SL t speleotuning 1\sigma = \text{rawD SL}[:,7]./1000; \# \text{ from years to kyrs}
```

In [47]:

Out[47]:



SprattLisiecki is tuned to LR04 (meaning it has relicts of orbital tuning). Jo has tuned the stack to the Sanbao speleothem, to have the GSL stack on the same age model as the Mediterranean record (Rohling). SHOULD WE USE Jo's AGE MODEL?

ii) Reverse dataset, to redefine from age to time

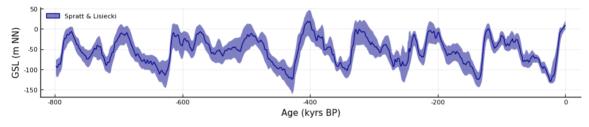
In [59]:

```
# For the analysis, we want to change age (increasing backwards) to time (increa
sing forward).
# We therefore reverse the dataset along the 1st dimension (rows)
revD SL = reverse(rawD SL, dims = 1) # reversing the dataset along the 1st dimen
sion (rows)
# Naming the arrays
t SL = - revD SL[:,1]/1000 # time (kyrs) # (converted from years to kyrs by
dividing with 1000. Negative since we define present as 0)
GSL mean SL = revD SL[:,2] # SeaLev longPC1 # indexation has been reversed. St
ays positive values
GSL 1\sigma SL = revD SL[:,3]
                                              # indexation has been reversed. St
ays positive values
                    # USE THIS?
GSL SL err lo = revD_SL[:,4] #
                                              # from MonteCarlo analysis (lower
 95% CI quantile)
                  # OR THIS?
GSL SL err up = revD SL[:,5] #
                                              # from MonteCarlo analysis (upper
 95% CI quantile)
# SprattLisiecki is tuned to LR04 (meaning it has relicts of orbital tuning)
# Jo has tuned the stack to the Sanbao speleothem, to have the GSL stack on the
same age model as the Mediterranean record (Rohling)
## SHOULD WE USE Jo's AGE MODEL?
SL t speleotuning = -revD SL[:,6]./1000 #from years to kyrs
SL t speleotuning 1\sigma = \text{revD SL}[:,7]./1000; \# \text{ from years to kyrs}
```

In [81]:

@save "../Koding/WrangledDataFiles/BasicArrays/SprattLisiecki.jld2" t_SL GSL_mea n_SL GSL_1 σ _SL GSL_SL_err_lo GSL_SL_err_up SL_t_speleotuning SL_t_speleotuning_1 σ

In [74]:



In [73]:

```
mean(GSL_1\sigma_SL) # (Article reads 1\sigma "Bootstrapping and random sampling yield mean uncertainty estimates of 9-12 m (1\sigma)").
# Checks out, kind of?
```

Out[73]:

12.070150187734665

bla # DISCUSSION on how to define the 95% confidence interval (2 σ) # Plot plot(# title = "SprattLisiecki GSL stack", xlabel = "Time (kyrs BP)", ylabel = "GSL (m NN)", size = (1000, 200)) plot!(t_SL, GSL_mean_SL, markersize = 0.1, ribbon = (GSL_mean_SL - GSL_SL_err_lo, GSL_SL_err_up - GSL_mean_SL), # Looks more or less like the 2 σ ribbon color = "blue", label = "Spratt&Lisiecki err_lo, err_up") plot!(t_SL, GSL_mean_SL, markersize = 0.1, ribbon = (GSL_1 σ _SL, GSL_1 σ _SL), label = "Spratt&Lisiecki 1 σ ") plot!(t_SL, GSL_mean_SL, markersize = 0.1, ribbon = (2 * GSL_1 σ _SL), label = "Spratt&Lisiecki 2 σ ") # Note: looks like the quantiles given in dataset are for 95% CI. (matches 2 σ) # How to most accurately describe 2 σ ? # GSL_2 σ _SL_v1 = (GSL_SL_err_up - GSL_SL_err_lo) / 2 # From MonteCarlo analysis # GSL_2 σ _SL_v2 = 2 * GSL_1 σ _SL # # are these the same ? # they look the same from the plot, but they are not exactly the same # GSL_2 σ _SL_v1 - GSL_2 σ _SL_v2 # (If they were, this should give an array of zeros) # SO WHICH ONE CHOULD WE CARRY ON IN uivD? ### NONE OF THESE, uivD is defined with 1 σ . 95% CI is then calculated after.

Notes on age model and uncertainties

The Spratt & Lisiecki record is tuned to the LR04 age model. As far as we can read, the LR04 age model uncertainties have already been incorporated in the dataset by Spratt & Lisiecki. We do not have a separate array for age uncertainties, as these are incorporated in the value-error through Monte Carlo analysis (equivalent to what we do with binned resampling).

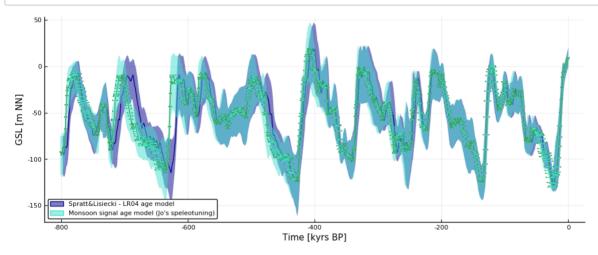
SprattLisiecki is tuned to LR04 (meaning it has relicts of orbital tuning)

Jo has tuned the stack to the Sanbao speleothem, to have the GSL stack on the same age model as the Mediterranean and Red Sea RSL records (Rohling and Grant)

SHOULD WE USE Jo's AGE MODEL?

In [97]:

```
# Plot of SprattLisiecki GSL
# Comparison of LR04 age model and Jo's tuning to monsoon signal in the Sanbao s
peleothem
plot SprattLisiecki agemodelcomparison =
plot(# title = "SprattLisiecki GSL stack",
    xlabel = "Time [kyrs BP]",
    ylabel = "GSL [m NN]", #NN given in unit in dataset -what does it mean?
    size = (1000, 400),
    legend = :bottomleft,
    #bg legend = :transparent
# Original LR04 age model
plot!(t SL, GSL mean SL,
   markersize = 0.1,
    ribbon = (GSL mean SL .- GSL SL err lo, GSL SL err up .- GSL mean SL), # use
this or 2 * 1sigma?
    color = :darkblue,
    label = "Spratt&Lisiecki - LR04 age model"
# Speleotuning age model
plot!(SL t speleotuning, GSL mean SL, ribbon = (GSL mean SL .- GSL SL err lo, GS
L SL err up .- GSL mean SL), color = :turquoise, fillalpha = 0.5, label = "Monso
on signal age model (Jo's speleotuning)")
plot!(SL t speleotuning, GSL mean SL, xerr = 2*SL t speleotuning 1\sigma, ms = 1, msc
olor = :turquoise, label = "")#"speleotuning age uncertainty (±2\u03c3)")
savefig("../../MASTER 2.0/figurar/RawData/GSL/SL agemodelcomparison.pdf")
```

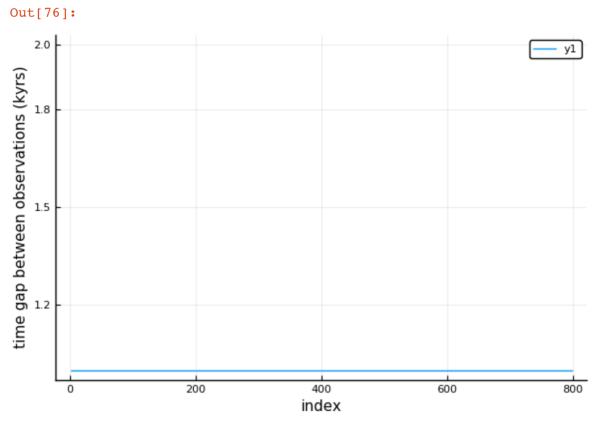


Interpolation?

In [76]:

```
# need for interpolation?
minimum(diff(t_SL))
mean(diff(t_SL))
maximum(diff(t_SL))

plot(diff(t_SL),
xlabel = "index",
ylabel = "time gap between observations (kyrs)")
```



No need to interpolate the SprattLisiecki record, as it is already on a regular grid with one value every 1000 year-grid

v) Redefining rD_SprattLisiecki as an UncertainIndexValueDataset

```
# Redefining ``rD_SprattLisiecki`` as an ``UncertainIndexValueDataset``

t_uiv_SL_noageunc = [UncertainValue(Normal, t_SL[i], 0) for i in 1:length(t_SL)]
# no age uncertainty

GSL_uiv_SL = [UncertainValue(Normal, GSL_mean_SL[i], GSL_1\subseteq_SL[i]) for i in 1:le

ngth(GSL_mean_SL)]

uivD_SL_noageunc = UncertainIndexValueDataset(t_uiv_SL, GSL_uiv_SL)

# with age uncertainty from LR04

t_uiv_SL_ageunc = [UncertainValue(Normal, t_SL[i], t_1\subseteq_SL[i]) for i in 1:length

(t_SL)] # SYSTEMATIC AGE UNCERTAINTY IN LR04 -INCLUDE HERE?

uivD_SL_ageunc = UncertainIndexValueDataset(t_uiv_SL_ageunc, GSL_uiv_SL)

# plot(uivD_SL)
```

In []:

```
# Save the relevant arrays of the LR04 record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/SprattLisiecki.jld2" t_SL
uivD_SL_ageunc uivD_SL_noageunc
```

In []:

```
plot(uivD_SL)
```

vii) Binned resampling on a 1 kyr timestep grid

In []:

```
# Define grid for binned resampling
grid_SL = ceil(minimum(t_SL)) + binsize/2 : binsize : floor(maximum(t_SL)) - bin
size/2

# Define the resampling method to draw 1000 resamples within each bin of the gri
d, and return a median with associated confidence
resampling_method_SL = BinnedResampling(grid_SL, 1000) # within each bin of the
grid_SL, draw a 1000 samples (with substitution)
```

Resample the uivD using the resampling method, to return one **mean/median** value with associated confidence for each bin in the grid

```
In [ ]:
```

```
@time SL_binned_fullength_ageunc = resample(uivD_SL_ageunc, resampling_method_SL
)
```

In []:

```
In [ ]:
```

```
@save "../../MASTER_2.0/Koding/WrangledDataFiles/Binned_ts_fullength/SprattLisiecki.jld2" SL_binned_fullength_ageunc SL_binned_fullength_noageunc
```

```
@load "../../MASTER_2.0/Koding/WrangledDataFiles/Binned_ts_fullength/SprattLisie
cki.jld2"
```

In []:

```
plot(SL_binned_fullength_noageunc)
```

Above is a plot of the record with fixed time (uncertainty in the time dimension is transposed to uncertainty in the value dimension through binned resampling).

Next, we compute the 0.5 and 0.975 quantiles, and make a plot with the 95% confidence interval ribbon.

In []:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
# Without the LR04 age model uncertainty
SL = SL binned fullength noageunc
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(SL.values, 0.5)
bin upper = quantile.(SL.values, 0.975) .- bin median
bin lower = bin median .- quantile.(SL.values, 0.025)
binmidpoints SL = [SL.indices[i].value for i in 1:length(SL)]
plot SL binned noageunc =
plot(binmidpoints SL, bin median,
    ribbon = (bin lower, bin upper),
    color = :darkblue,
    label = "Spratt & Lisiecki",
    xlabel = "Time [years BP]",
    ylabel = "GSL [m]",
    grid = false
```

```
### Plot the binned resampled uivD time series with the 95% confidence interval
SL = SL binned fullength ageunc
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(SL.values, 0.5)
bin upper = quantile.(SL.values, 0.975) .- bin median
bin lower = bin median .- quantile.(SL.values, 0.025)
binmidpoints SL = [SL.indices[i].value for i in 1:length(SL)]
plot SL binned ageunc =
plot(binmidpoints SL, bin median,
    ribbon = (bin lower, bin upper),
    color = :darkblue,
    label = "Spratt & Lisiecki",
    xlabel = "Time [years BP]",
    ylabel = "GSL [m]",
    grid = false
    )
```

2.2 - Elderfield GSL-record

- GSL record from Elderfield et al. (2012), denoted E for Elderfield
- Mg/Ca temperature deconvolution of $\delta^{18}O_b$, spanning the last 1.5 Myr.
- Data available from Pangaea, DOI: https://doi.org/10.1594/PANGAEA.786204
 (https://doi.org/10.1594/PANGAEA.786204)

In [109]:

```
# read in data
filepath Elderfield = "../../MASTER 2.0/data/sea-level/d180 ESL Elderfield.tab"
rawD E = readdlm(filepath Elderfield, skipstart = 16)
revD E = reverse(rawD E, dims = 1)
# name columns
t E = -revD E[:,1] # Age [ka BP]
                                                # note: AGE [ka BP] is interpol
ated age. From LR04? CHECK
                    # \delta180 H20 [% SMOW]
#revD[:,2]
                                                # What is this?
GSL_E = -revD_E[:,3] # Sea lev rel [m]
                                                # NOTE: MADE NEGATIVE TO MATCH
WITH OTHERS.
# uncertainty?? IMPORT AGE UNCERTAINTY FROM LR04?
                      # need to define anew. > <
# t \sigma LR04
```

Uncertainties

For the Elderfield record, no uncertainty is included in the public dataset. However, a note is made in the Elderfield article (2012), as well as an uncertainty analysis for this record is done in supplementary materials in Rohling et al. (2014). We will discuss these in the following.

Elderfield et al. (2012) report an "error in $\delta^{18}O_W$ of \pm 0.2 \textperthousand (per mille), from propagation of estimated temperature and $\delta^{18}O_C$ uncertainties.

- We interpret this as 1σ , and
- since $\delta^{18}O_W$ is the ice volume component of the signal, we define this same error on the GSL estimate.

In [106]:

```
# Create an array for 10 based on the uncertainty communicated in Elderfield et al. (2012)

GSL_1o_E__ = 0.0002 .* GSL_E # +- 0.2 per mille

; # THIS LOOKS WAY TOO SMALL in plot, but is all I can gather from the articl e. Something I'm not understanding?
```

Rohling et al. (2014) make a note of that uncertainties are poorly constrained in the Elderfield et al. (2012). They have therefore performed a probabilistic assessment of the Elderfield record. Rohling et al. (2014) find that the total uncertainty is of about ± 35 m (1σ):

- The bulk of the uncertainty (\pm 35 m) is random calibration uncertainty on temperature sensitivity (T_S), due to many unknowns.
- In addition there is the $\delta^{18}O_W$ to sea-level conversion uncertainty (\pm 10% (0.1 meters uncertainty per meter GSL change is the standard ratio).

Rohling et al. (2012) further state that this uncertainty "may appear large, but there is strong autocorrelation in the record, which leads to considerably tighter uncertainty limits to underlying 'mean' trends."

Sinces it is not the aboslute sea-level we are interested in, but rather the *dynamics* (relative changes) in sea level, we therefore don't include the random calibration uncertainty, but only the sea-level conversion uncertainty.

In [110]:

```
# Create an array for 1\sigma based on the uncertainty communicated in Rohling et al (2014)

GSL_1\sigma_E = zeros(length(GSL_E)) # Create an empty array of the length of the rec ord, to contain the 1\sigma

unc_calibrationTs = GSL_1\sigma_E[:] .= 35 # Uncertainty from temperature sensitivity calibration: adds up to \pm 35 meters GSL # BUT WE DON'T WANT TO USE THIS unc_conversion = GSL_E .* 0.1 # Uncertainty from conversion of d180_W to sea-level: 1\sigma = 10%

GSL_1\sigma_E_hiunc = unc_calibrationTs .+ unc_conversion # hiunc = high uncertainty - including uncertainties on temperature sensitivity

GSL_1\sigma_E_lounc = unc_conversion # lounc = low uncertainty - excluding uncertainties on temperature sensitivity;
```

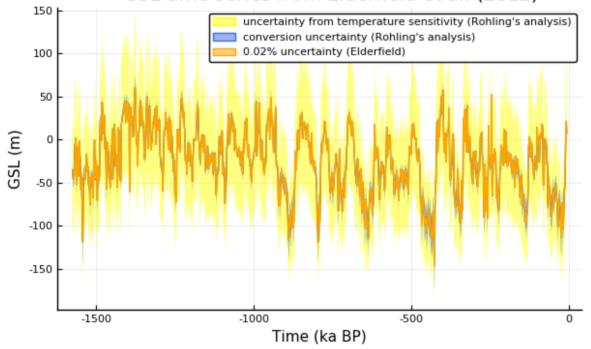
Note: weird that Elderfield communicates a 2%% (per mille) conversion uncertainty, and Rohling 10% conversion uncertainty. Comments?

In [109]:

```
# plot for visual comparison of the communicated uncertainty ranges
plot(title = "Comparison of communicated uncertainty ranges for the
    GSL time series from Elderfield et al. (2012)",
    xlabel = "Time (ka BP)",
    ylabel = "GSL (m)")
plot!(t_E, GSL_E,
    ribbon = (2*GSL 1\sigma E hiunc), # 95% confidence interval. Using the general 1\sigma
= +-35 meters, this adds up to +-140 meters...
    color = :yellow,
    label = "uncertainty from temperature sensitivity (Rohling's analysis)")
plot!(t E, GSL E,
    ribbon = (2*GSL 10 E lounc), # 95% CI. Using only conversion uncertainty (1
\sigma is 10%)
    color = :royalblue,
    label = "conversion uncertainty (Rohling's analysis)")
plot!(t E, GSL E,
    ribbon = (2*GSL 1\sigma E), # 95% CI. 1\sigma is 0.02% (0.02 permille)
    color = :orange,
    label = "0.02% uncertainty (Elderfield)")
```

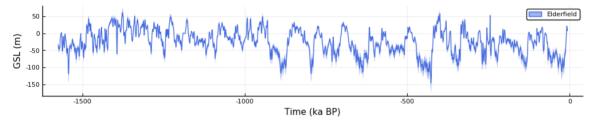
Out[109]:

Comparison of communicated uncertainty ranges for the GSL time series from Elderfield et al. (2012)



Note: We only take the *conversion uncertainty* (lounc, $\pm 10\%$) into account in running our analysis. We take care not to translate the uncertainty in determining the temperature sensitivity into uncertainty in the dynamics of sea level change (this would be an overcommunication of uncertainty). The reason for this is that we are not really interested in the absolute sea level change, but rather the *dynamics* of sea-level change.

In [111]:



Note on age model uncertainties

The Elderfield GSL record is correlated to the LR04 stack. Age model uncertainties on the LR04 stack are read in above. However, we make a note of the fact that the Elderfield record is a continuous stratigraphic core, implying next to no uncertainty in relative chronology. Therefore, propagations of age uncertainties from the LR04 age model should be done strictly monotonic (no age reversals allowed).

What will be the right way to include age uncertainties? Include in uivD or not? How does the BinnedResampling method treat age uncertainties?

Veiledningsmøte 9/6/2020: **Include age uncertainties in uivD** for now. We can later run sensitivity analyses to check if we get very different results with/without age uncertainties.

An array for the age model uncertainties will be created below, with the interpolation

In [121]:

```
# Make an array of the potential systematic deviations in the LR04 age model
t 4o LR04 = zeros(length(t LR04)) # maximum age model envelope (interpreted as t
he 95% confidence envelope, i.e. \pm 2\sigma)
t 4\sigma LR04[ t LR04 .> -1000] .= 4
                                                              # 4 ky from -1 Ma to 0
(present)
t 4\sigma LR04[(t LR04 .<= -1000) .& (t LR04 .> -3000)] .= 6 # 6 ky from -3 to -1 Ma
t 4\sigma LR04[(t LR04 .<= -3000) .& (t LR04 .> -4000)] .= 15 # 6 ky from -4 to -3 Ma
t 4\sigma \ LR04 \ (t \ LR04 \ .<= -4000) \ . \& \ (t \ LR04 \ .> -5000) \ ] \ .= \ 30 \ \# \ 6 \ ky \ from \ -3 \ to \ -1 \ Ma
t 4\sigma LR04[ t LR04 .<= -5000] .= 40
                                                              # 40 kv for before -5 M
а
t 1\sigma LR04 = t 4\sigma LR04 ./ 4 # 1\sigma
# interpolation
# create a continuous function with linear interpolation between every datapoint
# This allows us to choose the interpolated value for any point in time.
interpolate t 1σ LR04 = LinearInterpolation(t 1σ LR04, t LR04)
                                                                       # age model u
ncertainty (full)
# make a fine grained grid to contain the interpolated values
#fine grid LR04 = ceil(minimum(t LR04)) : 0.1 : floor(maximum(t LR04)) # One bin
for every 100 years (0.1 kyrs)
# make new arrays for interpolated data
# give a value from the interpolate function to every bin in the fine grid
#intpD t 1\sigma LR04 = [interpolate t 1\sigma LR04(i) for i in fine grid LR04]
# age model uncertainty from LR04
t 1\sigma E = [interpolate t 1\sigma LR04(i) for i in t E]
########## OR, define age model uncertainty directly
t 4\sigma E = zeros(length(t E)) # maximum age model envelope (interpreted as the 95%
confidence envelope, i.e, \pm 2\sigma)
t 4\sigma E[t E .> -1000] .= 4
                                                   # 4 ky from -1 Ma to 0 (present)
t 4\sigma E[(t E .<= -1000) .& (t E .> -3000)] .= 6 # 6 ky from -3 to -1 Ma
t_1\sigma_E = t_4\sigma_E ./ 4 # 1\sigma
```

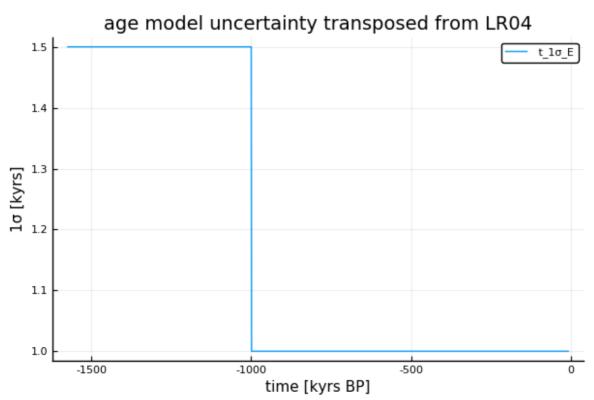
Out[121]:

```
1557-element Array{Float64,1}:
 1.5
 1.5
 1.5
 1.5
 1.5
 1.5
 1.5
 1.5
 1.5
 1.5
 1.5
 1.5
 1.5
 1.0
 1.0
 1.0
 1.0
 1.0
 1.0
 1.0
 1.0
 1.0
 1.0
 1.0
 1.0
```

In [122]:

```
plot(t_E, t_1\sigma_E, xlabel = "time [kyrs BP]", ylabel = "1\sigma [kyrs]", label = "t_1\sigma _E", title = "age model uncertainty transposed from LR04")
```

Out[122]:



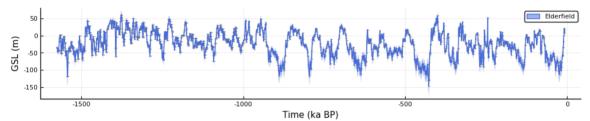
MA_NB1_PrepData

We see that we get $1\sigma = 1$ kyr for the youngest 1 Ma part of the record, and $1\sigma = 1.5$ kyr for the remainding part of the record. That is right according to the 95% confidence interval (4 σ) being 4 kyrs for 0-1 Ma and 6 kyrs for 1-3 Ma. We are thus sure to have created the age model uncertainty array correctly.

In [115]:

```
@save "../Koding/WrangledDataFiles/BasicArrays/Elderfield.jld2" t_E t_1\sigma_E GSL_E GSL_1\sigma_E_lounc
```

In [114]:



iv) Interpolation

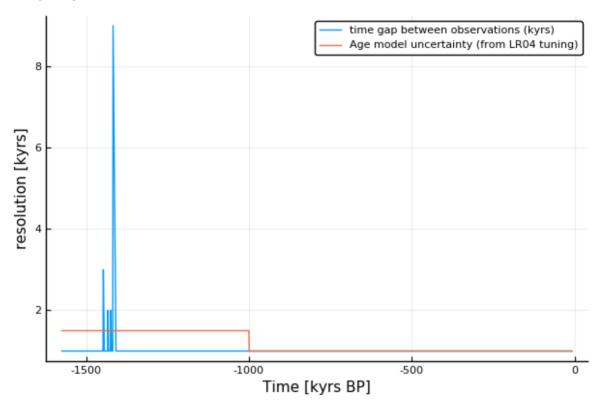
First, check if there is need for interpolation. (We want at least a millenial resolution for our analysis)

In [123]:

```
# need for interpolation?
minimum(diff(t_E)) # 1000 years
mean(diff(t_E)) # 1007 years
maximum(diff(t_E)) # 9000 years
# Yes, we need to interpolate

plot(t_E, diff(t_E),
    xlabel = "Time [kyrs BP]",
    ylabel = "resolution [kyrs]",
    label = "time gap between observations (kyrs)",
)
plot!(t_E, t_1\sigma_E, label = "Age model uncertainty (from LR04 tuning)")
```

Out[123]:



Mean time gap between observations is around 1000 years. The largest time gap between observations is 9000 years. To meet the requirement of the common grid we will later define we therefore need to interpolate, to fill in the missing values.

In the case of this record, the highest resolution is one value every 1000 years. I will therefore not do higher resolution analyses with this record.

LOGICAL QUESTION: making a higher resolution interpolation grid will be misleading when it comes to the resampling, won't it? Since it will estimate uncertainties based on fake data. Won't it be better here to use the interpolation grid as 1 kyr resolution?

In []:

```
# Interpolation of Elderfield record
interpolate t E = LinearInterpolation(t E, t E)
                                                      # interpolation function
for time
interpolate E = LinearInterpolation(GSL E, t E) # interpolation function
for GSL
                                                    # fine grained grid for whi
fine grid E = minimum(t E) : 1 : maximum(t E)
ch we want to interpolate values - one bin for every 100 years
                                                     # MAYBE WE SHOULD JUST HAVE
1 kyr interpolation? HIGHER WOULD ONLY BE FAKE DATA...
intpD t E = [interpolate t E(i) for i in fine grid E] # Array of the interpolat
ed time values
intpD GSL E = [interpolate E(i) for i in fine grid E]; # Array of the interpolat
ed GSL values
# carrying on equivalent array of uncertainties
intpD 1\sigma E = abs.(intpD GSL E) .* 0.1 # conversion uncertainty +/- 10% (lounc) #
WHAT's THE abs FOR??
# age model uncertainty from LR04
intpD t 1\sigma E = [interpolate t 1\sigma LR04(i) for i in fine grid E]
;# like discussed above, this looks right
```

In []:

```
intpD_t_E
```

```
In [ ]:
```

v) Redefine the interpolated array as an UncertainIndexValueDataset

In []:

```
# Redefining Elderfield interpolated data as uivD

# Note, here lounc is used (only d180 conversion uncertainty, not temperature se nsitivity),

GSL_uiv_E = [UncertainValue(Normal, intpD_GSL_E[i], intpD_1o_E[i]) for i in 1:le ngth(intpD_GSL_E)] # conversion uncertainty 10%, as reported in analysis of Roh ling et al. (2014)

# 2 time arrays are defined:
# - one with no uncertainty in time dimension
t_uiv_E_noageunc = [UncertainValue(Normal, intpD_t_E[i], 0) for i in 1:length(in tpD_t_E)] # LR04 age model uncertainty not included

# - and one with age model uncertainty from LR04
t_uiv_E_ageunc = [UncertainValue(Normal, intpD_t_E[i], intpD_t_1o_E[i]) for i in 1:length(intpD_t_E)] # time uncertainty from LR04 age model

uivD_E_ageunc = UncertainIndexValueDataset(t_uiv_E_ageunc, GSL_uiv_E)
uivD_E_noageunc = UncertainIndexValueDataset(t_uiv_E_noageunc, GSL_uiv_E)
```

```
In [ ]:
```

```
@time plot(uivD_E)
```

Save the relevant arrays of the wrangled Elderfield data in a .ild2 file

```
In [ ]:
```

```
# Save the relevant arrays of the Elderfield record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Elderfield.jld2" intpD_t_
E uivD_E_ageunc uivD_E_noageunc
```

vi) Binned resampling on a grid timesteps of 1000 years

```
In [ ]:
```

```
#Define the resampling method for the Elderfield record

# define grid

tmin_E = ceil(minimum(intpD_t_E)) + binsize/2

tmax_E = floor(maximum(intpD_t_E)) - binsize/2

grid_E = tmin_E : binsize : tmax_E

# In each bin of the grid, resample 1000 values

resampling_method_E = BinnedResampling(grid_E, 1000)
```

In []:

```
# resample the uncertain index value dataset with the resampling method defined
above
@time E_binned_fullength_ageunc = resample(uivD_E_ageunc, resampling_method_E)
```

In []:

```
# resample the uncertain index value dataset with the resampling method defined
  above
@time E_binned_fullength_noageunc = resample(uivD_E_noageunc, resampling_method_
E)
```

In []:

```
# Save the relevant arrays of the Elderfield record in a .jld2 file @save "../../MASTER_2.0/Koding/WrangledDataFiles/Binned_ts_fullength/Elderfield.jld2" E_binned_fullength_ageunc E_binned_fullength_noageunc
```

In []:

```
@load "../Koding/WrangledDataFiles/Binned_ts_fullength/Elderfield.jld2"
```

Plot of 95% confidence interval, including the LR04 age model uncertainty:

In []:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
E = E binned fullength ageunc
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(E.values, 0.5)
bin upper = quantile.(E.values, 0.975) .- bin median
bin lower = bin median .- quantile.(E.values, 0.025)
# time array
binmidpoints E = [E.indices[i].value for i in 1:length(E)]
plot E binned ageunc =
plot(binmidpoints E, bin median,
    ribbon = (bin lower, bin upper),
    color = :royalblue,
    label = "Elderfield",
    xlabel = "Time [years BP]",
    ylabel = "GSL [m]",
    grid = false
    )
```

Without the LR04 age model uncertainty:

In []:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
E = E binned fullength noageunc
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(E.values, 0.5)
bin upper = quantile.(E.values, 0.975) .- bin median
bin lower = bin median .- quantile.(E.values, 0.025)
# time array
binmidpoints E = [E.indices[i].value for i in 1:length(E)]
plot E binned noageunc =
plot(binmidpoints E, bin median,
    ribbon = (bin lower, bin upper),
    color = :royalblue,
    label = "Elderfield",
    xlabel = "Time [years BP]",
    ylabel = "GSL [m]",
    grid = false
    )
```

The difference seems small, even though we initially estimated the LR04 age model uncertainty to be overcommunicated.

We will use the version with age model uncertainty in our first round of analysis, for a conservative approach.

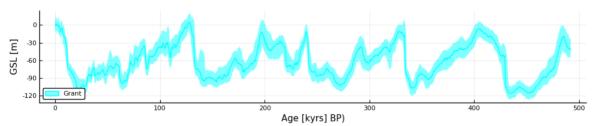
2.3 Grant RSL/GSL-record

- Grant et al. (2014), denoted G
- High resolution record of relative sea level at the straight of Bab el Mandab, Red Sea (equivalent to GSL). Spans the last 500 kyr.
- Data available from Nature, DOI: https://doi.org/10.1038/ncomms6076

In [123]:

```
# read in data
filepath Grant = "../../MASTER 2.0/data/sea-level/Grant(2014)/SD1 v1.txt"
rawD G = readdlm(filepath Grant, dims = (3951-14, 4), '\t', '\r', skipstart = 14
# naming columns
t_G_ = rawD_G[:,1]
                          # time kyr
RSL G = rawD G[:, 4]
                     # RSL_Pmax_m
RSL_q95lo_G = rawD_G[:,2] \# RSL_95low_m
RSL q95up G = rawD G[:,3] \# RSL 95high m
# define an array to carry on uncertainty - standard deviation (1\sigma):
RSL_1\sigma_G = (RSL_q95up_G .- RSL_q95lo_G) / 4 # 95%CI-quantiles include 2\sigma on e
ach side of mean, so we divide uncertainty range by 4 to get 1\sigma (standard deviat
ion)
;
# plot
plot_Grant_raw_age =
plot(#title = "Grant RSL record",
    xlabel = "Age [kyrs] BP)",
    ylabel = "GSL [m]",
    size = (1000, 200),
    legend = :bottomleft,
    #bg legend = :transparent
)
plot!(t G , RSL G ,
    ribbon = (RSL_G_ .- RSL_q95lo_G_, RSL_q95up_G_ .- RSL_G_), #plotting the 95%
confidence interval
    color = :cyan,
    label = "Grant")
```

Out[123]:



Reverse dataset to redefine from age to time

In [132]:

```
revD_G = reverse(rawD_G, dims = 1)
typeof(revD_G)  # Array{Float64,2}

# naming columns

t_G = -revD_G[:,1]  # time_kyr

RSL_G = revD_G[:,4]  # RSL_Pmax_m

RSL_q95lo_G = revD_G[:,2]  # RSL_95low_m

RSL_q95up_G = revD_G[:,3]  # RSL_95high_m

# define an array to carry on uncertainty - standard deviation (1\sigma):

RSL_1\sigma_G = (RSL_q95up_G .- RSL_q95lo_G) ./ 4  # 95%CI-quantiles include 2\sigma on each side of mean, so we divide uncertainty range by 4 to get 1\sigma (standard deviation)
;
```

In [145]:

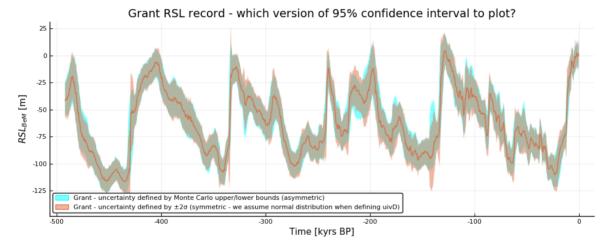
```
@save "../Koding/WrangledDataFiles/BasicArrays/Grant.jld2" t_G RSL_G RSL_1\sigma_G RSL_q951o_G RSL_q95up_G
```

Age uncertainty in the Grant record is incorporated in the value dimension through Monte Carlo analysis (like we do when binned resampling)

In [143]:

```
# plot
plot_Grant_raw =
plot(title = "Grant RSL record - which version of 95% confidence interval to plo
t?",
    xlabel = "Time [kyrs BP]",
    ylabel = string(L"RSL {BeM}", " [m]"),
    size = (1000, 400),
    legend = :bottomleft,
plot!(t G, RSL G,
    ribbon = (RSL G - RSL q951o G, RSL q95up G - RSL G), #plotting the 95% confi
dence interval
    color = :cyan,
    label = "Grant - uncertainty defined by Monte Carlo upper/lower bounds (asym
metric)")
plot!(t G, RSL G, ribbon = (2*RSL 1\sigma G, 2*RSL 1\sigma G), label = "Grant - uncertaint
y defined by \pm 2\sigma (symmetric - we assume normal distribution when defining uivD)"
)
```

Out[143]:



Which version of the 95% confidence interval should we plot?

- Monte carlo analysis (reported in dataset give slightly asymmetric uncertainty,
- while we assume normal distribution (σ)

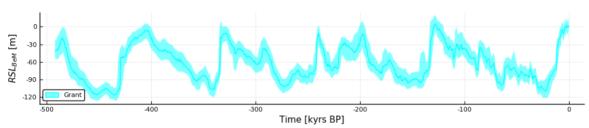
In [140]:

```
#@load "../Koding/WrangledDataFiles/BasicArrays/Grant.jld2" t_G RSL_G RSL_10_G

# plot
plot_Grant_raw =
plot(xlabel = "Time [kyrs BP]",
    ylabel = string(L"RSL_{BeM}", " [m]"),
    size = (1000,200),
    legend = :bottomleft,
    )
plot!(t_G, RSL_G,
    ribbon = (RSL_G - RSL_q95lo_G, RSL_q95up_G - RSL_G), #plotting the 95% confidence interval
    color = :cyan,
    label = "Grant")

#savefig("../../MASTER_2.0/figurar/RawData/GSL/plot_Grant.pdf")
```

Out[140]:



Interpolation. First, check if there is need for interpolation on the Grant data. (We want at least a millenial resolution for our analysis)

In []:

```
# need for interpolation?
minimum(diff(t_G)) # 125 years
mean(diff(t_G)) # 125 years
maximum(diff(t_G)) # 125 years

plot(diff(t_G),
xlabel = "index",
ylabel = "time gap between observations (kyrs)",
label = "diff(t_G)")
```

Grant is already on a regular time grid, with one observation every 125 years. We therefore have no need for interpolation on this dataset. We also note that we may use this time series to run higher resolution analyses.

Redefine as an UncertainIndexValueDataset

```
t_uiv_G = [UncertainValue(Normal, t_G[i], 0) for i in 1:length(t_G)] # no age un
certainty - incorporated in RSL uncertainty through monte carlo analysis
RSL_uiv_G = [UncertainValue(Normal, RSL_G[i], RSL_1\sigma_G[i]) for i in 1:length(RSL
_G)]
uivD_G = UncertainIndexValueDataset(t_uiv_G, RSL_uiv_G)
```

```
In [ ]:
```

```
# Save the relevant arrays of the Elderfield record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/WrangledData_Grant.jld2" t_G ui
vD_G
```

vi) Binned resampling

First, we bin the record on a grid with timesteps of 1000 years

```
In [15]:
@load "../../MASTER_2.0/Koding/WrangledDataFiles/WrangledData_Grant.jld2"
Out[15]:
2-element Array{Symbol,1}:
    :t_G
    :uivD_G

In []:

tmin_G = ceil(minimum(t_G))
    tmax_G = floor(maximum(t_G))
    grid_G = tmin_G + binsize/2 : binsize : tmax_G - binsize/2
    resampling_method_G = BinnedResampling(grid_G, 1000)

In []:
G binned fullength = resample(uivD G, resampling method G)
```

Plots

Plot of the uivD, here we see time uncertainty has been transposed into value uncertainty through the binned resampling

```
In []:

# plot(uivD_G)
a = rand(10)
plot(a, grid = (1,0))
```

I want some help to plot with a 1 kyr grid, to illustrate what happens when binned resampling

```
In [ ]:
```

```
#plot(G_binned_fullength, grid = ())
```

Now we plot the binned resampled uivD time series with the 95% confidence interval

In []:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
G = G binned fullength
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(G.values, 0.5)
bin_upper = quantile.(G.values, 0.975) .- bin_median
bin lower = bin median .- quantile.(G.values, 0.025)
# time array
binmidpoints_G = [G.indices[i].value for i in 1:length(G)]
plot G binned =
plot(binmidpoints G, bin median,
    ribbon = (bin lower, bin upper),
    color = :cyan,
    label = "Grant",
    xlabel = "Time [years BP]",
    ylabel = string(L"RSL {BeM}", " [m]"),
    grid = false
    )
```

Binned resampling of the record on a finer grid, to use in high resolution analyses. We will make two hr versions; one with timestep of 125 years, for analyses with La2004, and one with timesteps of 500 years, for high resolution analysis with the Martinez-García record.

• G binned fullength hr125, for hr analyses with La2004

```
# Binned resampling on a finer grid, for high resolution analyses

binsize_hr = 0.125
grid_G_hr = minimum(t_G) + binsize_hr/2 : binsize_hr : maximum(t_G) - binsize_hr
/2

# Intuitively, we would have the grid start at half a binsize *before* the first
datapoint (binmidpoint).

# But that might skew the first point in the record a bit (unless we extrapolate
for the first half of the bin, but that also imposes some assumptions).

# We therefore opt to cut the length of the timeseries by a datapoint in each en
d, to avoid skewedness or extrapolation.

resampling_method_G_hr = BinnedResampling(grid_G_hr, 1000)

G_binned_fullength_hr125 = resample(uivD_G, resampling_method_G_hr)
```

In []:

```
### Plot the high resolution binned resampled uivD with it's 95% confidence inte
rva1
G = G binned fullength hr
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(G.values, 0.5)
bin upper = quantile.(G.values, 0.975) .- bin median
bin lower = bin median .- quantile.(G.values, 0.025)
# time array
binmidpoints G = [G.indices[i].value for i in 1:length(G)]
plot G binned =
plot(binmidpoints_G, bin median,
    ribbon = (bin lower, bin upper),
    color = :cyan,
    label = "Grant",
    xlabel = "Time [years BP]",
    ylabel = string(L"RSL {BeM}", " [m]"),
    grid = false
```

• G_binned_fullength_hr500, for high resolution analysis with the Martinez-García record.

In [18]:

```
# Binned resampling on finer grid, for high resolution analyses with Martinez-Ga rcía

binsize_hr = 0.5 # time step 0.5 kyr
grid_G_hr = minimum(t_G) + binsize_hr/2 : binsize_hr : maximum(t_G) - binsize_hr
/2
# Intuitively, we would have the grid start at half a binsize *before* the first datapoint (binmidpoint).
# But that might skew the first point in the record a bit (unless we extrapolate for the first half of the bin, but that also imposes some assumptions).
# We therefore opt to cut the length of the timeseries by a datapoint in each en d, to avoid skewedness or extrapolation.

resampling_method_G_hr500 = BinnedResampling(grid_G_hr, 1000)

G_binned_fullength_hr500 = resample(uivD_G, resampling_method_G_hr500)
```

Out[18]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDataset} containing 983 uncertain values coupled with 983 uncertain indic

In [19]:

```
# Save the relevant arrays of the Elderfield record in a .jld2 file @save "../../MASTER_2.0/Koding/WrangledDataFiles/Binned_ts_fullength/Grant.jld2" G_binned_fullength G_binned_fullength_hr125 G_binned_fullength_hr500
```

```
In [20]:
```

```
@load "../Koding/WrangledDataFiles/Binned_ts_fullength/Grant.jld2" # This will b
e read in in our Toolbox notebook (NB3)

Out[20]:
3-element Array{Symbol,1}:
    :G_binned_fullength
    :G_binned_fullength_hr125
    :G_binned_fullength_hr500
```

1.4 Rohling (2014)

- denoted R
- Relative sea-level at straight of Gibraltar, record spanning 5.3 Ma.
- · Data available from... DOI...
- Note main text caveats and reasoning for which parts of the records will be used: Sapropelic intervals
 makes the record inapt for the the post-MPT. We could opt to exclude the sapropelic intervals, but this
 would call for interpolation over large time gaps, and compromise the confidence level of the results.
 Therefore, we leave the record as is, and interpret the results with a pinch of salt.

i) Read in data

In [159]:

```
# read in data
filepath Rohling = "../../MASTER_2.0/data/sea-level/Rohling(2014)/Rohling(2014)_
data fig2 v1 rslGib.txt"
# .txt file with column 1-6 from the published dataset.
rawD Rohling = readdlm(filepath Rohling, '\t', Any, '\r', skipstart = 8, dims =
(5331, 6))
# naming columns
# First two columns are Wang et al. Discontinuous due to sapropel intervals. We
will instead use the probabilistic analysis, column 3-6.
t R sapgap = rawD Rohling[:,1] # A: Med.RSL chronology (ky)
RSL_R_sapgap = rawD_Rohling[:,2] # B: Med RSL (m; gaps are sapropel intervals)
# ...we use the Probabilistic analysis of the Wang record, by Rohling et al. (20
14)
t R = rawD Rohling[:,3]
                              # chronology (ky)
RSL mean R = rawD Rohling[:,4] # Med. RSL MEDIAN
                                                       # (m; shifted so mean o
f last three ky = 0) (??)
q95 up R = rawD Rohling[:,5] # upper 95%CI quantile # Med. RSL upper bound
95% probability interval for the median (~equivalent to 2se)
q95 lo R = rawD Rohling[:,6] # lower 95%CI quantile # Med. RSL lower bound
95% probability interval for the median (~equivalent to 2se)
# plot Rohling
plot Rohling raw age =
plot(#title = "Mediterranean sea level stack (Rohling)",
   size = (1000, 200),
   xlabel = "Age (kyrs)",
   ylabel = "Sea level (m)")
plot!(t R , RSL mean R ,
   markersize = 0.5,
   ribbon = (RSL mean R \cdot - q95 lo R , q95 up R \cdot - RSL mean R ), # 95% CI
   color = :skyblue,
   label = "Rohling")
plot!(t R sapgap, RSL R sapgap)
```

Cannot convert SubString{String} to series data for plotting

Stacktrace:

- [1] error(::String) at ./error.jl:33
- [2] prepareSeriesData(::SubString{String}) at /Users/maria/.julia/packages/Plots/qZHsp/src/series.jl:14
- [3] convertToAnyVector(::SubString{String}, ::Dict{Symbol,Any}) at /Users/maria/.julia/packages/Plots/qZHsp/src/series.jl:26
- [4] (::getfield(Plots, Symbol("##152#155")){Dict{Symbol,Any}})(::Su
 bString{String}) at ./none:0
- [5] iterate(::Base.Generator{Array{Any,1},getfield(Plots, Symbol("#
 #152#155")){Dict{Symbol,Any}}}, ::Int64) at ./generator.jl:47
- [6] append_any(::Any, ::Vararg{Any,N} where N) at ./essentials.jl:7
 28
- [7] convertToAnyVector(::Array{Any,1}, ::Dict{Symbol,Any}) at /User s/maria/.julia/packages/Plots/qZHsp/src/series.jl:41
- [8] macro expansion at /Users/maria/.julia/packages/Plots/qZHsp/sr c/series.jl:129 [inlined]
- [9] apply_recipe(::Dict{Symbol,Any}, ::Type{Plots.SliceIt}, ::Array
 {Any,1}, ::Array{Any,1}, ::Nothing) at /Users/maria/.julia/packages/
 RecipesBase/zBoFG/src/RecipesBase.jl:275
- [10] _process_userrecipes(::Plots.Plot{Plots.PyPlotBackend}, ::Dict {Symbol,Any}, ::Tuple{Array{Any,1},Array{Any,1}}) at /Users/maria/.julia/packages/Plots/qZHsp/src/pipeline.jl:83
- [11] _plot!(::Plots.Plot{Plots.PyPlotBackend}, ::Dict{Symbol,Any},
 ::Tuple{Array{Any,1},Array{Any,1}}) at /Users/maria/.julia/packages/
 Plots/qZHsp/src/plot.jl:178
- [12] #plot!#143(::Base.Iterators.Pairs{Union{}, Union{}, Tuple{}, Name
 dTuple{(),Tuple{}}}, ::typeof(plot!), ::Plots.Plot{Plots.PyPlotBacke
 nd}, ::Array{Any,1}, ::Vararg{Array{Any,1},N} where N) at /Users/mar
 ia/.julia/packages/Plots/qZHsp/src/plot.jl:158
- [13] plot!(::Plots.Plot{Plots.PyPlotBackend}, ::Array{Any,1}, ::Var arg{Array{Any,1},N} where N) at /Users/maria/.julia/packages/Plots/qZHsp/src/plot.jl:155
- [14] #plot!#142(::Base.Iterators.Pairs{Union{},Union{},Tuple{},Name dTuple{(),Tuple{}}}, ::typeof(plot!), ::Array{Any,1}, ::Vararg{Array {Any,1},N} where N) at /Users/maria/.julia/packages/Plots/qZHsp/src/plot.jl:150
- [15] plot!(::Array{Any,1}, ::Array{Any,1}) at /Users/maria/.julia/packages/Plots/qZHsp/src/plot.jl:144
 - [16] top-level scope at In[159]:29

ii) Redefine from age to time by reversing the dataset

In [152]:

```
# Reverse the dataset
revD Rohling = reverse(rawD Rohling, dims = 1 )
#t R sapgap = -revD Rohling[:,1] # A: Med.RSL chronology (ky)
                                  # B: Med RSL (m; gaps are sapropel interval
#RSL R sapgap = revD Rohling[:,2]
s)
t R = -revD Rohling[:,3]
                               # chronology (ky)
                                                       # negative, so that time
runs forward with increasing indices
RSL mean R = revD Rohling[:,4] # Med. RSL MEDIAN
                                                       # (m; shifted so mean of
last three ky = 0) (??)
q95 up R = revD Rohling[:,5] # upper 95%CI quantile
                                                        # Med. RSL upper bound 9
5% probability interval for the median (~equivalent to 2se)
q95 lo R = revD Rohling[:,6] # lower 95%CI quantile
                                                        # Med. RSL lower bound 9
5% probability interval for the median (~equivalent to 2se)
# Reformulate the uncertainties from quantiles to standard deviance (1\sigma)
RSL 1\sigma R = (q95 up R .- q95 lo R) / 4
```

Age uncertainty has been incorporated in the RSL standard deviation through Monte Carlo probabilistic analysis.

In [162]:

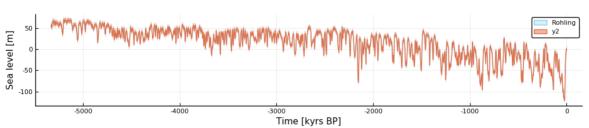
```
@save "../Koding/WrangledDataFiles/BasicArrays/Rohling.jld2" t_R RSL_mean_R q95_
lo_R q95_up_R
```

In [163]:

```
# plot Rohling
plot_Rohling_raw =
plot(#title = "Mediterranean sea level stack (Rohling)",
    size = (1000,200),
    xlabel = "Time [kyrs BP]",
    ylabel = "Sea level [m]")
plot!(t_R, RSL_mean_R,
    ribbon = (RSL_mean_R .- q95_lo_R, q95_up_R .- RSL_mean_R),
    fillalpha = 0.3,
    color = :skyblue,
    label = "Rohling")
#plot!(t_R, RSL_mean_R, ribbon = (2*RSL_1o_R) ) # Checking that same as above,
    95% CI. All good.

savefig("../../MASTER_2.0/figurar/RawData/GSL/plot_Rohling_raw.pdf")
```

Out[163]:



Interpolation

Note on sapropelic intervals

The Rohling record has a significant challenge, that it is riddled with **sapropelic intervals**. These intervals are anomalies in the GSL record, hypothesized to be periods of major surface freshwater dilution, and so the $\delta^{18}O_p$ signal is not indicative of sea level for these periods. For the Grant record, there were records keeping track of sapropelic intervals was kept for each individual datapoint ("using colour, core-scanning X-ray fluorescence(XRF), and magnetic data, in addition to stable isotope, organic carbon, and microfossil abundance data from previous studies"), allowing for a precise identification and exclusion of sapropelic intervals from the record. This does not exist for the Rohling record, however, and the *sapropelic intervals of the Mediterranean has instead been identified and removed through a signal processing approach.*

They also report that 3 sapropel intervals seem to not have been filtered out by the signal processing. These were detected by discord with other GSL records overlapping in time. A caveat to analyses with this record is that we presently have no way to know if there are more sapropelic intervals that has not been captured by the signal processing. Since this is the only high resolution GSL record we have that spans further back than Elderfield (1.5 Ma), such intervals would not be identified further back in the record.

would like to at least mark the sapropelic intervals, to showcase the problem. Have tried, but need help on how to.

```
In [160]:

t_R_sapgap;

In [161]:

using Missings
df = collect(Missings.replace(t_R_sapgap, NaN));
;
```

HELP. # MISSING VALUES in the first two columns (Sapropel intervals). Need to be dealt with #import Pkg; Pkg.add("Missings") using Missings Missings.replace(x, 1) Missings.EachReplaceMissing{Array{Union{Missing, Int64},1},Int64}(Union{Missing, Int64}[1, 2, missing], 1) julia> collect(Missings.replace(x, 1)) 3-element Array{Int64,1}: 1 2 1 julia> collect(Missings.replace(x, 1)) == coalesce.(x, 1) true

Missing values in the first two columns of the dataset (renamed sapgap) are sapropel intervals.

```
In [ ]:
plot(t_R_sapgap, diff(t_R_sapgap)) # Doesn't work because haven't replaced missi
ng values
```

Explore the resolution of the Rohling record when sapropelic intervals are cut out

```
### Explore the resolution of the Rohling record when sapropelic intervals are c
ut out ###
# Read in data from Rohling et al.
#### Note: I have cut out the sapropel intervals (marked by gaps in the first co
lumn) from this file.####
sapcutfile = "../../MASTER 2.0/data/sea-level/Rohling(2014)/Rohling(2014) data f
ig2 sapropelscutout.txt"
rawD Rohling sapcut = readdlm(sapcutfile,'\t', Any, '\r', skipstart = 8, dims =
(3736, 6))
# Reversing the dataset
revD Rohling sapcut = reverse(rawD Rohling sapcut, dims = 1 )# to convert from a
ge (increasing backwards) to time (increasing forward, with present defined as
 0)
# naming columns of interest
t R sapcut = -revD Rohling sapcut[:,3]
                                          # chronology (ky)
RSL mean R sapcut = revD Rohling sapcut[:,4] # Med. RSL MEDIAN
                                                                    # (m; shif
ted so mean of last three ky =0) (??)
q95 up R sapcut = revD Rohling sapcut[:,5] # upper 95%CI quantile
                                                                      # Med. RSL
upper bound 95% probability interval for the median (~equivalent to 2se)
q95 lo R sapcut = revD Rohling sapcut[:,6] # lower 95%CI quantile # Med. RSL
lower bound 95% probability interval for the median (~equivalent to 2se)
# Reformulate the uncertainties from quantiles to standard deviance (1\sigma)
RSL 1\sigma R sapcut = (q95 up R sapcut .- q95 lo R sapcut) / 4
# Explore the resolution of the Rohling record when sapropelic intervals are cut
print("The highest resolution after sapropel intervals cut out is ", minimum(dif
f(t R sapcut)), "
", "Average resolution after sapropel intervals are cut out is ", mean(diff(t R s
apcut)), "
", "Sapropelic intervals create gaps in the record of up to ", maximum(diff(t R s
apcut)), " kyrs.")
plot(t R sapcut, diff(t R sapcut),
   xlabel = "Age [kyrs BP]",
   ylabel = "time gap [kyrs]",
   title = "Resolution of Rohling sea-level record
   after cutting out reported sapropelic intervals ")
```

MA_NB1_PrepData

The resolution plot above shows the intervals that were filtered out by signal processing (Rohling et al. (2014).

As mentioned, *Rohling et al.(2014)* also report that 3 sapropel intervals seem to not have been filtered out by the signal processing. I did not find the data necessary to mark these intervals, but they are roughly at 450, 550 and 750 kyrs BP (marked in the figures of the Rohling article, and it is also evident in the comparison with the other GSL plots, which we will see below). These anomalies make the Rohling record inapt for analyses across the post-MPT.

Interpolation?

We could opt to exclude these "residual" sapropelic intervals. However, this would call for interpolation over large time gaps, since our method requires continuous time series, which would compromise the confidence level of the results/statistical power of the method. (How much interpolation is acceptable before a breakdown of statistical power in the method could be explored further with a sensitivity analysis, but that is outside the scope of this project.) Given that we do not have access to the exact time intervals of the three later identified sapropelic intervals, we instead leave the record *as is*, with the sapropelic intervals, and omit using it for the post-MPT. Additionally, since we cannot be sure there are other sapropelic intervals not captured by the signal processing, we interpret any results from analyses with the Rohling record with a pinch of salt. This is thus one of the records where we have to *bite i det sure eplet* and accept the limitations of our data.

Per now, there is one analysis with the Rohling record that gives significant results in the *opposite direction* from significant results with the other GSL records (Rohling \rightarrow Bereiter on the postMPT_500 grid). What to make of this, when both are *significant* results?

Sapropelic intervals are reportedly identified mainly for the first 700 years of the record, and we therefore planned not to include the Rohling record for analyses on the post-MPT.

Interpolation

The dataset was originally given on a 1 kyr regular time grid. If we cut out the sapropel intervals, there are gaps in the record we need to interpolate to run the analysis.

```
In [ ]:
```

```
# We have decided not to cut out the 3 sapropelic intervals
# Evaluate need for interpolation by checking the resolution

minimum(diff(t_R)) #1
mean(diff(t_R)) #1
maximum(diff(t_R)) #1

plot(diff(t_R),
xlabel = "index",
ylabel = "time gap between observations (kyrs)",
label = "diff(t_R)")
```

The Rohling record is already on a regular 1 kyr grid, so no need for interpolation.

Redefining revD Rohling as an UncertainIndexValueDataset

In []:

```
# Refining data as an ``UncertainIndexValueDataset``

t_uiv_R = [UncertainValue(Normal, t_R[i], 0) for i in 1:length(t_R)]

RSL_uiv_R = [UncertainValue(Normal, RSL_mean_R[i], RSL_lo_R[i]) for i in 1:length(RSL_mean_R)] #HELP ON DEFINING intpD_RSL_mean_R

uivD_R = UncertainIndexValueDataset(t_uiv_R, RSL_uiv_R);
```

In []:

```
plot(uivD_R)
```

In []:

Save the relevant arrays of the Rohling record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Rohling.jld2" t_R RSL_mea
n_R RSL_1 uivD_R

vi) Binned resampling

In []:

```
grid_R = t_R[1] - binsize/2 : binsize : t_R[end] + binsize/2
#= dette er ein ryddigare og finare måte å definere på
men då må eg kjøre alle binned resamplings på nytt
og ikkje noko problem med sånn det er gjort allereie
forutan at eg mister eit datapunkt i kvar ende
så kanskje berre Grant og Chalk det har noko for seg å redefinere =#
```

In []:

```
\label{eq:grid_R} \begin{split} &\text{grid}_{-R} = \text{ceil}(\text{minimum}(\text{t}_{-R})) + \text{binsize}/2 : \text{binsize} : \text{floor}(\text{maximum}(\text{t}_{-R})) - \text{binsiz} \\ &\text{e}/2 \end{split}
```

In []:

```
resampling_method_R = BinnedResampling(grid_R, 1000) # 1000 draws in each bin.
R_binned_fullength = resample(uivD_R, resampling_method_R)
@save "../Koding/WrangledDataFiles/Binned_ts_fullength/Rohling.jld2" R_binned_fullength
```

```
@load "../Koding/WrangledDataFiles/Binned_ts_fullength/Rohling.jld2"
```

In []:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
R = R binned full
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(R.values, 0.5)
bin upper = quantile.(R.values, 0.975) .- bin median
bin lower = bin median .- quantile.(R.values, 0.025)
# time array
binmidpoints R = [R.indices[i].value for i in 1:length(R)]
plot R binned =
plot(binmidpoints R, bin median,
    ribbon = (bin lower, bin upper),
    color = :skyblue,
    label = "Rohling",
    xlabel = "Time [years BP]",
    ylabel = string(L"RSL {Gib}", " [m]"),
    size = (1000, 200),
    grid = false
```

Comparative plots between GSL records

Comparative plot between SprattLisiecki and Rohling

```
# Comparative plot between SprattLisiecki and Rohling
plot(#title = "Comparison of Spratt&Lisiecki's & Rohling's GSL records",
    xlabel = "Time (kyrs BP)",
    ylabel = "Sea level (m)",
    size = (1000, 200),
    xtics = (0:100:1000),
    xlim = (-1000,0)) \# Plot only last 1 Myr BP
plot!(t R, RSL mean R,
    markersize = 0.5,
                                                                             # 95%
    ribbon = (RSL mean R .- q95 lo R, q95 up R .- RSL mean R),
CI
    color = :skyblue,
    label = "Rohling")
plot!(t SL, GSL mean SL,
    markersize = 0.5,
    ribbon = (GSL mean SL - GSL SL err lo, GSL SL err up - GSL mean SL),
CI
    # xerr = t \sigma SL, # age uncertainty from Jo (speleothem tuning of record) - U
SE OR NOT?
    color = :darkblue,
    label = "Spratt & Lisiecki")
```

Large deviations in GSL dynamics between the records around 750, 550 and 450 kyrs BP. These seem to be concordant in timing with the potential bands of sapropelic intervals mentioned in Rohling et al. (2014) as not having been filtered out by the signal-processing (marked in yellow in the article, but I couldn't find the indices in the dataset.)

In any case, this illustrates the point that the GSL records available are *significantly different*. That is the reason why we will run our analyses with many different records for GLS, so that we can make more robust conclusions.

Why are the uncertainties (confidence interval) so very different?

```
# Plot to compare all GSL reconstructions
plot GSL comparative 1500kyrs =
plot(#title = "Comparative plot of GSL records",
    xlabel = "Time [ka BP]",
    ylabel = "GSL [m]",
    grid = true,
                          # not working
    xlim = (-1500, 0), \# plot only a range 0:1000 ka
    ticks = (0:100:-1000), # not working
    size = (1000, 200),
    share = :(x,y),
                         # working?
    #legend = :bottomleft
plot!(t R, RSL mean R,
    ribbon = (2*RSL_1\sigma_R), # \pm 2\sigma, aka 95% confidence interval
    color = :skyblue,
    label = "Rohling"
#plot!(fine grid E, intpD GSL E, yerr = (2*t 1\sigma E), color = :royalblue, ms = 0.
1) # LR04 age model uncertainty
plot!(fine grid E, intpD GSL E,
    ribbon = (2*intpD 1\sigma E),
    color = :royalblue,
    label = "Elderfield"
#plot!(t SL, GSL mean SL, yerr = (2*t 1\sigma SL), color = :darkblue, ms = 0.1) # LR0
4 age model uncertainty
plot!(t SL, GSL mean SL,
    ribbon = (GSL mean SL - GSL SL err lo, GSL SL err up - GSL mean SL), # 95% C
Ι
    color = :darkblue,
    label = "Spratt & Lisiecki"
plot!(t G, RSL G,
                                                                            # 95% C
    ribbon = (RSL G - RSL q95lo G, RSL q95up G - RSL G),
Ι
    color = :cyan,
    label = "Grant"
savefig("../figurar/RawData/GSL/plot GSL comparative 1500kyrs.pdf")
```

In []:

```
# Compare GSL records with d180 record

plot(plot_GSL_comparative_full,
    plot_LR04_raw_time,
    layout = grid(2,1),
    link = :x)
```

In []:

help on making it pretty (common xlabel, and maybe ylabel)

3 - Insolation

- Insolation record La2004, from Laskar et al. (2004) (denoted Ins).
- Numerical solution for midsummer insolation at 65°N
- Time series computed using the AnalySeries computer software (Paillard et al., 1994)

In [43]:

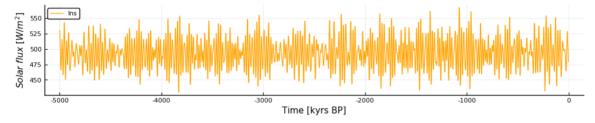
```
# Read in La2004

filepath_insol = "../../MASTER_2.0/data/insolation/Paillard_computation/Ins_65N_
Daily_ins21June_65N_past5Ma" # (nB) - La2004 solution, computed using AnalySerie s
rawD_insol = readdlm(filepath_insol, dims = (50001,2), skipstart = 2)

# reverse columns, to redefine from age (ka BP, increasing backwards) to time (k yrs BP, increasing forwards)
revD_insol = reverse(rawD_insol, dims = 1)

# name columns of interest
t_insol = -revD_insol[:,1] # time [kyrs BP]. Made negative to define 0 as present
insol_65N = revD_insol[:,2] # insolation: daily mean TOA solar flux at 65N summ er solstice, [W/m^2]
@save "../Koding/WrangledDataFiles/BasicArrays/La2004.jld2" t_insol insol_65N
```

In [44]:



iv) Interpolation to time grids equivalent to time steps we will use

Assigning the time series to grids with time steps we will use in analyses (1 kyr, 500 yrs, 125 yrs).

In [45]:

```
# interpolation on grids with time steps we will use
# Make a continuous function of interpolated values
interpolate t insol = LinearInterpolation(t insol, t insol)
interpolate insol 65N = LinearInterpolation(insol 65N, t insol)
grid insol 1 = t insol[1] : 1 : t insol[end] # 1 kyr, common time step for all a
nalyses
# Arrays with interpolated values for timesteps of 500 years
La2004 t fullength
                        = [interpolate t insol(i) for i in grid insol 1]
# time array, interpolated to high resolution
La2004 insol65N fullength = [interpolate insol 65N(i) for i in grid insol 1] #
insolation values array, interpolated for 500 year timestep analyses
binsize hr500 = 0.5 # for analyses with Lambert and Martinez-Garcia
grid insol hr500 = t insol[1] : binsize hr500 : t insol[end] # we set the grid f
or the interpolated data to give values concordant with the Grant and Chalk reco
rds for the high resolution analyses
# Arrays with interpolated values for timesteps of 500 years
La2004 t fullength hr500
                               = [interpolate t insol(i) for i in grid insol hr
            # time array, interpolated to high resolution
5001
La2004 insol65N fullength hr500 = [interpolate insol 65N(i) for i in grid insol
hr500] # insolation values array, interpolated for 500 year timestep analyses
# make a grid onto which we assign values from the interpolated function
binsize hr = 0.125 # for analyses with Grant and Chalk
grid insol cr = t insol[1] : binsize hr : t insol[end]
# Arrays with interpolated values for timesteps of 125 years
La2004 t fullength hr125
                               = [interpolate t insol(i) for i in grid insol cr
         # time array, interpolated to high resolution
La2004 insol65N fullength hr125 = [interpolate insol 65N(i) for i in grid insol
cr] # insolation values array, interpolated for 125 year timestep analyses
;
```

v) carry on arrays to analysis Note: Computed insolation values has no associated uncertainty on this timescale (certain up to >50 Ma into past and present). We therefore have no need in redefining the insolation data as UncertainIndexValueDataset. We can instead carry on the arrays t_insol and insol 65 directly to analysis. For high resolution analysis we carry on the equivalent interpolated arrays.

```
In [47]:
```

```
# Save the relevant arrays of the La2004 record in a .jld2 file
#### These are the arrays we will carry on for insolation:
# timestep 1 kyr - standard grid for analysis with all records
La2004 t fullength
                            # age array to define common time interval
La2004 insol65N fullength
                            # ... insolation values to be sent to causality tes
# interpolated arrays for high resolution analysis
# timestep 500 years - for hr analyses with Lambert, Martinez-Garcia and Bereite
La2004 t fullength hr500
La2004 insol65N fullength hr500
# timestep 125 years - for hr analysis with Chalk and Grant records
La2004 t fullength hr125
La2004 insol65N fullength hr125
@save "../../MASTER 2.0/Koding/WrangledDataFiles/La2004.jld2" La2004 t fullength
La2004 insol65N fullength La2004 t fullength hr125 La2004 insol65N fullength hr
125 La2004 t fullength hr500 La2004 insol65N fullength hr500
```

In [48]:

```
# checking that all 6 arrays included in the file to be carried on in NB3
@load "../../MASTER_2.0/Koding/WrangledDataFiles/La2004.jld2"
# checking, all good
```

Out[48]:

```
6-element Array{Symbol,1}:
:La2004_t_fullength
:La2004_insol65N_fullength
:La2004_t_fullength_hr125
:La2004_insol65N_fullength_hr125
:La2004_t_fullength_hr500
:La2004_insol65N_fullength_hr500
```

4 - pCO2 records

These are the reference records we will use for pCO2:

- Bereiter et al (2015), spanning the last 800 kyrs
- Chalk et al (2017), spanning the interval from 1.090 1.240 Ma
- Hönisch et al (2009), spanning the last 2 Myrs.

4.1 Bereiter pCO2 record

- pCO2 record from Bereiter et al. (2015), denoted B
- Direct measurement of pCO2 concetration in air bubbles from Epica Dome C (EDC) ice core.
- · denoted B
- age model AICC2012
- Data available from... got the file from Jo, where to reference?

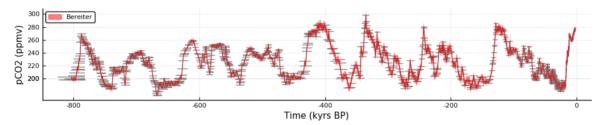
In [3]:

```
# Read in data from Bereiter
# AICC2012 age model
filepath Bereiter = "../../MASTER 2.0/data/CO2/Bereiter2015 CO2 ageunc.txt"
rawD B = readdlm(filepath Bereiter, '\t', Float64,'\n', dims = (1643,4), skipsta
rt = 1)# reading in the file
revD B = reverse(rawD B, dims = 1) # reversing along the first dimension (rows)
# naming columns
t B = -(revD B[:,1] /1000) # 1st column = Time (kyrs) - divided by 1000 to chang
e unit from years to kyrs. Negative because we define present as 0 kyrs.
t 1 o B = revD B[:,2] /1000 # 2nd column = age model uncertainty - divided by 100
0 to change unit from years to kyrs. \sigma remains a positive value
CO2 mean B = revD B[:,3] # 3rd column = pCO2 (ppmv) - remains in positive valu
CO2 1\sigma B = revD B[:,4] # 4th column = pCO2 uncertainty (ppmv) - \sigma remains a
positive value
@save "../Koding/WrangledDataFiles/BasicArrays/Bereiter.jld2" t B t 1\sigma B CO2 mea
n B CO2 1σ B
```

In [4]:

```
# overview plot of pCO2
@load "../Koding/WrangledDataFiles/BasicArrays/Bereiter.jld2" t_B t_1\u00f3_B CO2_mea
n B CO2 1\sigma B
#plot Bereiter raw ageunc =
plot(#title = "800 kyrs of pCO2",
    size = (1000, 200),
    xlabel = "Time (kyrs BP)",
    ylabel = "pCO2 (ppmv)",
    legend = :topleft)
plot!(#title = "Bereiter pCO2 record, from Epica Dome C ice core"
    t B, CO2 mean B, xerr = (2 * t 1\sigma B), # use 1 or 2 \sigma for?
    ms = 0.1, color = :grey,
    label = "", #"Age model uncertainty (\pm 2\sigma)
plot!(t B, CO2 mean B,
    ribbon = (2 * CO2 10 B), # 95% CI
    color = "red",
    alpha = 0.5,
    label = "Bereiter"
    )
```

Out[4]:



In [40]:

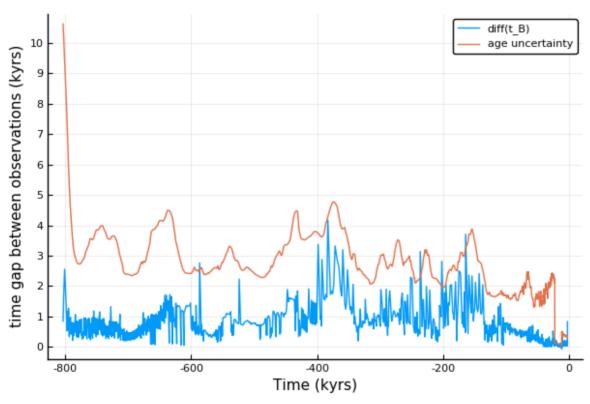
```
savefig("../../Master_2.0/figurar/RawData/pCO2/Bereiter_raw_timeunc.pdf")
```

iv) Interpolation

First, we check if there is need for interpolation of this record (We have decided a 1 kyr resolution for our analyses)

In [5]:

Out[5]:



The Bereiter record has irregular resolution, with gaps of up to 4000 years between observations. However, the age uncertainty is large enought to get some values in each bin through the binned resampling.

With a mean uncertainty of 490 years, we may also try out higher resolution analyses. We will however set the mean resolution as a lower boundary of the time steps used in analyses.

Redefining as UncertainIndexValueDataset

We redefine the arrays to an uivD, which carries the uncertainties as kernel density estimates (KDE). The binned resampling function can then resample from the probability distribution.

1. uivD including age model uncertainty

In [6]:

```
# Redefining as uivD_B

t_uiv_B = [UncertainValue(Normal, t_B[i], t_1\sigma_B[i]) for i in 1:length(t_B)]
CO2_uiv_B = [UncertainValue(Normal, CO2_mean_B[i], CO2_1\sigma_B[i]) for i in 1:lengt
h(CO2_mean_B)]
uivD_B = UncertainIndexValueDataset(t_uiv_B, CO2_uiv_B)
```

Out[6]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas et} containing 1643 uncertain values coupled with 1643 uncertain ind ices

Note: The Bereiter (pCO2) and Lambert (dust) records are both from the same ice core site (EDC). For analyses between these two time series, we can therefore exclude uncertainties in the age model. We therefore also prepare a version of the uivD that doesn't include uncertainties in age.

1. uivD without age model uncertainty

In [7]:

```
# uivD_B_EDC defined without age uncertainty, for analysis with other records fr
om the EDC core.

t_uiv_B_EDC = [UncertainValue(Normal, t_B[i], 0) for i in 1:length(t_B)] # age m
odel uncertainty not included

CO2_uiv_B = [UncertainValue(Normal, CO2_mean_B[i], CO2_1\sigma_B[i]) for i in 1:lengt
h(CO2_mean_B)]
uivD_B_EDC = UncertainIndexValueDataset(t_uiv_B_EDC, CO2_uiv_B)

# Save the relevant arrays of the Bereiter record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Bereiter_nointp.jld2" uiv
D_B uivD_B_EDC
```

```
In [ ]:
```

```
# plot(uivD_B)
```

vi) Binned resampling on grid

```
In [8]:
```

:uivD_B
:uivD_B_EDC

```
@load "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Bereiter_nointp.jld2"
Out[8]:
2-element Array{Symbol,1}:
```

In [9]:

```
uivD_B
#= Be sure to use the original data array
(original array (without interpolation) has 1643 datapoints
interpolated array has 8018 datapoints
=#
```

Out[9]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas et} containing 1643 uncertain values coupled with 1643 uncertain ind ices

1. Binned resampling with age model uncertainty

In [13]:

```
# binned resamplig on a 1 kyr grid
binsize = 1 # 1 kyr
tmin_B = ceil(minimum(t_B)) # first bin midpoint at a whole kyr
tmax_B = floor(maximum(t_B)) # last bon midpoint at a whole kyr
grid B = tmin B - binsize/2 : binsize : tmax B + binsize/2 # define the grid by
the ebin edges
resampling method B = BinnedResampling(grid B, 1000)
## standard version (1 kyr grid, with full uncertainties)
@time B binned fullength ageunc = resample(uivD B, resampling method B)
### Binned resampling on a 500 yr grid, for high resolution analysis of time ste
p 500 yrs with Grant, La2004, Lambert, Martinez-García
binsize_hr500 = 0.5 # mean resolution is one observation every 490 years
tmin_B = ceil(minimum(t_B)) # first bin midpoint at a whole kyr
tmax B = floor(maximum(t B)) # last bon midpoint at a whole kyr
grid B hr500 = tmin B - binsize hr500/2 : binsize hr500 : tmax B + binsize hr500
/2 # define the grid by bin edges
resampling method B hr500 = BinnedResampling(grid B hr500, 1000) # resample 1000
values within each bin of the grid.
@time B binned fullength ageunc hr500 = resample(uivD B, resampling method B hr5
00)
31.571611 seconds (34.45 M allocations: 7.322 GiB, 8.51% gc time)
28.384372 seconds (29.06 M allocations: 12.474 GiB, 13.74% gc time)
```

UndefVarError: resampling method not defined

Stacktrace:

- [1] top-level scope at util.jl:156
- [2] top-level scope at In[13]:33

In [14]:

```
#### Binned resampling on a 125 yr grid, for high resolution analysis of time s
tep 125 yrs with Grant, La2004,
binsize_hr125 = 0.125 # mean resolution is one observation every 490 years
tmin_B = ceil(minimum(t_B)) # first bin midpoint at a whole kyr
tmax_B = floor(maximum(t_B)) # last bon midpoint at a whole kyr
grid_B_hr125 = tmin_B - binsize_hr125/2 : binsize_hr125 : tmax_B + binsize_hr125
/2 # define the grid by bin edges

resampling_method_B_hr125 = BinnedResampling(grid_B_hr125, 1000) # resample 1000
values within each bin of the grid.
@time B_binned_fullength_ageunc_hr125 = resample(uivD_B, resampling_method_B_hr1
25)
```

```
68.614789 seconds (64.92 M allocations: 46.878 GiB, 19.33% gc time)
```

Out[14]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 6409 uncertain values coupled with 6409 uncertain ind
ices

1. Binned resampling without age model uncertainty.

Version without age model uncertainties is for analysis with other EDC ice core dust record (Lambert / IceDust)

In [15]:

```
#### version without age model uncertainties, for analysis with other EDC ice co
re dust record (Lambert IceDust)

# binned on the 1 kyr grid
@time B_binned_fullength_noageuncEDC_noIntp = resample(uivD_B_EDC, resampling_me
thod_B)

# binned on the 500 yr grid
@time B_binned_fullength_noageuncEDC_hr500_noIntp = resample(uivD_B_EDC, resampl
ing_method_B_hr500)

#binned on the 125 yr grid
#@time B_binned_fullength_noageuncEDC_hr125_noIntp = resample(uivD_B_EDC, resampl
ling_method_B_hr125)
```

```
6.276096 seconds (5.43 M allocations: 5.816 GiB, 24.78% gc time) 12.876414 seconds (11.52 M allocations: 11.521 GiB, 25.87% gc time)
```

Out[15]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 1603 uncertain values coupled with 1603 uncertain ind
ices

```
In [18]:
@time B binned fullength noageuncEDC_hr125_noIntp = resample(uivD_B_EDC, resampl
ing method B hr125)
 45.891576 seconds (42.39 M allocations: 45.387 GiB, 25.86% gc time)
Out[18]:
UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 6409 uncertain values coupled with 6409 uncertain ind
ices
In [19]:
@save "../../MASTER 2.0/Koding/WrangledDataFiles/Binned ts fullength/Bereiter no
intp.jld2" B binned fullength ageunc B binned fullength ageunc hr500 B binned fu
llength ageunc hr125 B binned fullength noageuncEDC noIntp B binned fullength no
ageuncEDC hr500 noIntp B binned fullength noageuncEDC hr125 noIntp
In [20]:
@load "../../MASTER 2.0/Koding/WrangledDataFiles/Binned ts fullength/Bereiter no
intp.jld2"
Out[20]:
6-element Array{Symbol,1}:
 :B binned fullength ageunc
 :B binned fullength ageunc hr500
 :B binned fullength ageunc hr125
```

Plot the binned resampled time series with the 95% confidence interval

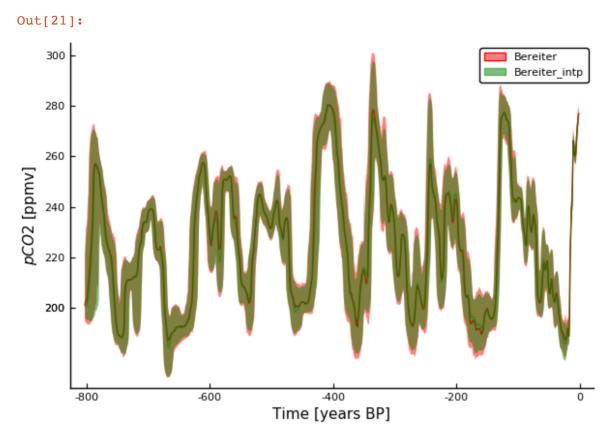
1. With the AICC2012 age model uncertainty

:B binned fullength noageuncEDC noIntp

:B_binned_fullength_noageuncEDC_hr500_noIntp
:B binned fullength noageuncEDC hr125 noIntp

In [21]:

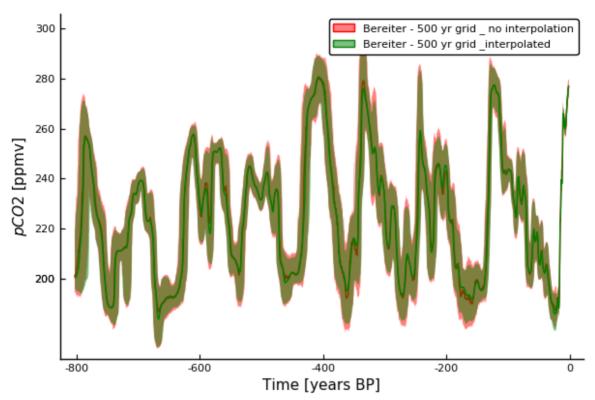
```
# 1 kyr grid, full age uncertainty,
########## no interpolation
@load "../../MASTER 2.0/Koding/WrangledDataFiles/Binned ts fullength/Bereiter no
intp.jld2" # no interpolation
### Plot the binned resampled uivD time series with the 95% confidence interval
B = B binned fullength ageunc
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
# time array
binmidpoints B = [B.indices[i].value for i in 1:length(B)]
#plot B binned ageunc =
plot(binmidpoints B, bin median,
   ribbon = (bin lower, bin upper),
   color = :red.
   label = "Bereiter",
   xlabel = "Time [years BP]"
   ylabel = string(L"pCO2", " [ppmv]"),
   grid = false
########## w/interpolation
@load "../../MASTER 2.0/Koding/WrangledDataFiles/Binned ts fullength/Bereiter.jl
d2" # with interpolation
### Plot the binned resampled uivD time series with the 95% confidence interval
B = B binned fullength ageunc
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
# time array
binmidpoints B = [B.indices[i].value for i in 1:length(B)]
#plot B binned ageunc intp =
plot! (binmidpoints B, bin median,
   ribbon = (bin lower, bin upper),
   color = :green, alpha = 0.5,
   label = "Bereiter intp",
   xlabel = "Time [years BP]"
   ylabel = string(L"pCO2", " [ppmv]"),
   grid = false
    )
```



In [117]:

```
# 500 yr grid, with age uncertainty
##### without interpolation,
@load "../../MASTER 2.0/Koding/WrangledDataFiles/Binned ts fullength/Bereiter no
intp.jld2"
### same time series without age uncertainty
B = B binned fullength ageunc hr500
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
binmidpoints B = [B.indices[i].value for i in 1:length(B)]# time array
plot(binmidpoints B, bin median,
   ribbon = (bin lower, bin upper),
   color = :red,
   label = "Bereiter - 500 yr grid no interpolation",
   xlabel = "Time [years BP]",
   ylabel = string(L"pCO2", " [ppmv]"),
   grid = false
    )
####### with interpolation
@load "../../MASTER 2.0/Koding/WrangledDataFiles/Binned ts fullength/Bereiter.jl
d2"
### same time series without age uncertainty
B = B binned fullength ageunc hr500
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
binmidpoints B = [B.indices[i].value for i in 1:length(B)]# time array
plot! (binmidpoints B, bin median,
   ribbon = (bin lower, bin upper),
   color = :green,
   label = "Bereiter - 500 yr grid interpolated",
   xlabel = "Time [years BP]"
   ylabel = string(L"pCO2", " [ppmv]"),
   grid = false
    )
```





These two versions give slightly different results in the predictive asymmetry analysis between Grant GSL and Bereiter CO2. Both give a mean predictive asymmetry from BerCO2 to GraSL, but

- with interpolation, confidence intervals overlap, leaving some chance (maybe 20%) that the predictive asymmetry is unsignificant.
- without interpolation, confidence intervals diverge completely, leaving an unambiguous indication of unidirectional coupling

2. Without age model uncertainty

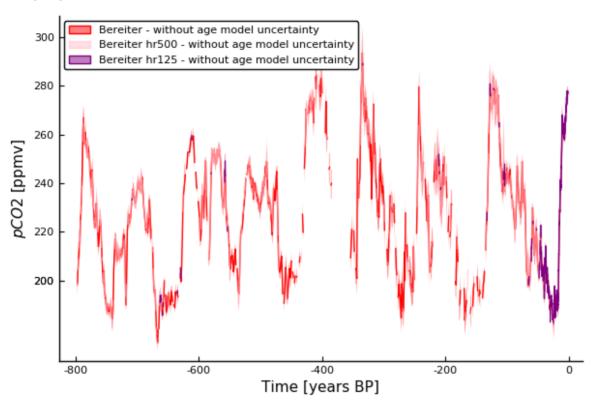
For analyses between EDC pCO2 and EDC dust - see that this calls for interpolation of data...

In [31]:

```
# 1 kyr, no age uncertainty
##### without interpolation
@load "../../MASTER 2.0/Koding/WrangledDataFiles/Binned_ts_fullength/Bereiter_no
intp.jld2"
B = B binned fullength noageuncEDC noIntp
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
# time array
binmidpoints B = [B.indices[i].value for i in 1:length(B)]
#plot B binned noageuncEDC intp =
plot(binmidpoints B, bin median,
    ribbon = (bin lower, bin upper),
    color = :red,
    label = "Bereiter - without age model uncertainty",
    xlabel = "Time [years BP]",
    ylabel = string(L"pCO2", " [ppmv]"),
    grid = false
    )
#### 500 yr grid
B = B binned fullength noageuncEDC hr500 noIntp
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
# time array
binmidpoints B = [B.indices[i].value for i in 1:length(B)]
plot!(binmidpoints B, bin median,
    ribbon = (bin lower, bin upper),
    color = :pink, alpha = 0.5,
    label = "Bereiter hr500 - without age model uncertainty",
    xlabel = "Time [years BP]",
    ylabel = string(L"pCO2", " [ppmv]"),
    grid = false
    )
#### 125 yr grid
B = B binned fullength noageuncEDC hr125 noIntp
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
# time array
binmidpoints B = [B.indices[i].value for i in 1:length(B)]
plot! (binmidpoints B, bin median,
    ribbon = (bin lower, bin upper),
    color = :purple,
    label = "Bereiter hr125 - without age model uncertainty",
    xlabel = "Time [years BP]",
    ylabel = string(L"pCO2", " [ppmv]"),
    grid = false
```

)
####### NEED FOR INTERPOLATION

Out[31]:



We need to interpolate values, to run the analysis without age uncertainty.

Interpolation of record when defined without age model uncertainty

Bereiter pCO2 and Lambert dust record are both from the EDC ice core. To better resolve the causal relationship between dust and pCO2, we can run analysis between these two records with the age uncertainty from lock-in depth of gas in ice, which is much smaller than the full age model uncertainty.

However, this calls for interpolation of some values on the Bereiter record, since there are gaps in the record that, when removing the age uncertainty, can no longer be covered in the binned resampling of data. To avoid over-interpolation, ee use the mean resolution of the record as the lower boundary for interpolation of data, which is 490 years for the Bereiter record.

In [179]:

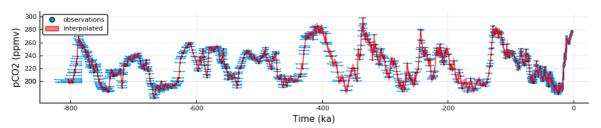
```
# interpolation of Bereiter data to ensure we have data in all bins
# defining interpolation functions
t B interpolate
                      = LinearInterpolation(t B, t B );
t_1\sigma_B_{interpolate}
                     = LinearInterpolation(t_1σ_B, t_B);
CO2 mean B interpolate = LinearInterpolation(CO2 mean B , t B );
CO2 1\sigma B interpolate = LinearInterpolation(CO2 1\sigma B , t B );
# make a time grid to contain the interpolated values for the Bereiter dataset.
fine grid B = minimum(t B): 0.1: maximum(t B) #### centennial resolution when
interpolated values are carried on
print(fine grid B)
# we then make arrays of the interpolated values corresponding to the bins of th
e fine time grid
intpD t B
                 = [t B interpolate(i) for i in fine grid B]
               = [t 1\sigma B interpolate(i) for i in fine grid B];
intpD t 1σ B
intpD CO2 mean B = [CO2 mean B interpolate(i) for i in fine grid B];
intpD CO2 1\sigma B = [CO2 1\sigma B interpolate(i) for i in fine grid B];
```

-803.70925:0.1:-2.00925

In [180]:

```
# plot interpolation to control
# NOTE: COMPUTATIONALLY HEAVY CELL when plotting interpolated age uncertainties
plot intpD B =
plot(#title = "Bereiter interpolation on fine grid",
    xlabel = "Time (ka)",
    ylabel = "pCO2 (ppmv)",
    size = (1000, 200))
scatter!(t_B, CO2_mean_B,
    \#ribbon = (CO2\_1\sigma\_B, CO2\_1\sigma\_B),
    xerr = 2 * t 1\sigma B,
    ms = 1,
    label = "observations")
plot!(intpD t B, intpD CO2 mean B,
    ribbon = (2 * intpD CO2 1\sigma B),
    color = "red",
    label = "interpolated")
#plot!(fine grid B, interpD CO2 mean B,
    # xerr = interpD t 2\sigma B, # plotting xerr is computationally heavy, and drown
s out plot info. Better to plot on "observations"
    \# ms = 0.1,
    # label = "age uncertainty",
    # color = "red")
```

Out[180]:



Define the interpolated data as an uivD

```
In [181]:
```

```
# uivD_B_EDC defined without age uncertainty, for analysis with other records fr om the EDC core.

t_uiv_B_EDC = [UncertainValue(Normal, intpD_t_B[i], 0) for i in 1:length(intpD_t_B)] # age model uncertainty not included

CO2_uiv_B = [UncertainValue(Normal, intpD_CO2_mean_B[i], intpD_CO2_lo_B[i]) for i in 1:length(intpD_CO2_mean_B)]

uivD_B_EDC = UncertainIndexValueDataset(t_uiv_B_EDC, CO2_uiv_B)

# Save the relevant arrays of the Bereiter record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Bereiter_noageuncEDC_int
p.jld2" uivD_B_EDC
```

In [174]:

```
length(t_uiv_B_EDC)
```

Out[174]:

1604

In [184]:

```
mean_res = mean(diff(t_B))
print("To limit the effect of false values, we do not interpolate values below t
he mean resolution of the record (", mean_res, " kyrs).
The highest resolution analyses with th is record will therefore be with the 500
yr time step.")
```

To limit the effect of false values, we do not interpolate values be low the mean resolution of the record ($0.48825185140073074~\rm kyrs$). The highest resolution analyses with th is record will therefore be with the 500 yr time step.

Binned resampling on the 1 kyr grid

In [185]:

```
# version without age model uncertainties, for analysis with other EDC ice core
  record (Bereiter pCO2)

# define the 1 kyr grid
binsize = 1 # 1 kyr
tmin_B = ceil(minimum(t_B)) # first bin midpoint at a whole kyr
tmax_B = floor(maximum(t_B)) # last bon midpoint at a whole kyr
grid_B = tmin_B - binsize/2 : binsize : tmax_B + binsize/2 # define the grid by
  the bin edges

# binned resamplig on the 1 kyr grid
resampling_method = BinnedResampling(grid_B, 1000) # draw 1000 samples in each b
in (uncertainties are now only in value)
@time B_binned_fullength_noageuncEDC = resample(uivD_B_EDC, resampling_method_B)
```

32.186467 seconds (26.09 M allocations: 28.265 GiB, 29.33% gc time)

Out[185]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas et} containing 802 uncertain values coupled with 802 uncertain indic es

Binned resampling on a higher resolution time grid (one value for every 500 years)

In [186]:

```
# Binned resampling on a 500 yr grid, for high resolution analysis with Lambert
EDC record (low age uncertainty)

binsize_hr500 = 0.5 # mean resolution is one observation every 490 years
tmin_B = ceil(minimum(t_B)) # first bin midpoint at a whole kyr
tmax_B = floor(maximum(t_B)) # last bon midpoint at a whole kyr
grid_B_hr500 = tmin_B - binsize_hr500/2 : binsize_hr500 : tmax_B + binsize_hr500
/2 # define the grid by bin edges

# binned resampling on the 500 yr grid
resampling_method_B_hr500 = BinnedResampling(grid_B_hr500, 1000) # resample 1000
values within each bin of the grid.
@time B_binned_fullength_noageuncEDC_hr500 = resample(uivD_B_EDC, resampling_method_B_hr500)
```

73.369700 seconds (51.91 M allocations: 55.877 GiB, 26.66% gc time)

Out[186]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 1603 uncertain values coupled with 1603 uncertain ind
ices

In [187]:

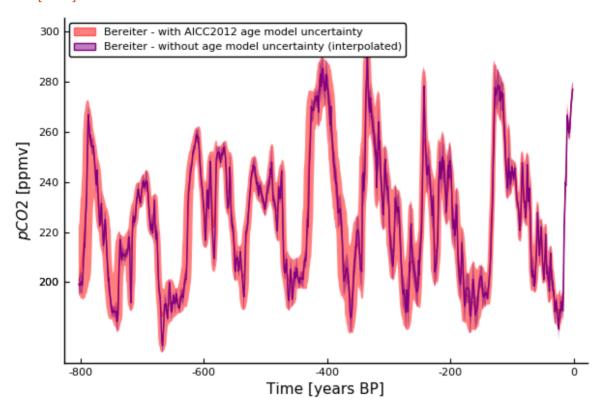
```
# Save the relevant arrays of the Bereiter record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/Binned_ts_fullength/Bereiter_no
ageuncEDC_intp.jld2" B_binned_fullength_noageuncEDC B_binned_fullength_noageuncE
DC_hr500
```

• Visualize the difference between the records with and without age model uncertainty

In [159]:

```
# 1 kyr grid, full age uncertainty,
######### no interpolation
@load "../../MASTER 2.0/Koding/WrangledDataFiles/Binned ts fullength/Bereiter no
intp.jld2" # no interpolation
### Plot the binned resampled uivD time series with the 95% confidence interval
B = B binned fullength ageunc
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
# time array
binmidpoints B = [B.indices[i].value for i in 1:length(B)]
#plot B binned ageunc =
plot(binmidpoints B, bin median,
   ribbon = (bin lower, bin upper),
   color = :red, alpha = 0.5,
   label = "Bereiter - with AICC2012 age model uncertainty",
   xlabel = "Time [years BP]",
   ylabel = string(L"pCO2", " [ppmv]"),
   grid = false
########## w/interpolation
@load "../../MASTER 2.0/Koding/WrangledDataFiles/Binned_ts_fullength/Bereiter_no
ageuncEDC intp.jld2" #with interpolation
### Plot the binned resampled uivD time series with the 95% confidence interval
B = B binned fullength noageuncEDC
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(B.values, 0.5)
bin upper = quantile.(B.values, 0.975) .- bin median
bin lower = bin median .- quantile.(B.values, 0.025)
# time array
binmidpoints B = [B.indices[i].value for i in 1:length(B)]
#plot B binned ageunc intp =
plot! (binmidpoints B, bin median,
   ribbon = (bin lower, bin upper),
   color = :purple, alpha = 1,
   label = "Bereiter - without age model uncertainty (interpolated)",
   xlabel = "Time [years BP]",
   ylabel = string(L"pCO2", " [ppmv]"),
   grid = false, legend = :topleft
    )
```

Out[159]:



4.2 - Chalk pCO2 record

- pCO2 record from Chalk et al. (2017), denoted C
- d11B proxy record for pCO2 spanning 1.090 1.240 Ma, (this is the only high resolution pCO2 record that is synchronous to the MPT).
- Data available from...DOI...

i) Read in data

Age [ka BP] CO2 [ppmv] (median CO2 (ppm)) CO2 [ppmv] CO2 [ppmv] CO2 [ppmv] (probablistic assessment of CO...) CO2 [ppmv] (probability maximum 95% lower...) CO2 [ppmv] (probability maximum 95% upper...)Parameter(s): AGE [ka BP] (Age) Carbon dioxide [ppmv] (CO2) # COMMENT: median CO2 (ppm) Carbon dioxide [ppmv] (CO2) # COMMENT: CO2 (ppm) probabilistic upper bound of 95% probability interval [sic.] Carbon dioxide [ppmv] (CO2) # COMMENT: CO2 (ppm) probabilistic upper bound of 95% probability interval Carbon dioxide [ppmv] (CO2) # COMMENT: probablistic assessment of CO2 (ppm) Carbon dioxide [ppmv] (CO2) # COMMENT: probability maximum 95% lower bound for mean (ppm) Carbon dioxide [ppmv] (CO2) # COMMENT: probability maximum 95% upper bound for mean (ppm)

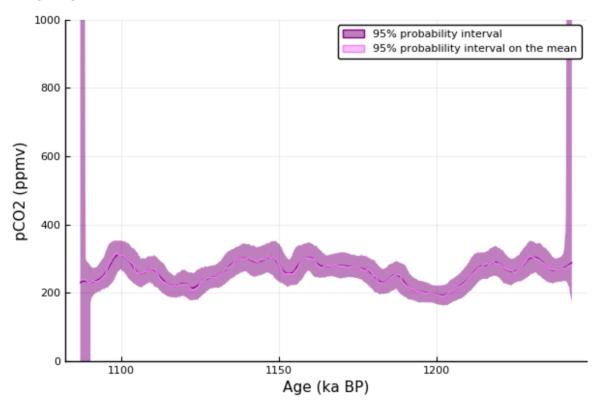
In [186]:

```
# CO2 Chalk (d11B proxy data)
data CO2 Chalk = DelimitedFiles.readdlm("../../MASTER 2.0/data/CO2/Chalk2017 CO
2.tab", Float64, skipstart = 18);
#Chalk et al dataset reports the following:
# 95% confidence interval
age C = data CO2 Chalk[:,1];
                                 # Age (ka BP)
CO2o mean C = data CO2 Chalk[:,2]; # median CO2 (ppmv)
                                # CO2 (ppm) probabilistic upper bound of 95%
CO2o lo C = data CO2 Chalk[:,3]
probability interval - [sic.] - probably mean lower bound]
CO2o_up_C = data_CO2_Chalk[:,4] # CO2 (ppm) probabilistic upper bound of 95%
probability interval
# 95% confidence interval on the mean
CO2p_mean_C = data_CO2_Chalk[:,5]; # probablistic assessment of CO2 (ppm)
CO2p lo C = data CO2 Chalk[:,6] # probability maximum 95% lower bound for mea
n (ppmv)
CO2p up C = data CO2 Chalk[:,7] # probability maximum 95% upper bound for mea
n (ppmv)
;
```

In [187]:

```
# plot Chalk
plot_Chalk =
plot(#title = "syn-MPT pCO2 record (Chalk)",
    xlabel = "Age (ka BP)",
    ylabel = "pCO2 (ppmv)",
    xticks = (:1050:50:1250),
    ylims = (0,1000)
)
plot!(age_C, CO2o_mean_C,
    ribbon = (CO2o_mean_C - CO2o_lo_C, CO2o_up_C - CO2o_mean_C),
    label = "95% probability interval",
    color = :purple)
plot!(age_C, CO2p_mean_C,
    ribbon = (CO2p_mean_C - CO2p_lo_C, CO2p_up_C - CO2p_mean_C),
    label = "95% probablility interval on the mean",
    color = :violet)
```

Out[187]:



To fully represent the uncertainty, we will use the 95% probability interval (purple ribbon in plot above).

However, there is obviously something off with the first and last datapoints reported in the Chalk dataset (a range from -10^5 ppmv to 5*10^4 ppmv is non-sensical). We therefore will not include the first and last datapoints of the record, where values deviate from the range in the remainder of the record. We have prepared a second version of the .tab-file, where we have cut out the improbable values (first 3 kyr and last 2 kyrs of observations)

```
In [40]:
```

```
\# data_CO2_Chalk[ (age_C .> 1090) .& (age_C .< 1241) ] \# Didn't work to cut this way, so we cut directly in the data file
```

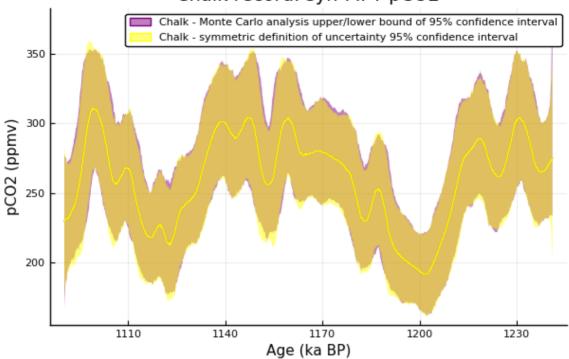
The pCO2 record we will use is plotted below:

In [188]:

```
# CO2 Chalk (d11B proxy data)
data_CO2_Chalk = DelimitedFiles.readdlm("../../MASTER_2.0/data/CO2/Chalk2017_CO2
v2.tab", Float64, skipstart = 18); # We have cut out the time interva
age C = data CO2 Chalk[:,1];
                                   # Age (ka BP)
CO2 mean C = data CO2 Chalk[:,2]; # median CO2 (ppmv)
CO2 lo C = data CO2 Chalk[:,3] # CO2 (ppm) probabilistic lower bound of 95%
probability interval
CO2 up C = data CO2 Chalk[:,4] # CO2 (ppm) probabilistic upper bound of 95%
probability interval
# defining an array for standard deviation, needed when defining our uivD
CO2 1\sigma C = (CO2 up C .- CO2 1\sigma C ) / 4 # St.dev (1\sigma) equals 1/4 of the 9
5% CI (2\sigma on each side of mean)input when defining the uivD
plot Chalk age =
plot(title = "Chalk record: syn-MPT pCO2",
    xlabel = "Age (ka BP)",
    ylabel = "pCO2 (ppmv)")
plot!(age C , CO2 mean C ,
    ribbon = (CO2_mean_C_ .- CO2_lo_C_, CO2_up_C_ .- CO2_mean_C_),
    label = "Chalk - Monte Carlo analysis upper/lower bound of 95% confidence in
terval",
    color = :purple)
plot!(age_C_, CO2_mean_C_,
    ribbon = (2*CO2_1\sigma_C_, 2*CO2_1\sigma_C_),
    label = "Chalk - symmetric definition of uncertainty 95% confidence interva
1",
    color = "yellow")
#= There are minor variations in the two ways of defining the 95% confidence in
terval.
This is due to asymmetry in upper and lower bounds
compared to a completely smooth hypothetical normal distribution.
We deem the differences to be so small that we can safely assume a normal distri
bution for the data. =#
savefig("../figurar/RawData/pCO2/Chalk rawD age.pdf")
```

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Chalk record: syn-MPT pCO2

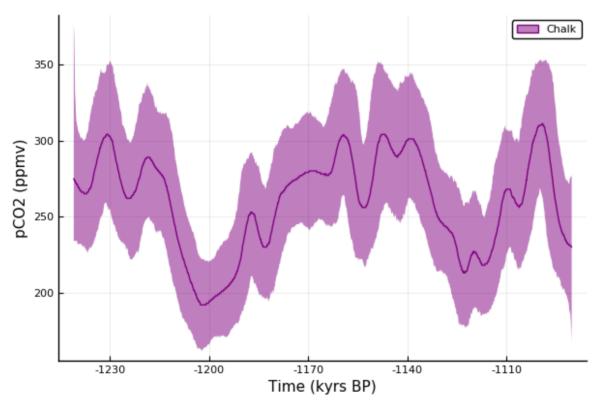


ii) Reverse arrays (from age to time)

In [189]:

```
# We reverse the arrays we will use to get the dataset of time (running forward
s)
t C = -reverse(age C , dims = 1)
CO2 mean C = reverse(CO2 mean C , dims = 1)
CO2 lo C
          = reverse(CO2_lo_C_, dims = 1)
CO2 up C
          = reverse(CO2_up_C_, dims = 1)
# defining an array for standard deviation, needed when defining our uivD
CO2 1\sigma C = (CO2 up C .- CO2 lo C) / 4 # St.dev (1\sigma) equals 1/4 of the 95% C
I (2\sigma on each side of mean) input when defining the uivD
@save "../Koding/WrangledDataFiles/BasicArrays/Chalk.jld2" t_C CO2_mean_C CO2_lo
_C CO2_up_C CO2 1σ C
# check by plotting
plot Chalk =
plot(xlabel = "Time (kyrs BP)",
    ylabel = "pCO2 (ppmv)")
plot!(t C, CO2 mean C,
    ribbon = (CO2 mean C .- CO2 lo C, CO2 up C .- CO2 mean C),
    color = :purple,
    label = "Chalk")
```

Out[189]:



iii) Age model uncertainty

The age model of the record comes from aligning the benthic foram $\delta^{18}O$ record from the same core (ODP999) to the the LR04 reference stack. [Check: it is unclear whether or not the age model uncertainty was incorporate in the Monte Carlo analysis of the by Chalk et al.] We therefore make an array add the LR04 age model uncertainty (95% confidence envelope = 6 kyrs for the Chalk time interval gives 1\sigma = 1.5 kyrs).

In [191]:

```
# age model uncertainty from LR04
t_1\sigma_C = [interpolate_t_1\sigma_LR04(i) for i in t_C];

UndefVarError: interpolate_t_1\sigma_LR04 not defined

Stacktrace:
[1] top-level scope at In[191]:1
```

Since our method operates with a lead-lag definition of causality, it is very sensitive to age reversals, and it is crucial to include uncertainties in the time series' age models.

The age model was constructed from a benthic $\delta^{18}O$ record from the same site, which has been tuned to the LR04 reference age model. We therefore add the LR04 age model uncertainty to the index uncertainty when redefining the Chalk record datatype to uivD.

The LR04 age model full uncertainty enveolpe uncertainty reported for 1-3 Ma BP is 6 kyrs (Lisiecki & Raymo, 2005). We interpreted this as the 95% confidence interval, aka 4σ . This gives us a 1σ age uncertainty value of 1.5 kyr for the entire span of the Chalk record.

In [194]:

```
# Create an array for age model uncertainty for the Chalk record

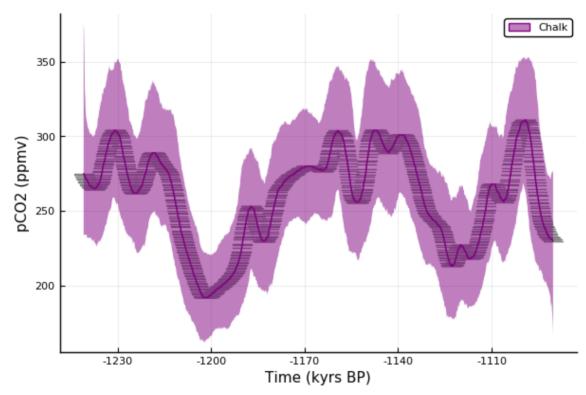
# t_1\sigma_LR04[t_C];
# Doesn't work to redefine LR04 age model uncertainty this way, probably due to
    dimension mismatch.

# We therefore create an empty array and fill it with the corresponding value fo
    r 1\sigma
    t_1\sigma_C = zeros(length(t_C))
    for i in 1:length(t_1\sigma_C)
        t_1\sigma = 1 kyr for the entire length of the Chalk time se
    res
    end
    t_1\sigma_C; # Check, all 1.5 = all good.
```

In [195]:

```
# check by plotting
plot_Chalk =
plot(xlabel = "Time (kyrs BP)",
    ylabel = "pCO2 (ppmv)")
plot!(t_C, CO2_mean_C, xerr = 2*t_1\sigma_C, color = :grey, ms = 0.1, label = "")
plot!(t_C, CO2_mean_C,
    ribbon = (CO2_mean_C - CO2_lo_C, CO2_up_C - CO2_mean_C),
    color = :purple,
    label = "Chalk")

savefig("../figurar/RawData/pCO2/plot_Chalk_rawD_timeuncLR04.pdf")
```



In [196]:

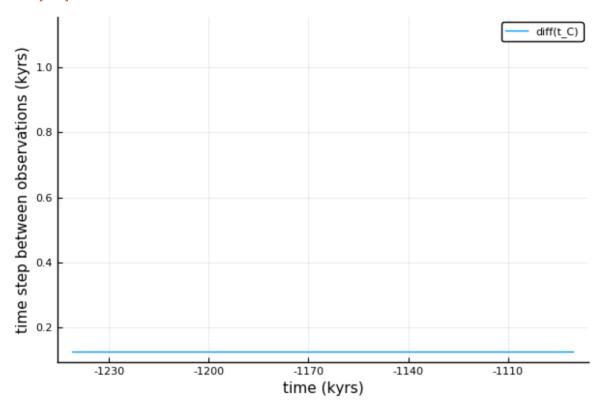
```
@save "../Koding/WrangledDataFiles/BasicArrays/Chalk.jld2" t_C t_1\sigma_C CO2_mean_C CO2_lo_C CO2_up_C CO2_1\sigma_C
```

Resolution

We now check the temporal resolution and evt need for interpolation on the Chalk dataset:

In [73]:

Out[73]:



The Chalk dataset is a high resolution record, requiring no interpolation. Observations are on a regular grid, with one observation for every 125 years, same as Rohling data, making it possible to do a more high-resolution analysis between these two time series.

Important note on resolution: These are not the raw data. Chalk et al. (2017) report original samples for every 3.5 - 4.5 years (Supporting information). The resolution they publish in the dataset is thus an order of magnitude above the sampled resolution. The over-sampling of the data may bias the results, as we use a data-driven method.

v) Redefining Chalk data as an UncertainIndexValueDataset **

```
In [21]:
```

```
# Redefining Chalk as uivD
t_uiv_C = [UncertainValue(Normal, t_C[i], t_1\sigma_C[i]) for i in 1:length(t_C)]
CO2_uiv_C = [UncertainValue(Normal, CO2_mean_C[i], CO2_1\sigma_C[i]) for i in 1: leng
th(CO2_mean_C)]
uivD_C = UncertainIndexValueDataset(t_uiv_C, CO2_uiv_C)
```

Out[21]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 1207 uncertain values coupled with 1207 uncertain ind
ices

In [23]:

```
# plot(uivD_C)
```

In [24]:

```
# Save the relevant arrays of the Chalk record in a .jld2 file @save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Chalk.jld2" uivD_C
```

vi) Binned resampling

```
In [25]:
```

```
@load "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Chalk.jld2"
Out[25]:
1-element Array{Symbol,1}:
:uivD C
```

Due to a false positive (with La2004 on 125 kyr timestep grid) we wish to use this record for sensitivity analysis. We therefore prepare many versions of the time series binned with different resolutions.

standard 1 kyr timestep grid

In [29]:

```
# make 1 kyr timestep grid
binsize = 1
tmin_C = ceil(minimum(t_C))
tmax_C = floor(maximum(t_C))
grid_C = tmin_C + binsize/2 : binsize : tmax_C - binsize/2
resampling_method_C = BinnedResampling(grid_C,1000)
```

Out[29]:

BinnedResampling{StepRangeLen{Float64, Base.TwicePrecision{Float64}, B
ase.TwicePrecision{Float64}}}(-1240.5:1.0:-1091.5, 1000)

In [30]:

```
@time C_binned = resample(uivD_C, resampling_method_C)
```

12.806340 seconds (16.77 M allocations: 1.680 GiB, 4.88% gc time)

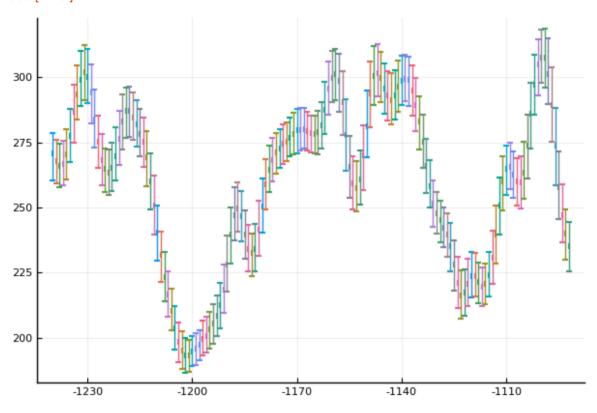
Out[30]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 149 uncertain values coupled with 149 uncertain indic
es

In [118]:

```
p1 = plot(C_binned)
#xgrid!(p1, :on, :red, 1, :dash)
```

Out[118]:



Make a high resolution version of the time series, for high resolution analyses.

• C binned hr125 with time step of 125 years, for hr analysis with La2004

In [31]:

```
binsize_hr = 0.125
grid_C_hr = tmin_C + binsize_hr/2 : binsize_hr : tmax_C - binsize_hr/2
resampling_method_C_hr = BinnedResampling(grid_C_hr, 1000)
@time C_binned_hr0125 = resample(uivD_C, resampling_method_C_hr)
```

17.373565 seconds (21.50 M allocations: 7.177 GiB, 13.13% gc time)

Out[31]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 1199 uncertain values coupled with 1199 uncertain ind
ices

• C binned hr500 with time step of 500 years, for hr analysis with Martinez-García

In [35]:

```
tmin_C = t_C[1]
tmax_C = t_C[end]
binsize_hr = 0.5
grid_C_hr500 = tmin_C + binsize_hr/2 : binsize_hr : tmax_C - binsize_hr/2
```

Out[35]:

-1240.75:0.5:-1090.75

In [34]:

```
resampling_method_C_hr500 = BinnedResampling(grid_C_hr500, 1000)
@time C_binned_hr500 = resample(uivD_C, resampling_method_C_hr500)
```

12.953482 seconds (16.60 M allocations: 2.432 GiB, 5.52% gc time)

Out[34]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 300 uncertain values coupled with 300 uncertain indic
es

vii) Save the binnedresampled timeseries

In [36]:

@save "../Koding/WrangledDataFiles/Binned_ts_fullength/Chalk.jld2" C_binned C_binned_hr0125 C_binned_hr500 $\#\ uivD_C$

In [37]:

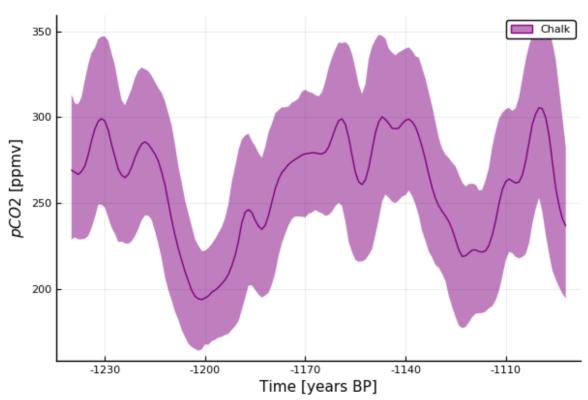
```
@load "../Koding/WrangledDataFiles/Binned_ts_fullength/Chalk.jld2"
```

```
Out[37]:
```

In [38]:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
C = C binned
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(C.values, 0.5)
bin upper = quantile.(C.values, 0.975) .- bin median
bin lower = bin median .- quantile.(C.values, 0.025)
# time array
binmidpoints C = [C.indices[i].value for i in 1:length(C)]
plot C binned =
plot(binmidpoints C, bin median,
    ribbon = (bin_lower, bin_upper),
    color = :purple,
    label = "Chalk",
    xlabel = "Time [years BP]",
    ylabel = string(L"pCO2", " [ppmv]"),
    #grid = false
```

Out[38]:



4.3 - Hönisch pCO2 record

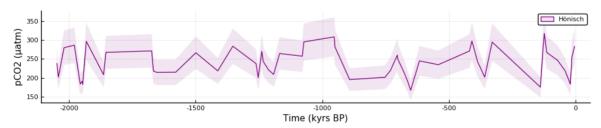
- From Hönisch et al. (2009), denoted H
- Low-resolution pCO2 record from d11B proxy, spanning ca 2 Ma.
- · Data available from ... DOI:..

Note: This is a very low resolution (only 54 datapoints over 2 Ma). Therefore contains little information on the dynamics of pCO2. It will therefore not be meaningful to run our analysis, but it is nice to have to get an idea

In [40]:

```
# CO2 Hönisch (2009)
rawD Honisch = readdlm("../../MASTER 2.0/data/CO2/Honisch2009 data v2.txt", '\t'
, Float64, '\r', dims = (54, 3))
# Note: have selected columns A, Q and R (specs below) from the pangaea file, an
d replaced empty cells with NaN.
revD H = reverse(rawD Honisch, dims = 1)
t H = - revD H[:,1] # A - redefining from age (ka BP) to time (kyrs) by making n
egative
CO2_H = revD_H[:,2]
                           # Q - pCO2 [\muatm] from calculations based on the varyi
ng alkalinity scenario (4%)
CO2 1\sigma H = revD H[:,3]
                        # R - standard deviation of pCO2
;
# plot
plot Honisch =
plot(xlabel = "Time (kyrs BP)",
    ylabel = "pCO2 (\muatm)",
    size = (1000, 200))
plot!(t H, CO2 H,
    color = :purple,
    fillalpha = 0.1,
    label = "Hönisch",
    \#yerr = (2*CO2 1\sigma H), \# vertical error bars
    ribbon = (2*CO2 1\sigma H, 2*CO2 1\sigma H), # ± 2\sigma
    markerstrokecolor = :black,
    ms = 1,
```

Out[40]:



In [41]:

savefig("../../Master 2.0/figurar/RawData/pCO2/plot Honisch ribbonscatter.pdf")

Now let's make a comparative plot of the pCO2 records

In [186]:

```
# comparative plot of pCO2 records
plot pCO2 comparative =
plot(xlabel = "Time (kyrs BP)",
    ylabel = "pCO2 (ppmv)",
    size = (1000, 400),
    legend = :topleft
# Chalk:
plot!(t C, CO2 mean C,
    ribbon = (2 * CO2_1\sigma_C), # plotting 2\sigma to illustrate that we will carry on a
95% confidence interval
    color = :violet,
    label = "Chalk")
# Bereiter:
plot!(t_B, CO2_mean_B, xerr = (2 * t_1\sigma_B), ms = 0.1, color = :black,
    label = "2\sigma age uncertainty" # how can I remove this label completely?
    )
plot!(t_B, CO2_mean_B,
    ribbon = (2 * CO2 1\sigma B), # 95% CI
    color = :red,
    fillalpha = 0.5,
    label = "Bereiter")
#savefig(".../.../Master 2.0/figurar/RawData/pCO2/plot pCO2 comparative B C ageun
c.pdf")
# Honisch:
twinx()
ylabel = "pCO2 (µatm)" # NEED HELP HERE on how to plot a second y-axis
scatter!(t_H, CO2_H,
    color = :purple,
    fillalpha = 0.1,
    label = "Hönisch",
    \#yerr = (2*CO2 1\sigma H), \# vertical error bars
    ribbon = (2*CO2 1\sigma H, 2*CO2 1\sigma H), # \pm 2\sigma
    markerstrokecolor = :purple,
    ms = 1
#savefig("../../Master 2.0/figurar/RawData/pCO2/plot pCO2 comparative wHonisch.p
df")
```

```
UndefVarError: CO2_1\u03c3_H not defined

Stacktrace:
[1] top-level scope at In[186]:30
```

In [187]:

```
# overview plot of pCO2 records
# overview plot of pCO2 records
plot pCO2 overview B C =
plot(plot Bereiter,
    plot Chalk,
    plot Honisch,
    layout = grid(3,1),
    link = :x,
    size = (1000, 400))
#savefig("../../Master 2.0/figurar/RawData/pCO2/plot pCO2 overview wHonisch.pd
f")
plot pCO2 overview B C =
plot(plot Bereiter,
   plot Chalk,
    #plot Honisch,
    layout = grid(2,1),
    link = :x,
    size = (1000, 400))
#savefig("../../Master 2.0/figurar/RawData/pCO2/plot pCO2 overview B C.pdf")
```

```
UndefVarError: plot_Bereiter not defined
Stacktrace:
[1] top-level scope at In[187]:1
```

5 - Dust records

- Dust concentration record from EDC ice core (Lambert et al., 2008), spanning the last 800 kyr.
- Record of Fe mass accumulation rate in the Southern Ocean (*Martinez-Garcia et al.,2011*), spanning the last 4.3 Myr.

5.1 - Lambert dust record

- Ice core record of dust (flux?) spanning the last 800 kyrs, from Epica Dome C (East Antarctic ice sheet).
- From Lambert et al. (2008), denoted L.
- Dust concentration record from EDC ice core (East Antarctica), spanning the last 800 kyr
- data available from Pangaea, DOI: https://doi.org/10.1594/PANGAEA.695995).
- Note: we are using the lpc-data (laser scatter method), since this has higher resolution downcore.

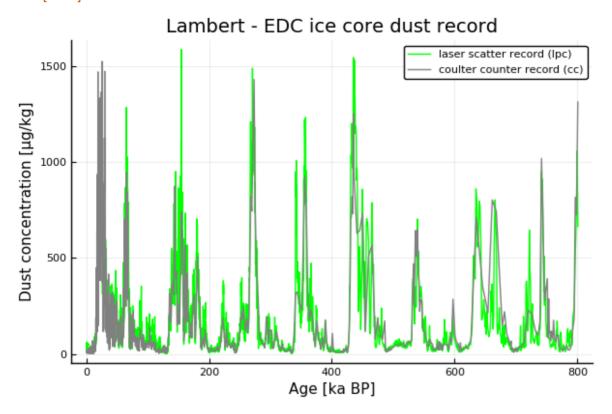
i) Load in data

Loading in data from Lambert, and checking which dataset is better to use - cc (coulter counter) or lpc (laser scatter)?

In [179]:

```
# Loading in data from Lambert
# and checking which dataset is better to use - cc (coulter counter) or lpc (las
er scatter)?
# load in Lambertdataset 1: EDC dust cc.tab
# From Pangaea data repository: Lambert et al (2008), file name EDC dust cc.tab
EDC dust cc = "../../MASTER 2.0/data/dust/Lambert(2008) dust EDC/datasets/EDC du
st cc.tab"
rawD L cc = readdlm(EDC dust cc, '\t', Float64, dims = (1154, 3), skipstart = 16
) #reading in the raw dataset, skipping the first 16 rows (descriptive)
# renaming columns of interest
depth L cc = rawD L cc[:,1] # depth (m)
age L cc = rawD L cc[:,2]; # age (ka) from Lambert dataset (EDC3 age model)
dust L cc = rawD L cc[:,3]; # dust concentration(µg/kg)
##### load in Lambert dataset 2: EDC dust lpc
# From Pangaea data repository: Lambert et al (2008), file name EDC dust lpc.tab
EDC dust lpc = "../../MASTER 2.0/data/dust/Lambert(2008)_dust_EDC/datasets/EDC_d
ust lpc.tab"
rawD L lpc = readdlm(EDC dust lpc, '\t', Float64, dims = (5163, 3), skipstart =
15) #reading in the raw dataset, skipping the first 16 rows (descriptive)
# renaming columns of interest
depth L lpc = rawD L lpc[:,1] # depth (m)
age_L_lpc = rawD_L_lpc[:,2]; # age (ka) (from EDC3 age model)
dust L lpc = rawD L lpc[:,3]; # dust concentration (µg/kg)
# Which data to use, lpc or cc?
length(dust_L_cc) # 1154
length(dust L lpc) # 5163
# lpc has more observations
# looks like cc is concentrated in the upper end of the core
# based on visuals, seems that lpc has better resolution down-core
plot(xlabel = "Age [ka BP]", ylabel = "Dust concentration [\mug/kg]", title = "Lam
bert - EDC ice core dust record")
plot!(age L lpc, dust L lpc, color = :lime, label = "laser scatter record (lp
plot!(age L cc, dust L cc, color = :grey, label = "coulter counter record (cc)")
# We decide to use the lpc-data (laser sensor),
# since this is the dataset with highest resolution back in time.
```

Out[179]:



Dust concentration vs dust flux.

"Because of the low accumulation rate at Dome C (ca 3cm/yr water equivalent), dry deposition is dominant and **the atmospheric dust load is best represented by the dust flux**" (Lambert et al. (2008), referring to the work of Wolff et al (2006)). "(...)the flux is the better analogue for atmospheric concentration at Dome C, although its use could induce an error in the change in concentration between glacial and interglacial of 10-20%." (Wolff et al 2006, in supplementary discussion)

Unfortunately, Lambert et al. has **only published data as dust concentration, and not included it's conversion to dust flux**. They cite Wolff et al (2006). for how they have done the conversion from concentration to flux. Wolff et al (2006) report the following on how to to calculate the flux:

"flux, J, is given by:

$$J = vd \bullet C_{air} + K \bullet P \bullet C_{air}$$

where C_{air} is the concentration of the chemical in air, P is the snow precipitation rate, vd and K are constants of proportionality (vd is known as the dry deposition velocity, and K is related to a mass-based scavenging ratio). The first term is the dry deposited flux and the second term is the wet deposited flux.

The concentration in ice, C_{ice} is simply the flux divided by the average precipitation rate for the period represented by the ice sample."

Since we do not have access to the parameters Lambert et al. have used in the conversion (most importantly the precipitation rate), we check by comparing the Lambert dust concentration record to the Wolff Fe flux record.

Or, Lambert et al report a mean precipitation rate for Dome C to be ca 3 cm/yr. However, reduced to a constant, this will not make any change to the dynamics of the time series, will it? All it would change is we could add the 10-20 \% uncertainty proposed [but was only proposed between glacial and interglacial] (supplementary discussion by Wolff et al.,2006)

In [180]:

```
fn = "/Users/maria/Downloads/41586_2006_BFnature04614_MOESM3_ESM_missingscut.tx
t"
data = readdlm(fn, '\t', dims = (272,5), Any, '\r', skipstart = 6)
Age = data[:,1]
Feflux = data[:,3]
Feflux[Feflux .== ""] .= NaN
```

Out[180]:

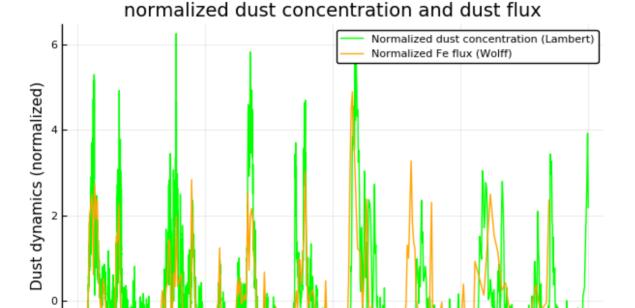
```
0-element view(::Array{Any,1}, Int64[]) with eltype Any
```

In [181]:

```
# normalization to get comparable scale for data
norm_Feflux = ((Feflux .- mean(Feflux)) / std(Feflux))
norm_dust_L = ((dust_L_lpc .- mean(dust_L_lpc)) / std(dust_L_lpc))

# Plot concentration and flux to compare
plot(title = "normalized dust concentration and dust flux", xlabel = "Time (kyrs BP", ylabel = "Dust dynamics (normalized)")
plot!(age_L_lpc, norm_dust_L, color = :lime, label = "Normalized dust concentrat ion (Lambert)")
plot!(Age, norm_Feflux, color = :orange, label = "Normalized Fe flux (Wolff)")
```

Out[181]:



Lambert dust concentration data seems more "spikey" than the Wolff Fe flux data.

200

Set to a logarithmic scale

(why? if the data was normalized already, this was a dynamics thing, not relict of units)

400

Time (kyrs BP

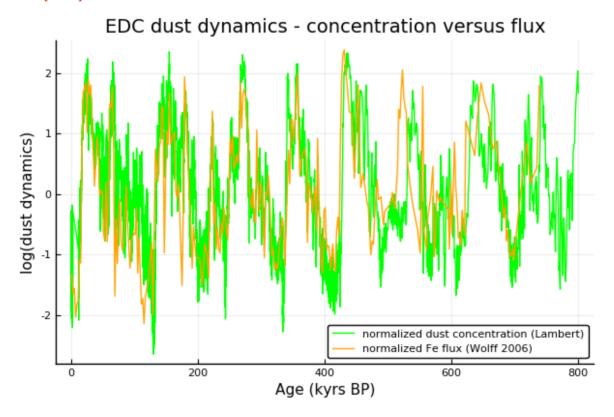
600

800

In [182]:

```
# NB can only run this cell once
# normalize Fe flux & put on a logarithmic scale
index = Feflux .!== NaN
norm Fe = Feflux
norm Fe[index] = log.(norm Fe[index]) # log causes a -Inf
ix = (norm Fe .!== -Inf) .& (norm Fe .!== NaN)
norm Fe[ix] = (norm Fe[ix] .- mean(norm Fe[ix])) ./ std(norm Fe[ix])
# normalize dust concentration & put on a logarithmic scale
index = dust L lpc .!== NaN
norm_dustconc = dust_L_lpc
norm_dustconc[index] = log.(norm_dustconc[index]) # log causes a -Inf
ix = (norm dustconc .!== -Inf) .& (norm dustconc .!== NaN)
norm dustconc[ix] = (norm dustconc[ix] .- mean(norm dustconc[ix])) ./ std(norm d
ustconc[ix])
plot(xlabel = "Age (kyrs BP)", ylabel = "log(dust dynamics)", title = "EDC dust
dynamics - concentration versus flux")
plot!(age L lpc, norm dustconc, label = "normalized dust concentration (Lamber
t)", color = :lime)
plot!(Age, norm Fe, label = "normalized Fe flux (Wolff 2006)", color = :orange )
```

Out[182]:



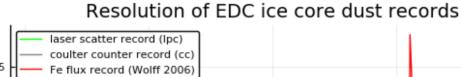
Dust concentration and flux are in fairly good agreement. There seem to be some differences in the age model, but this may be due to the Lambert record being on the EDC3 age model, while the Wollf data is on the EPC2 age model.

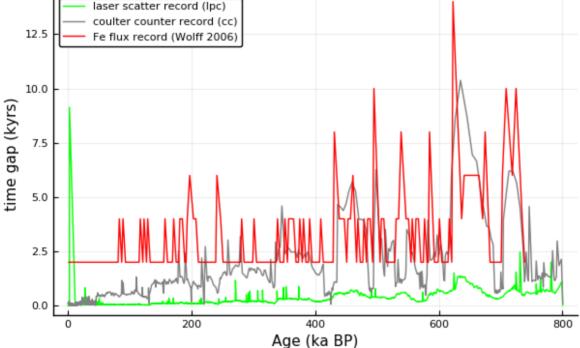
Check record resolutions

In [166]:

```
plot(xlabel = "Age (ka BP)", ylabel = "time gap (kyrs)", title = "Resolution of
EDC ice core dust records")
plot!(age L lpc, diff(age L lpc), color = :lime,
                                                  label = "laser scatter record
 (lpc)")
plot!(age L cc, diff(age L cc), color = :grey, label = "coulter counter record
 (CC)")
plot!(Age, diff(Age), color = :red, label = "Fe flux record (Wolff 2006)") # low
resolution, not suitable for our method.
```

Out[166]:





The Wolff (2006) Fe flux data has an order of magnitude lower resolution, which is what is of greater importance to our method (dynamical information). We therefore, in spite of not being as much of a direct process proxy, choose to use the Lambert dust concentration record further in our analyses.

We decide to use the lpc-data over the cc-data, since this is the dataset with highest resolution back in time.

Cutting away the low resolution interval at the youngest part of the record

There is a large time gap in the most recent part of the record, shown in the plot above. But where exactly is the gap? If we run the cell below (commented out for visual reasons) we can see that the observational gap is between 2.424 and 11.554 kyrs BP.

```
In [195]:
```

```
show(age L lpc)
```

Now, we would like to cut the time series to only keep the high resolution part of the record.

In order to cut the time series, we need to find the corresponding index of where we want to cut. Therefore, we make an array of the observations up until that age gap point.

In [196]:

```
beforetimegap = [0.371, 0.383, 0.442, 0.454, 0.466, 0.478, 0.49, 0.503, 0.516,
0.529, 0.542, 0.555, 0.568, 0.581, 0.594, 0.606, 0.619, 0.632, 0.672, 0.701, 0.7
15, 0.729, 0.744, 0.759, 0.772, 0.786, 0.799, 0.825, 0.837, 0.862, 0.875, 0.889,
0.903, 0.917, 0.932, 0.947, 0.961, 0.976, 0.99, 1.005, 1.019, 1.032, 1.047, 1.06
1, 1.075, 1.089, 1.104, 1.118, 1.133, 1.147, 1.161, 1.174, 1.188, 1.201, 1.215,
1.229, 1.244, 1.258, 1.273, 1.287, 1.302, 1.316, 1.33, 1.345, 1.359, 1.373, 1.38
7, 1.401, 1.416, 1.432, 1.448, 1.463, 1.479, 1.494, 1.51, 1.525, 1.541, 1.556,
1.572, 1.588, 1.604, 1.621, 1.637, 1.654, 1.67, 1.686, 1.702, 1.737, 1.754, 1.77
1, 1.787, 1.804, 1.822, 1.839, 1.856, 1.873, 1.889, 1.905, 1.922, 1.938, 1.956,
1.973, 1.99, 2.007, 2.025, 2.044, 2.061, 2.078, 2.095, 2.113, 2.13, 2.147, 2.165
, 2.184, 2.202, 2.219, 2.237, 2.253, 2.271, 2.289, 2.306, 2.327, 2.347, 2.367,
2.405, 2.424] # copied from running the cell above.
length(beforetimegap) # 126
# checking the age value of the index before and after time gap
age L lpc[126] # 2.424
age L lpc[127] # 11.554 # high resolution array starts at index 127
# From the plot above, the time gap should be around 9 kyrs long.
age L lpc[127] - age L lpc[126]
# So this looks right
```

Out[196]:

9.13

We have found the large gap in resolution, it's between index 126 and 127. We therefore cut the first 126 observations out of the time series.

In [197]:

```
# cutting the first 126 elements of the arrays, and rename the new, shorter arra
ys
# age array
print("The original array is ", length(age L lpc), " elements long") # long vers
ion is 5163 elements
                                                                      # age (ka)
age_L_EDC3 = age_L_lpc[(age_L_lpc .> age_L_lpc[126])]
(from EDC3 age model)
print("
The cut array is ", length(age L EDC3), " elements long")
# depth array
depth_L = depth_L_lpc[(age_L_lpc .> age_L_lpc[126])];
# checking that we cut right
depth L lpc[126] # 99 m
depth L lpc[127] # 358 m
# large gap, all good
# dust array
dust L = dust L lpc[(age L lpc .> age L lpc[126])]
# checking that we cut right
length(dust L lpc) # 5163
length(dust L) # 5037
#same as above, all good
```

The original array is 5163 elements long The cut array is 5037 elements long

Out[197]:

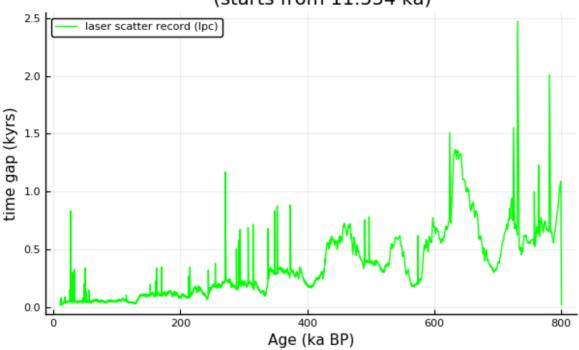
5037

In [198]:

```
plot(xlabel = "Age (ka BP)", ylabel = "time gap (kyrs)", title = "High resolutio
n part of Lambert lpc record
        (starts from 11.554 ka)")
plot!(age_L_EDC3, diff(age_L_EDC3), color = :lime, label = "laser scatter record (lpc)")
```

Out[198]:





Adjustments to age model

Note: We have two datasets (dust and CO2) from the same ice core (Epica Dome C). However, different age models are used on the two datasets. The dust record from Lambert et al. (2008) uses the older EPC3 age model, while the CO2 record from Bereiter et al. (2015) uses the revised AICC2012 age model. In the following, we want to get both datasets on the same age model (AICC2012).

In [199]:

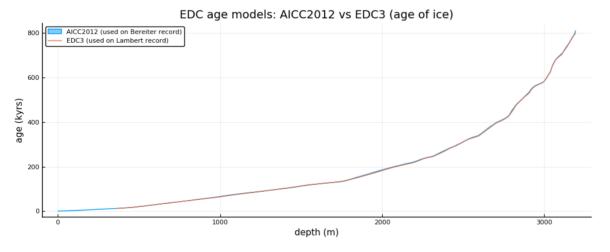
```
# read in data for the AICC2012 age model
filepath AICC2012 = "../../MASTER_2.0/Koding/age_models/EDC_AICC2012_chron.tab"
rawD AICC2012 = readdlm(filepath AICC2012, skipstart = 23, dims = (1936,8))
#columns
depth
          = Array{Float64,1}(rawD AICC2012[:,1]) # Depth ice/snow [m]
         = Array{Float64,1}(rawD_AICC2012[:,2]) # Age [ka BP]
age ice
age_ice_1\sigma = Array{Float64,1}(rawD_AICC2012[:,3]) # Age std dev [±]
        = Array{Float64,1}(rawD AICC2012[:,4])  # Gas age [ka BP]
age gas 1\sigma = Array\{Float64,1\}(rawD AICC2012[:,5]) # Gas age std dev [±]
# acc rate = rawD AICC2012[:,6]
                                 # Acc rate ice per year [m/a]
                                   # Thinning functionTF
# tf = rawD AICC2012[:,7]
# lidie
             = Array{Float64,1}(rawD AICC2012[:,8])
                                                       # Lock-in depth in ice
equivalent (LIDIE) [m] #
```

In [200]:

```
# Show difference between AICC2012 and EDC3 age models

plot_agemodels =
plot(title = "EDC age models: AICC2012 vs EDC3 (age of ice)",
    size = (1000, 400),
# xlims = (2500,2510),
    xlabel = "depth (m)",
    ylabel = "age (kyrs)")
!(depth, age_ice,
    label = "AICC2012 (used on Bereiter record)",
    ribbon = (age_ice_lo, age_ice_lo),
    fillalpha = 0.5)
plot!(depth_L, age_L_EDC3,
    label = "EDC3 (used on Lambert record)")
#plot!(depth, lidie)
```

Out[200]:



In []:

The differences between the two age models are very subtle, barely noticeable in the plot above. However, we make a point out of having the Lambert and Bereiter records on the same age model, to minimize/eliminate the age uncertainty between these records. Another advantage of the newer AICC2012 age model is that it includes a quantification of the age uncertainty. The quality of age models are of particular importance for our approach, as we operate with a lead-lag-based definition of causality.

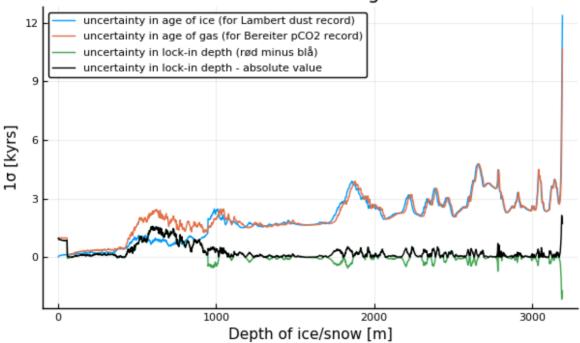
Age uncertainty from lock-in depth of gas in ice: For analyses between Lambert and Bereiter, both from the EDC core, we shall only use the age uncertainty from lock-in time of the gas bubbles in the ice. This is not reported directly in the dataset with the age model, but we can derive an array representing this uncertainty.

In [214]:

```
# age uncertainty between age of gas and age of ice
edc gasice ageunc 1\sigma = age gas 1\sigma - age ice 1\sigma
# for some reason, which uncertainty is larger varies. we therefore take the ab
solute value
edc gasice ageunc 1\sigma absolutevalue = [abs(i) for i in edc gasice ageunc 1\sigma]
#plot age uncertainties
plot(xlabel = "Depth of ice/snow [m]",
    ylabel = 10 \text{ [kyrs]},
    title = "Age uncertainties from lock-in depth
    for the AICC2012 age model")
plot!(depth, age ice 10, label = "uncertainty in age of ice (for Lambert dust re
cord)")
plot!(depth, age gas 1\sigma, label = "uncertainty in age of gas (for Bereiter pCO2 r
ecord)")
plot!(depth, edc gasice ageunc 10, label = "uncertainty in lock-in depth (rød mi
nus blå)", ms = :dot)
plot!(depth, edc gasice ageunc 10 absolutevalue, label = "uncertainty in lock-in
depth - absolute value", c = :black)
```

Out[214]:

Age uncertainties from lock-in depth for the AICC2012 age model

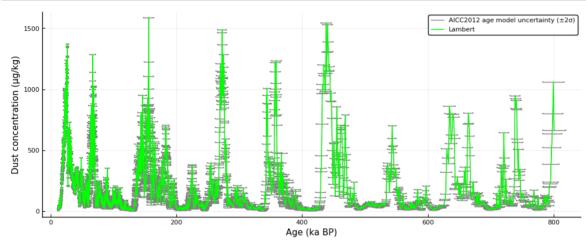


In [217]:

```
# Interpolation of data to the AICC2012 age model
# (reformulates from array of discrete ages for discrete depths, to ages given a
s a continuous function of any depth)
interpolate ageice
                      = LinearInterpolation(age ice, depth) # age of ice as
 a function of depth
interpolate ageice 1\sigma = LinearInterpolation(age ice 1\sigma, depth) # with associate
d uncertainties (1\sigma)
interpolate agegas
                      = LinearInterpolation(age gas, depth) # age of gas as
 a function of depth
interpolate agegas 1\sigma = LinearInterpolation(age gas 1\sigma, depth) # with associate
d uncertainties (1\sigma)
interpolate icegas ageunc 1\sigma = LinearInterpolation(edc gasice ageunc 1\sigma absolute
value, depth) # age uncertainty between gas and ice (1σ)
# Give corresponding age values from interp AICC2012 to the depth of observation
s in the Lambert dataset
newages L
            = [interpolate ageice(i) for i in depth L]
                                                            # array of ages (ka)
 for the Lambert dust data, according to the AICC2012 age model for ice.
newages 1\sigma L = [interpolate ageice <math>1\sigma(i) for i in depth L] # associated age unce
rtainties
newages 1\sigma L edc = [interpolate icegas ageunc 1\sigma(i) for i in depth L] # 1\sigma age un
certainty between gas and ice - to be used for analysis between Bereiter and Lam
bert records (both from edc)
;
```

In [218]:

```
# plot
#plot_Lambert_newagemodel =
plot(#title = "EDC dust concentration on AICC2012 age model",
    size = (1000, 400),
    xlabel = "Age (ka BP)",
    ylabel = "Dust concentration (µg/kg)")
plot!(newages_L, dust_L,
    xerr = 2 * newages_1\sigma_L, # 2\sigma = 95\% confidence interval on age
    ms = 1,
    color = :grey,
    label = "AICC2012 age model uncertainty (±2\sigma)")
plot!(newages_L, dust_L,
    color = :lime,
    label = "Lambert")
#savefig("../figurar/RawData/Dust/Lambert_aicc2012_wAgeUnc.pdf")
```

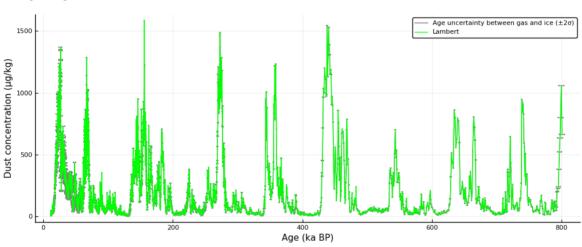


We note that the age uncertainty is quite large in this record. What effect may this have on the high resoultion analysis?

In [219]:

```
# plot
#plot_Lambert_newagemodel =
plot(#title = "EDC dust concentration on AICC2012 age model",
    size = (1000, 400),
    xlabel = "Age (ka BP)",
    ylabel = "Dust concentration (µg/kg)")
plot!(newages_L, dust_L,
    xerr = 2 * newages_lo_L_edc, # ±2\sigma = 95\% confidence interval
    ms = 1,
    color = :grey,
    label = "Age uncertainty between gas and ice (±2\sigma)")
plot!(newages_L, dust_L,
    color = :lime,
    label = "Lambert")
savefig("../figurar/RawData/Dust/Lambert_aicc2012_wAgeUnc_between_gasice.pdf")
```

Out[219]:



ii) Reverse dataset

Redefine from age (increasing backwards) to time (increasing forwards)

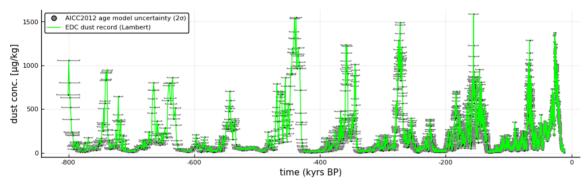
In [221]:

```
# redefine Lambert dataset from age (increasing backwards with indexation) to ti
me (increasing forwards with indexation)
t_L = - reverse(newages_L) # time is negative because we are defining present
as 0
t_lo_L = reverse(newages_lo_L)
t_lo_L_edc = reverse(newages_lo_L_edc)
dust_L = reverse(dust_L)
;
@save "../Koding/WrangledDataFiles/BasicArrays/Lambert.jld2" t_L t_lo_L t_lo_L_e
dc dust_L
```

DERIVERE FOR Å FÅ FLUX? HJELP

In [131]:

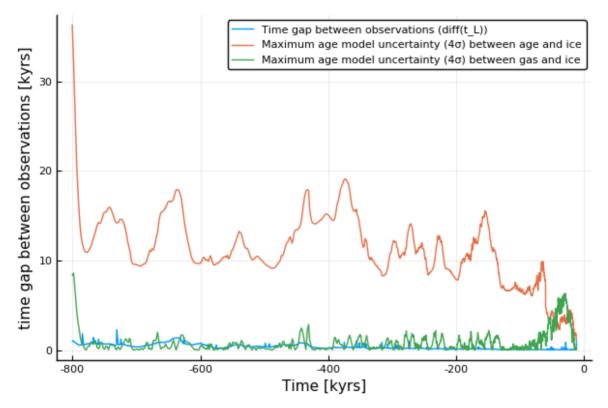
```
# plot Lambert time series
@load "../Koding/WrangledDataFiles/BasicArrays/Lambert.jld2"
plot Lambert time =
plot(#title = "EDC aeolian dust record (Lambert lpc record)",
    size = (1000, 300),
    xlabel = "time (kyrs BP)",
    ylabel = "dust conc. [\mu g/kg]")
scatter!(t L, dust L,
    xerr = 2 * t 1\sigma L, # AICC2012 age uncertainty (2\sigma)
    ms = 1,
    color = :grey,
    label = "AICC2012 age model uncertainty (2\sigma)"
plot!(t_L, dust_L,
    color = :lime,
    label = "EDC dust record (Lambert)"
    #label = "Lambert"
    )
savefig("../figurar/RawData/Dust/Lambert aicc2012 wTimeUnc.pdf")
```



iv) Interpolation of data, to ensure we have data in all bins.

In [143]:

Out[143]:



In [136]:

```
maximum(diff(t_L)) # 2.3 kyrs # (9.3 kyrs if not cutting the youngest part of t
he record)
minimum(diff(t_L)) # 20 years
mean(diff(t_L)) # 157 years # (155 years if not cutting the youngest part of
the record)
# Overall high resolution, but yes, we need interpolation for some bins
```

Out[136]:

0.1566320492454329

Resolution of the Lambert data is on average one datapoint every 150-200 years, but there are gaps of over 2000 years between observations. This means we need to interpolate.

No need to interpolate, as long as age uncertainty is larger than the time gaps - values will then still be drawn for each bin through resampling. We might still need to interpolate for the version that is to be run analysis with EDC pCO2 (reduced age uncertainty). Let's check.

In [142]:

```
##### redefining the data as an UncertainIndexValueDataset
t uiv L = [UncertainValue(Normal, t L[i], t 1\sigma L[i]) for i in 1:length(t L)] # a
ge uncertainty between gas and ice
dust uiv L = [UncertainValue(Normal, dust L[i], 0) for i in 1:length(dust L)]
uivD L noIntp = UncertainIndexValueDataset(t uiv L, dust uiv L)
# uivD L EDC is to be analysed with another EDC core, and therefore we don't inc
lude the age model uncertainties
t uiv L EDC = [UncertainValue(Normal, t_L[i], t_1\sigma_L_edc[i]) for i in 1:length(t
L)] # age uncertainty between gas and ice
dust uiv L = [UncertainValue(Normal, dust L[i], 0) for i in 1:length(dust L)]
uivD L EDC noIntp = UncertainIndexValueDataset(t uiv L EDC, dust uiv L)
##### Defining the grid for binned resampling
binsize = 1 # each timestep is 1000 years
tmin L = ceil(minimum(t L)) # first bin midpoint
tmax L = floor(maximum(t L)) # last bin midpoint
grid L = tmin L - binsize/2 : binsize : tmax L + binsize/2 # these must be the b
in edges if the bin midpoints shall be at every whole 1 kyr
# resampling on the grid
resampling method L = BinnedResampling(grid L, 1000) # resample 1000 draws (with
substitution) in each bin of the grid
@time L binned full noIntp = resample(uivD L noIntp, resampling method L)
#@time L binned full noIntp edc = resample(uivD L EDC noIntp, resampling method
L)
```

70.109201 seconds (58.62 M allocations: 19.675 GiB, 15.16% gc time)

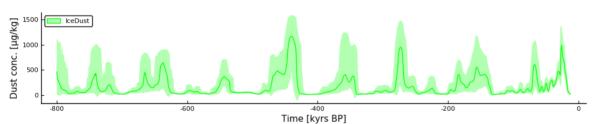
Out[142]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 789 uncertain values coupled with 789 uncertain indic
es

In [152]:

```
####### plot the binned resampled Lambert dust time series with reduced age un
certainty
ts = L binned full noIntp
binmidpoints ts = [ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot L edc =
plot(binmidpoints ts, bin median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :lime,
    label = "IceDust",
    xlabel = "Time [kyrs BP]",
    ylabel = "Dust conc. [\mu g/kg]",
    grid = false,
    size = (1000, 200),
    legend = :topleft
    )
```

Out[152]:



• check if the binned resampling could draw a value in every bin when Lambert record was defined with the smaller gas-ice age uncertainty

In [149]:

```
show(L_binned_full_noIntp_edc)
```

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 789 uncertain values coupled with 789 uncertain indic
es

In [150]:

```
####### plot the binned resampled Lambert dust time series with reduced age un
certainty
ts = L binned full noIntp edc
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot L edc =
plot(binmidpoints ts, bin median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :olive,
    label = "IceDust - age uncertainty between gas and ice",
    xlabel = "Time [kyrs BP]",
    ylabel = "Dust conc. [\mu q/kq]",
    grid = false,
    size = (1000, 200),
    legend = :left
# ...contains empty cells..., we need to interpolate..
```

ArgumentError: invalid index: nothing of type Nothing

```
Stacktrace:
```

```
[1] to index(::Nothing) at ./indices.jl:270
 [2] to index(::StepRangeLen{Float64, Base. TwicePrecision{Float64}, Ba
se.TwicePrecision{Float64}}, ::Nothing) at ./indices.jl:247
 [3] to indices at ./indices.jl:298 [inlined]
 [4] to_indices at ./indices.jl:295 [inlined]
 [5] getindex(::StepRangeLen{Float64, Base. TwicePrecision{Float64}, Ba
se.TwicePrecision{Float64}}, ::Nothing) at ./abstractarray.jl:981
 [6] quantile(::UncertainScalarKDE{Float64}, ::Float64) at /Users/ma
ria/.julia/packages/UncertainData/PbltS/src/uncertain values/Uncerta
inScalarsKDE.jl:113
 [7] broadcast getindex evalf at ./broadcast.jl:625 [inlined]
 [8] broadcast getindex at ./broadcast.jl:598 [inlined]
 [9] getindex at ./broadcast.jl:558 [inlined]
 [10] copyto nonleaf!(::Array{Float64,1}, ::Base.Broadcast.Broadcast
ed{Base.Broadcast.DefaultArrayStyle{1}, Tuple{Base.OneTo{Int64}}, type
of(quantile), Tuple{Base.Broadcast.Extruded{Array{Any,1},Tuple{Bool},
Tuple{Int64}},Float64}},::Base.OneTo{Int64},::Int64,::Int64) at
./broadcast.jl:982
 [11] copy at ./broadcast.jl:836 [inlined]
 [12] materialize(::Base.Broadcast.Broadcasted(Base.Broadcast.Defaul
tArrayStyle{1}, Nothing, typeof(quantile), Tuple{Array{Any,1},Float6
4}}) at ./broadcast.jl:798
 [13] top-level scope at In[150]:4
```

The edc version contains empty bins. Since we need continuous time series for our method, we need to interpolate some values if we are to use this record.

In [84]:

```
# Create continuous functions that linearly interpolates between each data point s
interpolate_t_L = LinearInterpolation(t_L, t_L)  # function to inte
rpolate age array
interpolate_t_lo_L = LinearInterpolation(t_lo_L, t_L)  # function to inte
rpolate age model uncertainties
interpolate_t_lo_L_edc = LinearInterpolation(t_lo_L_edc, t_L) # function to inte
rpolate age uncertainties between gas and ice
interpolate_dust_L = LinearInterpolation(dust_L, t_L); # function to inte
rpolate dust concentration
```

The otherwise high resolution of the data (one observation every 150 years on average) makes it defendable to draw interpolated values on a high resolution/centennial resolution grid without creating many false values. This will allow us to run higher resolution analyses on this dataset.

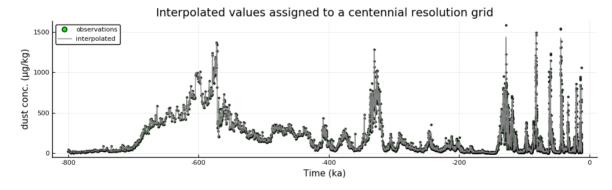
WHAT RESOLUTION TO ASSIGN ON?

In [99]:

```
# make a fine-grained time-grid to which we assign interpolated values for the L
ambert dataset.
fine grid L = minimum(t L) : 0.1 : maximum(t L) # 1 bin for every 100 years
# we then put the interpolated values in a new object intpD L, containing one va
lue in every bin of the fine time grid
intpD t L = [interpolate t L(i) for i in fine grid L];
intpD t 1\sigma L = [interpolate t 1\sigma L(i) for i in fine grid L];
intpD t 1\sigma L edc = [interpolate_t_1\sigma_L_edc(i) for i in fine_grid_L];
intpD dust L = [interpolate dust L(i) for i in fine grid L];
print(length(intpD t L))
# let's plot to check if interpolation is ok
plot(title = "Interpolated values assigned to a centennial resolution grid",
    xlabel = "Time (ka)",
    ylabel = "dust conc. (\mu g/kg)",
    size = (1000, 300)
scatter!(t L, dust L,
    \#xerr = 2 * t 1\sigma L
    ms = 2,
    color = :lime,
    label = "observations")
plot!(intpD t L, intpD dust L,
                           # too computationally heavy and too cluttering to pl
    \#xerr = intpD t 2\sigma B,
ot xerr for every 100 years
    color = :grey,
    label = "interpolated")
```

7888

Out[99]:



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Some observed peak values are not captured by the interpolation when we assign the values on a centennial grid.

To minimize the loss of nuance in data, we may choose a higher resolution fine grid for assigning the interpolated values.

wouldn't that also require higher n_draws when moving on to BinnedResampling?

I think: finer grid = less loss of nuance, but also TRADE-OFF with creating a fake smooth data where we have larger time gaps.

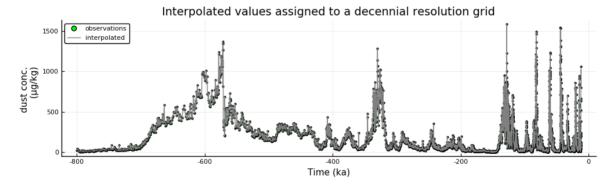
DISCUSS with Bjarte and Jo the best way to go about this.

In [100]:

```
# make a fine-grained time-grid to contain interpolated values for the Lambert d
ataset: Decennial resolution
fine grid L = minimum(t L) : 0.01 : maximum(t L) # 1 bin for every 10 years
# we then put the interpolated values in a new object intpD L, containing one va
lue in every bin of the fine time grid
t L intpDdec = [interpolate t L(i) for i in fine grid L];
t 1σ L intpDdec = [interpolate t 1σ L(i) for i in fine grid L];
dust L intpDdec = [interpolate dust L(i) for i in fine grid L];
print(length(t L intpDdec))
# let's plot to check if interpolation is ok
plot(title = "Interpolated values assigned to a decennial resolution grid",
    xlabel = "Time (ka)",
    ylabel = "dust conc.
    (\mu g/kg)",
    size = (1000,300)
scatter!(t L, dust L,
    \#xerr = 2 * t 1\sigma L
    ms = 2,
    color = :lime,
    label = "observations")
plot!(t L intpDdec, dust L intpDdec,
    \#xerr = intpD \ t \ 2\sigma \ B, \# \ too \ computationally \ heavy \ and \ too \ cluttering \ to \ pl
ot xerr for every 100 years
    ms = 0.1,
    color = :grey,
    label = "interpolated")
```

78880

Out[100]:



Redefining the time series as an UncertainIndexValueDataset

In [101]:

```
# redefining the interpolated dataset `intpD_L` (centennial resolution) as an u ivD

t_uiv_L = [UncertainValue(Normal, intpD_t_L[i], intpD_t_lo_L[i]) for i in 1:leng th(intpD_t_L)] # age uncertainty from AICC2012 age model dust_uiv_L = [UncertainValue(Normal, intpD_dust_L[i], 0) for i in 1:length(intpD_dust_L)] # no uncertainties reported concerning measurements uivD_L = UncertainIndexValueDataset(t_uiv_L, dust_uiv_L)

# We now have our uncertain index value Dataset for the dust record from Lambert et al.(2008)
```

Out[101]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDataset} containing 7888 uncertain values coupled with 7888 uncertain indices

Given that the Bereiter (pCO2) and Lambert (dust) records are both from the same ice core and now on the same age model, we can ignore the age model uncertainty between them and analyse directly on depth. We will use only the smaller uncertainties stemming from the lock-in time of the gas bubbles in the ice. We therefore also prepare a uivD version of the Lambert record with this small age uncertainty, for analysis between that specific time series pair.

In [104]:

```
# uivD_L_EDC is to be analysed with another EDC core, and therefore we don't inc
lude the age model uncertainties

t_uiv_L_EDC = [UncertainValue(Normal, intpD_t_L[i], intpD_t_lo_L_edc[i]) for i i
n 1:length(intpD_t_L)] # age uncertainty between gas and ice
dust_uiv_L = [UncertainValue(Normal, intpD_dust_L[i], 0) for i in 1:length(intpD_dust_L)]
uivD_L_EDC = UncertainIndexValueDataset(t_uiv_L_EDC, dust_uiv_L)
```

Out[104]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 7888 uncertain values coupled with 7888 uncertain ind
ices

In [105]:

```
plot(uivD_L_EDC, xlabel = "UncertainIndex (Time in kyrs BP)", ylabel = "UncertainValue (dust conc. in \mu g/kg)")
```

InterruptException:

```
Stacktrace:
```

- [1] get_clims(::Plots.Subplot{Plots.PyPlotBackend}) at /Users/mari a/.julia/packages/Plots/qZHsp/src/utils.jl:556
- [2] get_clims(::Plots.Subplot{Plots.PyPlotBackend}, ::Plots.Series) at /Users/maria/.julia/packages/Plots/qZHsp/src/utils.jl:569
- [3] py_add_series(::Plots.Plot{Plots.PyPlotBackend}, ::Plots.Serie s) at /Users/maria/.julia/packages/Plots/qZHsp/src/backends/pyplot.j 1:402
- [4] _before_layout_calcs(::Plots.Plot{Plots.PyPlotBackend}) at /Use rs/maria/.julia/packages/Plots/qZHsp/src/backends/pyplot.jl:975
- [5] prepare_output(::Plots.Plot{Plots.PyPlotBackend}) at /Users/mar ia/.julia/packages/Plots/qZHsp/src/plot.jl:254
- [6] show(::Base64.Base64EncodePipe, ::MIME{Symbol("image/png")}, ::
 Plots.Plot{Plots.PyPlotBackend}) at /Users/maria/.julia/packages/Plo
 ts/qZHsp/src/output.jl:198
- [7] #base64encode#3(::Nothing, ::typeof(Base64.base64encode), ::Fun ction, ::MIME{Symbol("image/png")}, ::Vararg{Any,N} where N) at /Use rs/sabae/buildbot/worker/package_macos64/build/usr/share/julia/stdlib/v1.2/Base64/src/encode.jl:206
- [8] base64encode(::Function, ::MIME{Symbol("image/png")}, ::Vararg {Any,N} where N) at /Users/sabae/buildbot/worker/package_macos64/build/usr/share/julia/stdlib/v1.2/Base64/src/encode.jl:203
- [9] _ijulia_display_dict(::Plots.Plot{Plots.PyPlotBackend}) at /Use rs/maria/.julia/packages/Plots/gZHsp/src/ijulia.jl:50
- [10] display_dict(::Plots.Plot{Plots.PyPlotBackend}) at /Users/mari
 a/.julia/packages/Plots/qZHsp/src/init.jl:83
 - [11] #invokelatest#1 at ./essentials.jl:790 [inlined]
 - [12] invokelatest at ./essentials.jl:789 [inlined]
- [13] execute_request(::ZMQ.Socket, ::IJulia.Msg) at /Users/maria/.julia/packages/IJulia/F1GUo/src/execute request.jl:112
- [14] #invokelatest#1 at ./essentials.jl:790 [inlined]
- [15] invokelatest at ./essentials.jl:789 [inlined]
- [16] eventloop(::ZMQ.Socket) at /Users/maria/.julia/packages/IJuli
 a/F1GUo/src/eventloop.jl:8
- [17] (::getfield(IJulia, Symbol("##15#18")))() at ./task.jl:268

In [106]:

```
# Save the relevant arrays of the Lambert record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Lambert.jld2" uivD_L uivD
_L_EDC
```

vi) Binned resampling on grid

```
In [107]:
```

```
@load "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/Lambert.jld2"
```

Out[107]:

```
2-element Array{Symbol,1}:
   :uivD_L
   :uivD_L_EDC
```

• For general analyses, bin on 1 kyr timestep

In [131]:

```
binsize = 1 # each timestep is 1000 years
tmin_L = ceil(minimum(intpD_t_L))
tmax_L = floor(maximum(intpD_t_L))
grid_L = tmin_L + binsize/2 : binsize : tmax_L - binsize/2 # these must be the b
in edges if the bin midpoints shall be at every whole 1 kyr
```

Out[131]:

```
-799.5:1.0:-12.5
```

In []:

In [120]:

```
resampling_method_L = BinnedResampling(grid_L, 1000) # resample 1000 draws (with
substitution) in each bin of the grid
@time L_binned_full = resample(uivD_L, resampling_method_L)
```

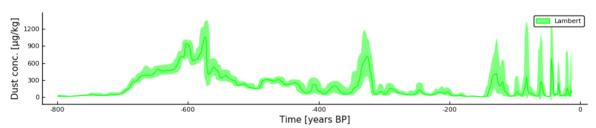
Out[120]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 787 uncertain values coupled with 787 uncertain indic
es

In [122]:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
L = L binned full
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(L.values, 0.5)
bin upper = quantile.(L.values, 0.975) .- bin median
bin lower = bin median .- quantile.(L.values, 0.025)
# time array
binmidpoints L = [L.indices[i].value for i in 1:length(L)]
plot L binned =
plot(size = (1000, 200),
    binmidpoints L, bin median,
    ribbon = (bin lower, bin upper),
    color = :lime,
    label = "Lambert",
    xlabel = "Time [years BP]",
    ylabel = "Dust conc. [\mu g/kg]",
    grid = false
```

Out[122]:



For analysis with the Bereiter record, which is from the same ice core of Epica Dome C, we prepare a version small age uncertainty (uncertainty from lock-in depth co gas bubbles in ice)

In [125]:

```
@time L_binned_full_EDC = resample(uivD_L_EDC, resampling_method_L) # version wi
th small age uncertainty
```

```
39.533399 seconds (38.53 M allocations: 27.983 GiB, 16.77% gc time)
```

Out[125]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas et} containing 787 uncertain values coupled with 787 uncertain indic es

• For higher resolution analyses, we prepare a finer grid on which to bin the time series.

In [189]:

```
binsize_hr125 = 0.125 # 125 year bins, for hr analysis with time step like the r
esolution of the Grant hr record
grid_L_hr125 = tmin_L - binsize_hr125/2 : binsize_hr125 : tmax_L + binsize_hr125
/2 # bin edges, for the bin midpoints to start and end at whole 1 kyrs
```

Out[189]:

```
-800.0625:0.125:-11.9375
```

In [191]:

```
resampling_method_L_hr125 = BinnedResampling(grid_L_hr125, 1000)
@time L_binned_full_hr125 = resample(uivD_L, resampling_method_L_hr125)
```

```
224.662006 seconds (234.18 M allocations: 216.257 GiB, 20.17% gc time)
```

Out[191]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 6305 uncertain values coupled with 6305 uncertain ind
ices

Since the mean resolution was a littlebit below the interplated hr resolution, we also prepare a hr version with 500 yr resolution, to see if these give any different results in analyses.

prepare version for analyses with time step of 500 years

In [192]:

```
binsize_hr500 = 0.5 # time steps like the Bereiter hr record
grid_L_hr500 = tmin_L + 0.5/2 : 0.5 : tmax_L - 0.5/2
resampling_method_L_hr500 = BinnedResampling(grid_L_hr500, 1000)
```

Out[192]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 1575 uncertain values coupled with 1575 uncertain ind
ices

In [196]:

```
@time L_binned_full_hr500 = resample(uivD_L, resampling_method_L_hr500)
```

Out[196]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 1575 uncertain values coupled with 1575 uncertain ind
ices

 Prepare version with smaller age uncertainty for higher resoltuion analysis (time step 500 years) with the Bereiter pCO2 EDC record.

```
In [195]:
@time L_binned_full_hr500_edc = resample(uivD_L_EDC, resampling_method_L_hr500)
67.615267 seconds (62.83 M allocations: 54.653 GiB, 22.72% gc time)
Out[195]:
```

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 1575 uncertain values coupled with 1575 uncertain ind
ices

Save all the binned resampled versions of the Lambert time series in a .jld2 file

```
In [197]:
```

```
# save all the binned resampled versions of the Lambert time series
@save "../Koding/WrangledDataFiles/Binned_ts_fullength/Lambert.jld2" L_binned_fu
ll L_binned_full_EDC L_binned_full_hr125 L_binned_full_hr500 L_binned_full_hr500
_edc
```

```
In [200]:
```

```
# check that they are all there
@load "../Koding/WrangledDataFiles/Binned_ts_fullength/Lambert.jld2"
```

Out[200]:

```
5-element Array{Symbol,1}:
  :L_binned_full
  :L_binned_full_EDC
  :L_binned_full_hr125
  :L_binned_full_hr500
  :L binned_full_hr500 edc
```

4.2 - Martinez-Garcia

- 4.2 Ma marine sediment core of dust (Fe-proxy)
- Data available from Pangaea. DOI: https://doi.org/10.1594/PANGAEA.767460
 (https://doi.org/10.1594/PANGAEA.767460). Supplement of article from Martinez-García et al. (2011).

In [1]:

```
# Read in Martinez-Garcia dataset

filepath_MG = "../../MASTER_2.0/data/dust/Martinez-Garcia_2011/datasets/177-1090
_Fe_dust_acc.tab"
rawD_MG = readdlm(filepath_MG, skipstart = 1, dims = (6165,3))

# name columns
age_MG_ = rawD_MG[:,1]  # Age [ka BP]
Fe_MG_ = rawD_MG[:,2] / 1000  # Accumulation rate Fe [mg/m**2/a]  # * 1/10
00 gives SI-units [g/m^2/year]
dust_MG_ = rawD_MG[:,3] / 10  # Accumulation rate dust [g/cm**2/ka]  # * 1/10
gives SI-units [g/m^2/year]
;

plot(xlabel = "Age [ka BP]", ylabel = string("Accumulation rate ",L"[mg/m^2/a]"
))
plot(age_MG_, Fe_MG_, label = "Fe")
plot!(age_MG_, Fe_MG_, label = "dust")
```

UndefVarError: readdlm not defined

Stacktrace:

[1] top-level scope at In[1]:4

ii) Redefine age as time

In [125]:

```
# reverse the arrays
t_MG = -reverse(age_MG_)  # negative to increase towards present (defined a
s 0)
Fe_MG = reverse(Fe_MG_)  # Accumulation rate Fe [g/m^2/year]
dust_MG = reverse(dust_MG_)  # Accumulation rate dust [g/m^2/year]
;
```

In [227]:

```
# plot Martinez-Garcia dust and Fe record

plot_MG_FeDust =
    plot(xlabel = "Time (kyrs BP)",
        ylabel = L"MAR \ [g/m^2/year]",
        size = (1000,200))

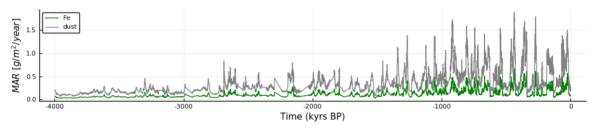
plot!(t_MG, Fe_MG,
        color = :green,
        label = "Fe")

#savefig("../figurar/RawData/Dust/Martinez-Garcia_Fe.pdf")

plot!(t_MG, dust_MG,
        color= :grey,
        label = "dust")

#savefig("../figurar/RawData/Dust/Martinez-Garcia_FeDust.pdf")
```

Out[227]:



Uncertainties

Uncertainties was not included in the dataset, but was reported in the article as following: "After error propagation, we find that the analytical component of the uncertainty is 7.8% [...] of the final value (1 σ) for Fe [...]" (Martinez-García et al. (2011), under Methods summary).

In [126]:

```
# Analytical uncertainty Fe_1\sigma_MG = Fe_MG .* (0.078) # 1\sigma = 7.8% of final value for Fe dust_1\sigma_MG = dust_MG .* (0.084) # 1\sigma = 8.4% of final value for Ti (we don't include this record) (dust);
```

Potential systematic deviations in the age model, as reported in article:

Maximum age model envelope:

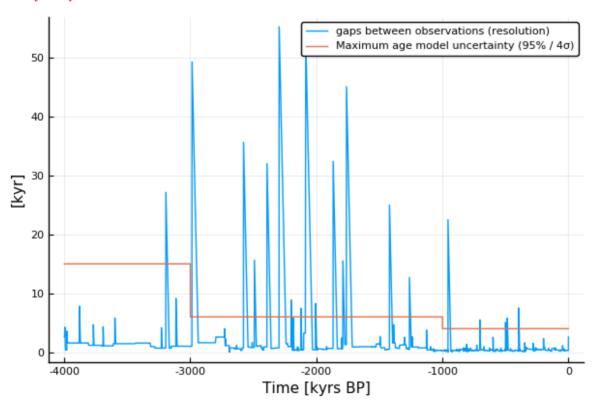
- 4 kyr for the interval 0-1 Myr ago
- 6 kyr for the interval 1-3 Myr ago
- 15 kyr for the interval 3-4 Myr ago

We *interpret* "maximum age model envelope" as the 95% confidence interval, that is 4σ ($\pm 2\sigma$). We should therefore divide by 4 to get 1σ .

In [129]:

```
# Potential systematic deviations in age model
# Maximum age model envelope of 4 kyr for the interval 0-1 Myr ago # (we interpr
et this as +-2\sigma = 4\sigma)
# Maximum age model envelope of 6 kyr for the interval 1-3 Myr ago
# Maximum age model envelope of 15 kyr for the interval 3-4 Myr ago
t 1\sigma MG = zeros(length(t MG));
t_1\sigma_MG[t_MG .> -1000] .= 4/4
                                                        # 4\sigma = 4 kyrs, so divide by
 4 to get 1\sigma
t 1\sigma \ MG[(t \ MG .<= -1000) .& (t \ MG .> -3000)] .= 6/4
t_1\sigma_MG[t_MG .<= -3000] .= 15/4
plot(t MG, diff(t MG), label = "gaps between observations (resolution)")
plot!(t MG, 4*t 10 MG, # Showing stepwise definition of age uncertainty, good.
    xlabel = "Time [kyrs BP]",
    ylabel = "[kyr]",
    label = "Maximum age model uncertainty (95% / 4\sigma)")
```

Out[129]:



Save the arrays

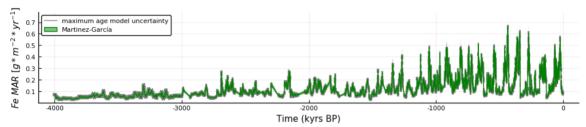
In [232]:

```
@save "../Koding/WrangledDataFiles/BasicArrays/MartinezGarcia.jld2" t_MG t_1\sigma_MG Fe_MG Fe_1\sigma_MG
```

Plot Fe MAR record with value andcertainty and age model uncertainty

```
In [130]:
```

```
# plot of only Fe record
@load "../Koding/WrangledDataFiles/BasicArrays/MartinezGarcia.jld2"
plot Fe MG ageunc =
plot(xlabel = "Time (kyrs BP)",
    ylabel = L"Fe \ MAR \ [g*m^{-2}*yr^{-1}]", # flux = accumulation rate (SYNON)
YMS?)
    size = (1000, 200))
plot!(t MG, Fe MG, xerr = 2 * t 1σ MG, color = :grey, label = "maximum age model
uncertainty") # age uncertainty (95% confidence interval)
plot! (t MG, Fe MG,
    ribbon = (2 * Fe 1\sigma MG, 2 * Fe 1\sigma MG), # analytical uncertainty (95% confid
ence interval)
    color = :green,
    label = "Martinez-García",
    #label = "Martinez-García et al. (2011)",
savefig("../figurar/RawData/Dust/MG Fe noageunc.pdf")
```

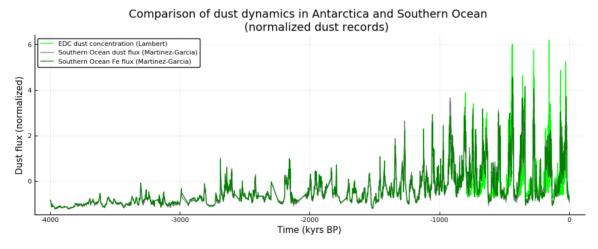


Normalize and plot Lambert and Martinez-Garcia dust records to compare dynamics:

In this normalized plot, we see that, even though the signal strength differs between the sites, the dynamics are the same for wind-born dust in the Southern Ocean (Martinez-Garcia marine core) and on East Antarctica (Lambert EDC ice core).

In [78]:

```
# Plot Lambert and Martinez-Garcia to compare
# normalize first, to avoid cluttering denominations and keep only dynamical inf
ormation
             = (Fe MG .- mean(Fe MG)) / std(Fe MG)
norm Fe MG
norm dust MG = (dust MG .- mean(dust MG)) / std(dust MG)
norm dust L = (dust L .- mean(dust L)) /std(dust L)
# plot
plot(title = "Comparison of dust dynamics in Antarctica and Southern Ocean
    (normalized dust records)",
   xlabel = "Time (kyrs BP)",
   ylabel = "Dust flux (normalized)",
   size = (1000, 400)
    # twinx(), # HOW TO GIVE A SECOND Y-AXIS LABEL TO THE RIGHT?
    # not needed if we normalize the data
plot!(t_L, norm_dust_L,
   color = :lime,
    label = "EDC dust concentration (Lambert)"
plot!(t MG, norm dust MG,
   color = :grey,
    label = "Southern Ocean dust flux (Martinez-Garcia)"
plot!(t_MG, norm_Fe_MG,
   color = :green,
    label = "Southern Ocean Fe flux (Martinez-Garcia)"
savefig("../../Master 2.0/figurar/plot dust comparative normalized.pdf")
```

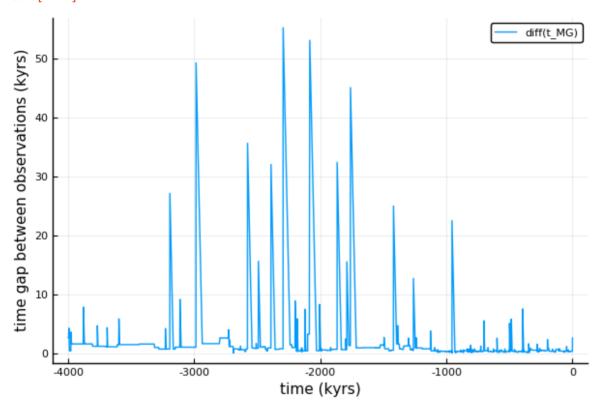


Notes on resolution

Is interpolation needed for the Martinez-Garcia data? We need at least one observation per 1000 years for our grid. Let's check the resolution of this time series:

In [120]:

Out[120]:



There are large variations in resolution of the marine dust record from Martinez-García. There are gaps of several tens of thousands of years at places, which call for interpolation in order to run analyses on millenial scale.

However, it is *less clear what resolution we should give our interpolated data*. On average the record has a high resolution, up to decadal resolution in parts, and with an average of one observation per 640 years.

Will it be acceptable to interpolate to for example a centennial resolution and run high resolution analyses?

We may later determine this by a *senistivity analysis*, testing if we get different results using different interpolation grids.

Interpolation

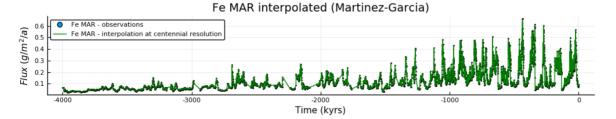
In [84]:

```
# interpolation
# define functions to interpolate
t MG interpolate = LinearInterpolation(t MG, t MG)
t 1 o MG interpolate = LinearInterpolation(t 1 o MG, t MG)
Fe MG interpolate = LinearInterpolation(Fe MG, t MG);
Fe 1\sigma MG interpolate = LinearInterpolation(Fe 1\sigma MG, t MG)
# defining a fine-grained grid that will be used to collect the interpolated val
ues
fine grid MG = ceil(minimum(t MG)) : 0.1 : floor(maximum(t MG)) # one value per
 100 years.
print(fine_grid_MG)
# we then put the interpolated values in a new object interpD MG, one value in e
very bin of the fine time grid
intpD t MG = [t MG interpolate(i) for i in fine grid MG]
intpD t 1\sigma MG = [t 1\sigma MG interpolate(i) for i in fine grid MG]
intpD Fe MG = [Fe MG interpolate(i) for i in fine grid MG]
intpD Fe 1\sigma MG = [Fe 1\sigma MG interpolate(i) for i in fine grid MG]
```

-4000.0:0.1:-1.0

In [89]:

Out[89]:



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NB! it appears some observations are not captured by the interpolation at millennial scale....

Question: From a previous plot (hidden above) it appeared that Some *observations were not captured by the interpolation*, not even when decreasing binsize to year scale (data is given on decade scale). I cannot see if that's the case from this coarse plot, but I only have low-resolution visuals. How can I control that this is not still the case?

Redefining the interpolated dataset as an uivD

In [90]:

```
# redefining the interpolated Fe record intpD_MG as an uivD

t_uiv_MG = [UncertainValue(Normal, intpD_t_MG[i], intpD_t_1\sigma_MG[i]) for i in 1:1
ength(intpD_t_MG)] # Potential deviations in age model, as reported in article
Fe_uiv_MG = [UncertainValue(Normal, intpD_Fe_MG[i], (0.078.*intpD_Fe_MG[i])) for
i in 1:length(intpD_Fe_MG)] # 1\sigma = 7.8\%, reported in article.
uivD_FeMG = UncertainIndexValueDataset(t_uiv_MG, Fe_uiv_MG)
```

Out[90]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 39991 uncertain values coupled with 39991 uncertain i
ndices

Save the relevant arrays in a .jld2 file

```
In [91]:
```

```
# Save the relevant arrays of the Martinez-García record in a .jld2 file
@save "../../MASTER_2.0/Koding/WrangledDataFiles/uivDs/MartinezGarcia.jld2" uivD
_FeMG t_MG
```

vi) Binned resampling

```
In [6]:
```

```
@load "../Koding/WrangledDataFiles/uivDs/MartinezGarcia.jld2"
Out[6]:
```

```
3-element Array{Symbol,1}:
   :intpD_t_MG
   :uivD_FeMG
   :t MG
```

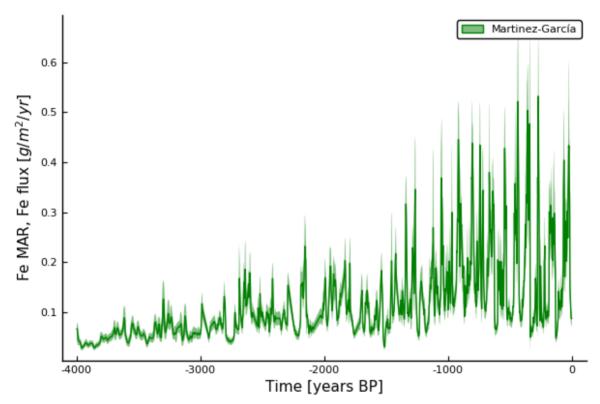
```
In [94]:
```

```
# Binned resampling for the main 1 kyr grid
binsize = 1
tmin MG = ceil(minimum(t MG))
tmax MG = floor(maximum(t MG))
grid MG = tmin MG + binsize/2: binsize : tmax MG - binsize/2
resampling method MG = BinnedResampling(grid MG, 1000)
@time MG binned fullength = resample(uivD FeMG, resampling method MG)
@save "../../MASTER 2.0/Koding/WrangledDataFiles/Binned ts fullength/MartinezGar
cia.jld2" MG binned fullength
727.512050 seconds (685.45 M allocations: 691.831 GiB, 24.55% gc tim
e)
In [7]:
@load "../Koding/WrangledDataFiles/Binned ts fullength/MartinezGarcia.jld2"
Out[7]:
1-element Array{Symbol,1}:
 :MG binned fullength
```

In [96]:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
MG = MG binned fullength
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(MG.values, 0.5)
bin upper = quantile.(MG.values, 0.975) .- bin median
bin lower = bin median .- quantile.(MG.values, 0.025)
# time array
binmidpoints MG = [MG.indices[i].value for i in 1:length(MG)]
plot MG binned =
plot(binmidpoints MG, bin median,
    ribbon = (bin lower, bin upper),
    color = :green,
    label = "Martinez-García",
    xlabel = "Time [years BP]",
    ylabel = string("Fe MAR, Fe flux ", L"[\{g/m^{2}\}/yr]"),
    grid = false
    )
```

Out[96]:



--

Points of discussion

- SHOULD we do hr analysis of this record?
- if so, what is a reasonable BINSIZE for hr analysis?

In [121]:

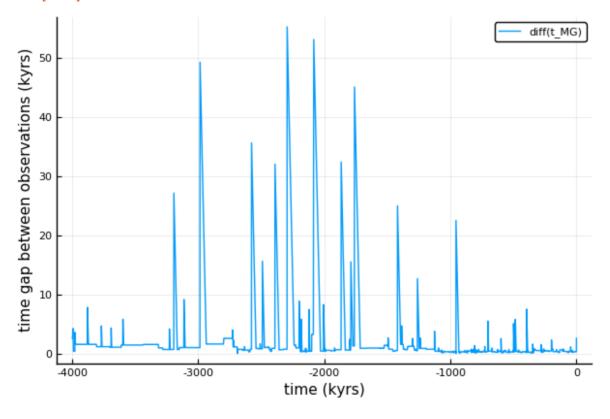
```
print("Mean resolution of the Martinez-García record is ", mean(diff(t_MG)), "
  (650 years).
Highest resolution of the Martinez-García record is ", minimum(diff(t_MG)), " (3
0 years).
largest time gap is ", maximum(diff(t_MG)), " (55 000 years)." )
plot_MG_resolution
```

Mean resolution of the Martinez-García record is 0.6488822193380921 (650 years).

Highest resolution of the Martinez-García record is 0.02999999999997 2715 (30 years).

largest time gap is 55.219999999999 (55 000 years).

Out[121]:



Is the record suitable for high resolution analysis?

Even though Martinez-Garcia record is a high resolution record at the baseline, the large time gaps represent a problem for using the record for high resolution analysis. However, we notice that the first part of the record (post-MPT) has smaller time gaps and overall higher resolution. Perhaps we can run higher resolution analyses for selected time intervals of the record? Let's check.

The other records suitable for high resolution analysis are the Grant GSL record, the La2004 record and the Chalk syn-MPT pCO2 record (and possibly Lambert and Bereiter?). We therefore want to have a closer look at the resolution for the time intervals overlapping with these records.

1. From the plot above, the **post-MPT time interval** does seem to have high resolution for the Martinez-García record. Let's check the resolution for the relevant time interval.

Mostly, the resolution is just above a datapoint per 500 years.

In [8]:

```
#common time interval for the Grant and Martinez-García records
tmin short = -800
tmin G = -491 # recall for now, since haven't read in Grant data in this iterati
on
tmax short = 0
tmin MG = t MG[1]
tmax MG = t MG[end]
maximum([tmin short, tmin MG]) : minimum([tmax short, tmax MG])
Out[8]:
-800.0:1.0:-1.0
In [31]:
# Cut the relevant interval from the MG record
t MG short = t MG[(t MG .> tmin short)] # MG record for the past 800 kyrs
# make a plot to show the resolution of the interval
plot(t MG short, diff(t MG short),
    label = "resolution",
    xlabel = "Time (kyrs BP)",
    xticks = (-800:100:0),
    ylabel = "time gaps [kyrs]",
    yticks = (0:1:8),
    title = "MG record resolution over the past 800 kyrs")
```

```
UndefVarError: t_MG not defined
Stacktrace:
[1] top-level scope at In[31]:1
```

In [10]:

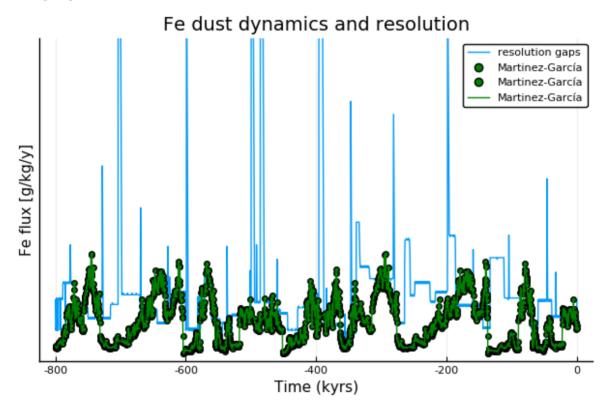
```
print("For the last 800 kyrs of the Martinez-García record,
mean resolution is ", mean(diff(t_MG_short)), " (330 years).
Highest resolution for the period is ", minimum(diff(t_MG_short)), " (140 year
s).
Largest time gap is ", maximum(diff(t_MG_short)), " (ca. 7500 years)." )
```

```
For the last 800 kyrs of the Martinez-García record, mean resolution is 0.3290329218106996 (330 years). Highest resolution for the period is 0.13999999999998636 (140 years). Largest time gap is 7.519999999999982 (ca. 7500 years).
```

Is it acceptable to run high resolution analysis on this time interval? Let's first also check if the record in general has fast or slow dynamics.

In [19]:

Out[19]:



I'd say the time series looks like it has fast dynamics where it has high resolution, which does not speak in favour of interpolating for higher resolution analyses.

Input? Interpret this as fast dynamics, or "too small fluctuations to care about"?

But let's give it a try anyways with high resolution analysis for fun, but keep in mind the assumptions made when interpreting the results. We have more or less the same case with the Lambert time series, so that might add *some* robustness to the result.

binsize_hr = 0.125 (like Grant record) will be too fine. Let's stay above the mean resolution, and interpolate values for the high resolution grid to one value each 500 years.

Can fake data from linear interpolation might bias the result to a causal connection ("precede and predict"))?

In [12]:

```
# common grid MG - G with hr binsize 0.5 kyrs
    # NB. this implies we have to make a version for Grant hr at 500 as well

binsize_hr500 = 0.5

grid_MG_hr500 = tmin_short - binsize_hr500/2 : binsize_hr500 : tmax_MG + binsize
_hr500/2 # to get values at binmidpoints, substract half a binsize from tmin, an d add half a binsize to tmax
resamplingmethod_MG_hr500 = BinnedResampling(grid_MG_hr500, 1000)

@time MG_binned_postmpt_hr500 = resample(uivD_FeMG, resamplingmethod_MG_hr500)
# (191 seconds to run BinnedResampling on this grid)
```

```
282.095583 seconds (289.81 M allocations: 277.983 GiB, 19.30% gc time)
```

Out[12]:

UncertainIndexValueDataset{UncertainIndexDataset,UncertainValueDatas
et} containing 1600 uncertain values coupled with 1600 uncertain ind
ices

In [84]:

```
# If we've binned correctly these should be the same:
print(tmin_short : tmax_MG, "
   ", MG_binned_postmpt_hr500.indices[1].value : MG_binned_postmpt_hr500.indices[e
nd].value) # first and last binmidpoints of binned time series
#hurra!
```

```
-800.0:1.0:-1.0
-800.0:1.0:-1.0
```

In [20]:

@save "../Koding/WrangledDataFiles/Binned_ts_fullength/MartinezGarcia.jld2" MG_b
inned fullength MG_binned_postmpt_hr500

```
In [22]:
```

```
# check that saved the 2 binned uivDs properly
@load "../Koding/WrangledDataFiles/Binned_ts_fullength/MartinezGarcia.jld2"
# all good
```

Out[22]:

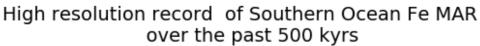
```
2-element Array{Symbol,1}:
:MG_binned_fullength
:MG_binned_postmpt_hr500
```

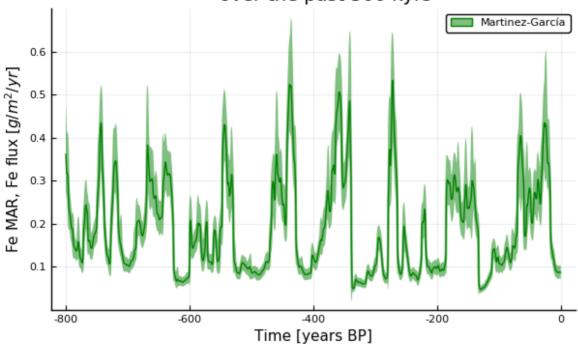
is 1000 draws enough to give a proper probability distribution?

In [28]:

```
### Plot the binned resampled uivD time series with the 95% confidence interval
MG = MG binned postmpt hr500
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(MG.values, 0.5)
bin upper = quantile.(MG.values, 0.975) .- bin median
bin lower = bin median .- quantile.(MG.values, 0.025)
# time array
binmidpoints MG = [MG.indices[i].value for i in 1:length(MG)]
plot MG binned =
plot(title = "High resolution record of Southern Ocean Fe MAR
    over the past 500 kyrs",
    binmidpoints MG, bin median,
    ribbon = (bin lower, bin upper),
    color = :green,
    label = "Martinez-García",
    xlabel = "Time [years BP]",
    ylabel = string("Fe MAR, Fe flux ", L"[\{g/m^{2}\}/yr^{3}]"),
    \#xticks = (-800:0.5:0)
    )
```

Out[28]:





--

· High resolution record syn-MPT

It would also be interesting to do a high resolution analysis of the dynamics between Fe dust, insolation and pCO2 **during the MPT**. The records available for this are Martinez-García, La2004 and Chalk. How is the resolution of the Martinez-García Fe MAR record across the MPT?

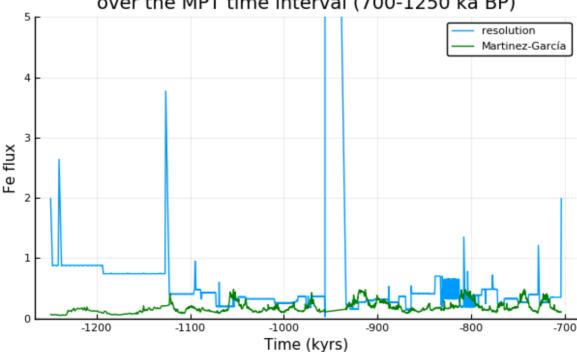
In [35]:

```
# MG record resolution during the MPT
# cut out the time interval of the MG record
t MG synMPT = t MG[(t MG \cdot > -1250) .& (t MG \cdot < -700)] # relevant interval for an
alysis with La2004
print("For the syn-MPT time interval
Mean resolution is ", mean(diff(t_MG_synMPT)), " (350 years).
Smallest time gap is ", minimum(diff(t_MG_synMPT)), " (30 years).
Largest time gap is ", maximum(diff(t MG synMPT)), " (22.5 kyrs)." )
# plot to visualize resolution and if there are significant time gaps
plot(t MG synMPT, diff(t MG synMPT), ylims = (0,5), label = "resolution", xlabel
= "Time (kyrs BP)", ylabel = "time gaps [kyrs]", title = "MG record resolution
over the MPT time interval (700-1250 ka BP)")
#=Let's also check if the record in general has fast or slow dynamics
(if slow dynamics in the high resolution parts of the record, it might be defend
able to interpolate for a higher resolution). =#
Fe MG synMPT = Fe MG[(t MG \cdot > -1250) \cdot \& (t MG \cdot < -700)]
plot!(t MG synMPT, Fe MG synMPT, color = :green, xlabel = "Time (kyrs)", ylabel
= "Fe flux", label = "Martinez-García") # title = "Fe dust dynamics 1.250 - 1.08
0 Ma BP")
```

For the syn-MPT time interval
Mean resolution is 0.37929117442668525 (350 years).
Smallest time gap is 0.0299999999972715 (30 years).
Largest time gap is 22.5 (22.5 kyrs).

Out[35]:

MG record resolution over the MPT time interval (700-1250 ka BP)



With a mean resolution of 380 years, we deem it reasonable to run high resolution analysis with the Martinez-García dust record over the syn-MPT, with a timestep to 0.5 kyrs.

• Let's check if the Martinez-García Fe MAR record has sufficiently high resolution for the time interval overlapping with the Chalk record (1080-1250 ka BP).

In [36]:

```
# cut out the time interval of the MG record overlapping with the Chalk pCO2 rec
ord
tmin_C = -1242
tmax_C = -1088

t_MG_synC = t_MG[(t_MG .>= tmin_C) .& (t_MG .<= tmax_C)]
t_MG_synC[1] : t_MG_synC[end] # (does not concord because t_MG is not binned on
a regular time grid yet)</pre>
```

Out[36]:

-1241.98:1.0:-1088.98

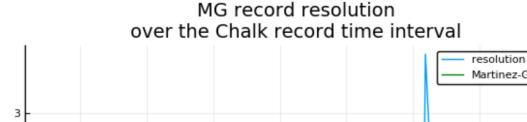
In [39]:

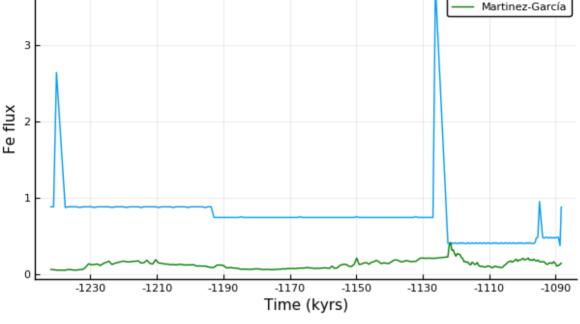
```
# MG record resolution for the Chalk record time interval
print("Mean resolution of the Martinez-García record over the Chalk record time
interval is ", mean(diff(t MG synC)), " (680 years).
Highest resolution for the period is ", minimum(diff(t_MG_synC)), " (320 years).
Largest time gap is ", maximum(diff(t MG synC)), " (3.770 years)." )
# plot to visualize resolution
plot(t MG synC, diff(t MG synC), label = "resolution", xlabel = "Time (kyrs BP)"
, ylabel = "time gaps [kyrs]", title = "MG record resolution
over the Chalk record time interval")
#=Let's also check if the record in general has fast or slow dynamics
(if slow dynamics in the high resolution parts of the record, it might be defend
able to interpolate for a higher resolution). =#
Fe MG synC = Fe MG[(t MG .>= tmin C) .& (t MG .<= tmax C)]
plot!(t MG synC, Fe MG synC, color = :green, xlabel = "Time (kyrs)", ylabel = "F
e flux", label = "Martinez-García", title = "Fe dust dynamics 1.250 - 1.080 Ma B
P")
plot!(title = "MG record resolution
over the Chalk record time interval")
plot!(xticks = (-1250:20:1080))
```

Mean resolution of the Martinez-García record over the Chalk record time interval is 0.680132743362832 (680 years). Highest resolution for the period is 0.3700000000011823 (320 years).

Largest time gap is 3.76999999999982 (3.770 years).

Out[39]:





We see that most of the record has a coarser resolution than 500 years for most of the MTP time interval (mean resolution of one value per 680 years). Only the very first part of the record has a 500 yr resolution (ca 1090-1120 kyrs BP). We could consider running a high resolution analysis over that time interval, but the time series length would unfortunately be too short to give robust results (need at least 100 datapoints, and this interval would give us only 30 kyrs/500 yrs = 60 datapoints). We therefore conclude that **the preMPT syn-Chalk time interval of the record is not suitable for a higher resolution analysis**. We will therefore run the analysis only on the 1 kyr timestep grid, for which the Martinez-Garcia record has sufficient resolution.

• Finally, let's check the resolution of the record for the pre-MPT which overlaps with the Elderfield record

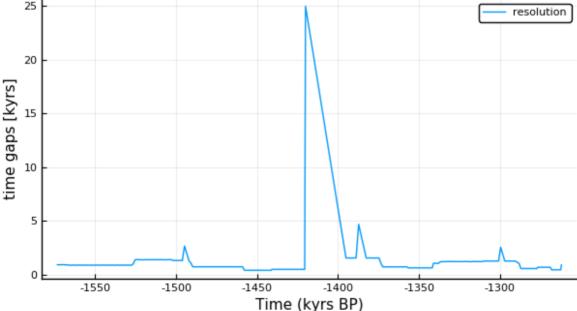
In [53]:

```
# MG record resolution pre-MPT
# cut out the pre MPT time interval of the MG record that overlaps with the Elde
rfield record
t MG preMPT 1500 = t MG[(t MG \cdot > -1574) \cdot \& (t MG \cdot < -1250)] # relevant interval
 for analysis with La2004
print("Mean resolution of the Martinez-García record over the syn-MPT time inter
valis ", mean(diff(t_MG_preMPT_1500)), " (900 years).
Highest resolution for the period is ", minimum(diff(t_MG_preMPT_1500)), " (400
Largest time gap is ", maximum(diff(t_MG_preMPT 1500)), " (25 kyrs)." )
# plot to visualize resolution and if there are significant time gaps
plot(t MG preMPT 1500, diff(t MG preMPT 1500),
    label = "resolution",
    xlabel = "Time (kyrs BP)",
    ylabel = "time gaps [kyrs]",
    title = "MG record resolution
    over the Elderfield pre-MPT time interval
    (1574-1250 ka BP)")
```

Mean resolution of the Martinez-García record over the syn-MPT time intervalis 0.8976945244956772 (900 years). Highest resolution for the period is 0.4099999999998545 (400 years). Largest time gap is 24.960000000000036 (25 kyrs).

Out[53]:





With a mean resolution of 900 years between each datapoint, and a time gap of 25 kyrs, this record is **not suitable for higher resolution analyses** over the pre-MPT time period. We will only use the standard 1 kyr grid already prepared.

Finally, let's check the resolution of the record for the pre-MPT which overlaps with the Elderfield record

In [1]:

```
# MG record resolution pre-MPT
# cut out the pre MPT time interval of the MG record that overlaps with the Elde
rfield record
t MG preMPT 4000 = t MG[(t MG .> tmin) .& (t MG .< -1250)] # relevant interval f
or analysis with La2004
print("Mean resolution of the Martinez-García record over the syn-MPT time inter
valis ", mean(diff(t MG preMPT 4000)), " (900 years).
Highest resolution for the period is ", minimum(diff(t MG preMPT 4000)), " (400
years).
Largest time gap is ", maximum(diff(t MG preMPT 4000)), " (25 kyrs).")
# plot to visualize resolution and if there are significant time gaps
plot(t MG preMPT 4000, diff(t MG preMPT 4000),
    label = "resolution",
   xlabel = "Time (kyrs BP)",
   ylabel = "time gaps [kyrs]",
   title = "MG record resolution
   over the Elderfield pre-MPT time interval
    (1574-1250 ka BP)")
```

```
UndefVarError: t_MG not defined
Stacktrace:
[1] top-level scope at In[1]:1
```

Summary NB1

- We now have all the data on a regular time grid with time steps of 1 kyr. We have also some higher resolution regular grids, where we considered there was sufficient resolution in hte dataset for this.
- Associated uncertainties in time has been shuffled over to the value index, by resampling on the regular grid. The uncertainties carried on as kernel density estimates, which allows further resampling to estimate confidence intervals onwards.
- The wrangled data (the time series binned resampled on regular time grids) have been saved as .jld2 files, and can easily be imported in other notebooks for analysis.

The next step for the wrangled data will be to define a common time interval for the time series. Once the time series are on the same time grid (resolution and time interval), we can compute the predictive asymmetry from one to the other. This is done in the NBRs.

In []:	
In []:	
III [].	
In []:	
In []:	
In []:	
In []:	
In []:	
In []:	

@save "../Koding/WrangledDataFiles/Binned_ts_fullength/ALL_ts_binned.jld2"
LR04_binned_fullength_fullageunc LR04_binned_fullength_noageunc La2004_insol65N_fullength
La2004_insol65N_fullength_hr SL_binned_fullength_ageunc SL_binned_fullength_noageunc
E_binned_fullength_ageunc E_binned_fullength_noageunc G_binned_fullength G_binned_fullength_hr
R_binned_full B_binned_fullength_ageunc B_binned_fullength_ageunc_hr B_binned_fullength_noageuncEDC
B_binned_fullength_noageuncEDC_hr C_binned_1kyrgrid C_binned_hr L_binned_full L_binned_full_EDC
L_binned_full_hr MG_binned_fullength MG_binned_postmpt_hr500 # missing L_binned_full_EDC_hr

```
In [54]:
```

```
@load "../Koding/WrangledDataFiles/Binned_ts_fullength/ALL_ts_binned.jld2"
```

Out[54]:

```
22-element Array{Symbol,1}:
 :LR04 binned fullength fullageunc
 :LR04 binned fullength noageunc
 :La2004 insol65N fullength
 :La2004_insol65N_fullength hr
 :SL binned fullength ageunc
 :SL binned fullength noageunc
 :E binned fullength ageunc
 :E binned fullength noageunc
 :G binned fullength
 :G binned fullength hr
 :R binned full
 :B binned fullength ageunc
 :B binned fullength ageunc hr
 :B binned fullength noageuncEDC
 :B binned fullength noageuncEDC hr
 :C binned 1kyrgrid
 :C binned hr
 :L binned full
 :L binned full EDC
 :L binned full hr
 :MG binned fullength
 :MG binned short hr
```

save the wrangled data as a .jld2 file @save "../../MASTER_2.0/Koding/WrangledDataFiles/WrangledData.jld2" uivD_SL t_SL uivD_E intpD_t_E uivD_G t_G uivD_R t_R uivD_LR04 intpD_t_LR04 uivD_La2004 t_La2004 La2004 uivD_B uivD_B_EDC intpD_t_B uivD_C t_C uivD_L uivD_L_EDC intpD_t_L intpD_dust_L uivD_FeMG intpD_t_MG # This data can now be loaded in and read in any notebook, through the @load command: @load "/Users/maria/Dropbox/MASTER_2.0/Koding/WrangledData.jld2"

In [55]:

This data can now be loaded in and read in any notebook, through the @load com
mand:
@load "/Users/maria/Dropbox/MASTER 2.0/Koding/WrangledData.jld2"

SystemError: opening file "/Users/maria/Dropbox/MASTER_2.0/Koding/Wr angledData.jld2": No such file or directory

Stacktrace:

- [1] #systemerror#44(::Nothing, ::typeof(systemerror), ::String, ::B
 ool) at ./error.jl:134
 - [2] systemerror at ./error.jl:134 [inlined]
- [3] #open#311(::Bool, ::Bool, ::Bool, ::Bool, ::typeof(ope n), ::String) at ./iostream.jl:289
 - [4] #open at ./none:0 [inlined]
- [5] JLD2.MmapIO(::String, ::Bool, ::Bool, ::Bool) at /Users/maria/.julia/packages/JLD2/hB4ya/src/mmapio.jl:100
- [6] openfile at /Users/maria/.julia/packages/JLD2/hB4ya/src/JLD2.j
 l:194 [inlined]
- [7] #jldopen#9(::Bool, ::Bool, ::typeof(jldopen), ::String, ::Bool, ::Bool, ::Bool, ::Type{JLD2.MmapIO}) at /Users/maria/.julia/package s/JLD2/hB4ya/src/JLD2.jl:231
- [8] jldopen at /Users/maria/.julia/packages/JLD2/hB4ya/src/JLD2.jl: 203 [inlined] (repeats 2 times)
- [9] #jldopen#10(::Base.Iterators.Pairs{Union{},Union{},Tuple{},Name
 dTuple{(),Tuple{}}}, ::typeof(jldopen), ::String, ::String) at /User
 s/maria/.julia/packages/JLD2/hB4ya/src/JLD2.jl:293
- [10] jldopen at /Users/maria/.julia/packages/JLD2/hB4ya/src/JLD2.j
 1:288 [inlined] (repeats 2 times)
- [11] @load(::LineNumberNode, ::Module, ::Any, ::Vararg{Any,N} where
- N) at /Users/maria/.julia/packages/JLD2/hB4ya/src/loadsave.jl:99

Recalling notations

The time arrays (prefix t_, or intpD_t_ for interpolated datasets) contain the age values, formulated as negative values (kyrs from present). The data (prefix uivD_) contain the measured/reconstructed values, with their associated uncertainties carried on as kernel density estimates (KDE). The suffixes are abbreviations for the records, recalled here below:

- **LR04** = *Lisiecki & Raymo (2005)*. $\delta^{18}O$ global reference stack, spanning the last 5.3 Myr.
- SL = Spratt & Lisiecki (2016). GSL stack (PCA of 5 different GSL records). Millennial resolution over the last 798 kyr.
- **E** = *Elderfield et al. (2012)*. GSL record, spanning the last 1.5 Myr (Mg/Ca temperature deconvolution of d180).
- **G** = *Grant et al. (2014)*. Red Sea RSL record, spanning the last 500 kyr.
- **R** = Rohling et al. (2014). Mediterranean RSL record, spanning the last 5.3 Myr
- La2004 = Laskar (2004). Numerical solution for northern hemisphere summer insolation (top of atmosphere solar flux for 21st of June at 65°N). 50 Myr back and forth in time from present without significant uncertainty associated (uncertain beyond that).
- **B** = Bereiter et al. (2015). pCO2 record from ice core at Epica Dome C, spanning ca 800 ka.
- **C** = Chalk et al. (2017). pCO2 record from $\delta^{11}B$ proxy, spanning ca 1.090-1.240 Ma.
- **L** = *Lambert et al (2008)*. Dust record from dust concentration in ice core at Epica Dome C, spanning ca 800 kyr. (Note, the data has been set to updated age model since publication.)
- **MG** = *Martinez-Garcia et al. (2011)*. Record of iron mass accumulation rate (Fe MAR) in the Southern Ocean, spanning the last 4 Myr.

Overview plots of raw data

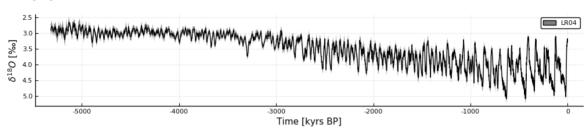
In []:

@load "../data/sea-level/LR04/LR04 wrangled.jld2"

In [3]:

```
# LR04
@load "../data/sea-level/LR04/LR04_wrangled.jld2"
plot_LR04_noageunc = plot(
    t_LR04, d180_LR04, ribbon = (2*d180_1o_LR04),
    xlabel = "Time [kyrs BP]", ylabel = L"$\delta^{18}O$ [$\perthousand$]",
    yflip = true, # by convention, we plot d180 on a reversed axis #(mirror read ability of GSL, which constitutes on average 70% of the signal)
    color = :black, label = "LR04", size = (1000,200))
#plot!(t_LR04, d180_LR04, yerr = 2*t_1o_LR04) # 95% confidence interval of age m odel uncertainty
```

Out[3]:



sea level records

gsl overview plot

In []:

```
plot_gsl_overview =
    plot(plot_SL_binned, plot_E_binned, plot_G_binned, plot_R_binned,
    link = :x,
    layout = grid(4,1),
    xlim = (-2000,0)
)
```

In []:

```
# Subplot of GSL records

#plot_GSL_overview =
plot(#title = "Global sea level records"
    plot_SprattLisiecki_raw_time,
    plot_Elderfield_raw_ageunc,
    plot_Grant_raw,
    plot_Rohling_raw,
    layout = (4,1),
    #share= :x,
    link = :x,
    size = (1000,400),
    xlims = (-1500:0), #(0:1500) # Why won't it cut at -1500 kyrs?
    ticks = false #(-1500:100:0)
    )
#savefig("../../Master_2.0/figurar/RawData/GSL/plot_GSL_overview.pdf")
```

gsl comparative plot

```
# Plot to compare all GSL reconstructions
plot GSL comparative full =
plot(#title = "Comparative plot of GSL records",
    xlabel = "Time [ka BP]",
    ylabel = "GSL [m]",
    \#x \lim = (-1000, 0),
                         # plot only a range 0:1000 ka
    ticks = (0:100:-1000), # not working
    grid= :true,
                         # not working
    size = (1000, 200),
    share = :(x,y),
                         # working?
    #legend = :bottomleft
plot!(t R, RSL mean R,
    ribbon = (2*RSL 1\sigma R), # \pm 2\sigma, aka 95% confidence interval
    color = :skyblue,
    label = "Rohling"
#plot!(fine grid E, intpD_GSL_E, yerr = (2*t_1\sigma_E), color = :royalblue, ms = 0.
1) # LR04 age model uncertainty
plot!(fine grid E, intpD GSL E,
    ribbon = (2*intpD 1\sigma E),
    color = :royalblue,
    label = "Elderfield"
#plot!(t SL, GSL mean SL, yerr = (2*t 1\sigma SL), color = :darkblue, ms = 0.1) # LR0
4 age model uncertainty
plot!(t SL, GSL mean SL,
    ribbon = (GSL mean SL - GSL SL err lo, GSL SL err up - GSL mean SL), # 95% C
Ι
    color = :darkblue,
    label = "Spratt & Lisiecki"
    )
plot!(t_G, RSL G,
                                                                            # 95% C
    ribbon = (RSL_G - RSL_q95lo_G, RSL_q95up_G - RSL_G),
Ι
    color = :cyan,
    label = "Grant"
savefig("../figurar/RawData/GSL/plot comparative GSL.pdf")
```

gsl - LR04 overview

In []:

```
# Compare GSL records with d180 record

plot_overview_LR04_GSL =
plot(plot_LR04_raw_time,
    plot_GSL_comparative_full,
    layout = grid(2,1),
    link = :x)
savefig("../figurar/RawData/LR04/overview_LR04_GSL.pdf")
```

```
In []:
In []:
In []:
# plot Lambert time series
# plot Lambert time series
```

```
# plot Lambert time series
# @load "../data/dust/Lambert_wrangled_raw.jld2"

plot_Lambert_time_noageunc =
plot(#title = "EDC aeolian dust record (Lambert lpc record)",
    size = (1000, 200),
    xlabel = "time (kyrs BP)",
    ylabel = "dust conc. [µg/kg]")

#scatter!(t_L, dust_L, xerr = 2 * t_1\sigma_L, ms = 1, color = :grey, label = "AICC 2012 age model uncertainty (2\sigma)")
plot!(t_L, dust_L,
    color = :lime,
    #label = "EDC dust record (Lambert)"
    label = "Lambert")

#savefig("../figurar/RawData/Dust/Lambert_aicc2012_noageunc.pdf")
```

```
In [ ]:
```

```
# plot of only Fe record
@load "../data/dust/MG_raw.jld2"

plot_Fe_MG_noageunc =
  plot(xlabel = "Time (kyrs BP)",
      ylabel = L"Fe \ MAR \ [g*m^{-2}*yr^{-1}]", # flux = accumulation rate (SYNON YMS?)
      size = (1000,200))
#plot!(t_MG, Fe_MG, xerr = 2 * t_1\sigma_MG, color = :grey, label = "maximum age mode l uncertainty") # age uncertainty (95% confidence interval)
plot!(t_MG, Fe_MG,
      ribbon = (2 * Fe_1\sigma_MG), # analytical uncertainty (95% confidence interval)
      color = :green,
      label = "Martinez-García",
      #label = "Martinez-García et al. (2011)",
      )

#savefig("../figurar/RawData/Dust/MG_noageunc.pdf")
```

Normalize and plot Lambert and Martinez-Garcia dust records to compare

```
# Plot Lambert and Martinez-Garcia to compare
# normalize first, to avoid cluttering denominations and keep only dynamical inf
ormation
norm Fe MG = (Fe MG .- mean(Fe MG)) / std(Fe MG)
norm dust MG = (dust MG .- mean(dust MG)) / std(dust MG)
norm dust L = (dust L .- mean(dust L)) /std(dust L)
# We choose not to plot age uncertainty, not to clutter the plot
plot dust comparative L FeMG = # wMGdust =
plot(#title = "Comparative plot of dust records",
    xlabel = "Time (kyrs BP)", ylabel = "Dust flux (normalized)", size = (1000,
200))
    # twinx(), # HOW TO GIVE A SECOND Y-AXIS LABEL TO THE RIGHT?# not needed if
we normalize the data
# MG dust record (will not be used in analyses)
#plot!(t MG, norm dust MG, color = :grey, label = "Martinez-García dust", # labe
1 = "Southern Ocean dust flux (Martinez-Garcia)"
    ribbon = (2*0.084*norm dust MG))
                                                                            # 1\sigma =
8.4%. We plot the 95% condifence interval (\pm 2\sigma)
# MG Fe record
plot!(t MG, norm Fe MG, color = :green, label = "Martinez-García", # label = "S
outhern Ocean Fe flux (Martinez-Garcia)"
                                                                       \# 1\sigma = 7.
    ribbon = (2*0.078*norm Fe MG))
8%. We plot the 95% condifence interval (\pm 2\sigma)
# Lambert dust record
plot!(t_L, norm_dust_L, color = :lime, label = "Lambert" ) #label = "EDC dust co
ncentration (Lambert)")
#savefig("../figurar/plot dust comparative.pdf")
```

In this normalized plot, we see that, even though the signal strength differs between the sites, the dynamics are the same for wind-born dust in the Southern Ocean (Martinez-Garcia marine core) and on East Antarctica (Lambert EDC ice core).

```
plot(plot_LR04_raw_time, # rerun above. Please, save the necessary stuff
    plot_GSL_comparative,
    plot_insolation, # rerun above. Please, save the necessary stuff
    plot_pC02_comparative_B_C,
    #plot_pC02_comparative_wHonisch,
    plot_dust_comparative_LFeMG,
    #plot_dust_comparative_wMGdust
    layout = grid(5,1)
)
```

Overview plots of wrangled time series

(interpolated and resampled on a regular grid)

```
In [2]:
```

```
#### d180 / sea level / ice volume proxies
@load "../WrangledDataFiles/Binned ts fullength/LR04.jld2" # LR04 - d180 stack
@load "../WrangledDataFiles/Binned ts fullength/SprattLisiecki.jld2" # Spratt&Li
siecki - global sea level stack, spanning last 800 kyrs (post-MPT)
@load "../WrangledDataFiles/Binned ts fullength/Elderfield.jld2" # Elderfield -
 global sea level record, spanning last 1.5 Myrs (pre- syn- & post-MPT)
@load "../WrangledDataFiles/Binned ts fullength/Grant.jld2" # Grant - Red Sea RS
L record, spanning last 500 kyrs (post-MPT)
@load "../WrangledDataFiles/Binned ts fullength/Rohling.jld2" # Rohling - Medite
rranean RSL record, spanning last 5.3 Myrs (pre- syn- & post-MPT)
#### insolation time series
@load "../WrangledDataFiles/La2004.jld2" # La2004 - numerical solution for north
ern hemisphere summer insolation, last 5 Ma computed using AnalySeries (Paillar
d, 1994). (pre- syn- & post-MPT).
#### pCO2 records/proxies
@load "../WrangledDataFiles/Binned ts fullength/Bereiter nointp.jld2" # Bereiter
- post-MPT pCO2 record - with age model uncertainty
@load "../WrangledDataFiles/Binned ts fullength/Bereiter noageuncEDC intp.jld2"
# Bereiter - post-MPT pCO2 record - without age model uncertainty (for analysis
with Lambert EDC dust record). NB. this time series containts interpolated valu
@load "../WrangledDataFiles/Binned ts fullength/Chalk.jld2" # Chalk - syn-MPT pC
02 record (time interval 1.088-1.242 Ma BP, updated cut to 1.092-1.240 Ma BP)
#### dust records
@load "../WrangledDataFiles/Binned ts fullength/Lambert.jld2" # Lambert - post-M
PT Antarctic dust record (spanning last 800 kyrs)
@load "../WrangledDataFiles/Binned ts fullength/MartinezGarcia.jld2" # Martinez-
García - 4 Ma Southern Ocean Fe dust record (pre- syn- & post-MPT)
```

Bereiter BinnedResampled without interpolation

Define the first and last time point in each record. These will be used in the NBRs to determine the common grid.

In [3]:

```
# Define the first and last time point in each record
##### d180 record
# LR04
LR04 = LR04 binned fullength fullageunc
# assign to tmin LR04 the first time value in LR04 (that is, the LR04 record sta
rts at ``tmin LR04`` years BP)
tmin LR04 = LR04.indices[1].value # the binned LR04 record starts at -802 kyrs
BP (binmidpoint)
tmax LR04 = LR04.indices[end].value # -0.3 kyrs BP ...last age value in the LR04
record (binmidpoint)
print("The LR04 d180 record spans from ", tmin LR04, " to ", tmax LR04, " kyrs B
P.")
##### insolation time series
# La2004 (Ins)
La2004 = La2004 insol65N fullength
#= recall, insolation is not an uncertain index value dataset type.
We therefore use two arrays (time index and insolation value) =#
tmin La2004 = La2004 t fullength[1] # 5000
tmax La2004 = La2004 t fullength[end] # 0 kyrs BP (present day)
print("
The La2004 insolation time series (Ins) spans from ", tmin La2004, " to ", tmax
La2004, " kyrs BP.")
##### GSL records
# SprattLisiecki (SpraSL)
SL = SL binned fullength ageunc
tmin SL = SL.indices[1].value # -797.0 kyrs BP
tmax SL = SL.indices[end].value # -1.0 kyrs BP
print("
Spratt Lisiecki GSL stack (SpraSL) spans from ", tmin SL, " to ", tmax SL, " ky
rs BP.")
# Elderfield (EldSL)
E = E binned fullength ageunc
tmin E = E.indices[1].value # -1574 kyrs BP
tmax E = E.indices[end].value # -8 kyrs BP
print("
The Elderfield GSL record (EldSL) spans from ", tmin E, " to ", tmax E, " kyrs B
P.")
# Grant (GraSL)
G = G binned fullength
tmin G = G.indices[1].value # -491 kyrs BP
tmax G = G.indices[end].value # -1 kyrs BP
The Grant sea level record (GraSL) spans from ", tmin_G, " to ", tmax_G, " kyrs
 BP.")
# Rohling (RohSL)
R = R_binned_full
tmin R = R.indices[1].value # -5329 kyrs BP
tmax R = R.indices[end].value # -1.0 kyrs BP
print("
The Rohling sea level record (RohSL) spans from ", tmin R, " to ", tmax R, " kyr
```

```
s BP.")
##### pC02 records
# Bereiter (BerCO2)
B = B binned fullength ageunc
tmin B = B.indices[1].value # -802 kyrs BP
tmax B = B.indices[end].value # -3 kyrs BP
print("
The Bereiter pCO2 record (BerCO2) spans from ", tmin B, " to ", tmax B, " kyrs B
P.")
# Chalk (ChaCO2)
C = C binned
tmin C = C.indices[1].value # - 1242 kyrs BP
tmax C = C.indices[end].value # -1088 kyrs BP
The Chalk pCO2 record (ChaCO2) spans from ", tmin C, " to ", tmax C, " kyrs BP."
##### dust records
# Lambert (IceDust)
L = L binned full
tmin L = L.indices[1].value # -799 kyrs BP
tmax L = L.indices[end].value # - 2 kyrs BP
print("
The Lambert dust record (IceDust) spans from ", tmin L, " to ", tmax L, " kyrs B
P.")
# Martinez-García (MarFe)
MG = MG binned fullength
tmin MG = MG.indices[1].value # -3999 kyrs BP (4 Myrs)
tmax MG = MG.indices[end].value # -2 kyrs BP
print("
The Martinez-Garcia Fe flux record (MarFe) spans from ", tmin MG, " to ", tmax M
G, " kyrs BP.")
MG hr = MG binned postmpt hr500
tmin MG hr = MG hr.indices[1].value # -3999 kyrs BP (4 Myrs)
tmax MG hr = MG hr.indices[end].value # -2 kyrs BP
print("
The higher resolution part of the MG record spans from ", tmin MG hr, " to ", tm
ax MG hr, " kyrs BP.")
```

The LR04 d180 record spans from -5320.0 to 0.0 kyrs BP.

The La2004 insolation time series (Ins) spans from -5000.0 to -0.0 k vrs BP.

Spratt Lisiecki GSL stack (SpraSL) spans from -797.0 to -1.0 kyrs B P.

The Elderfield GSL record (EldSL) spans from -1574.0 to -8.0 kyrs B P.

The Grant sea level record (GraSL) spans from -491.0 to -1.0 kyrs B P.

The Rohling sea level record (RohSL) spans from -5329.0 to -1.0 kyrs BP.

The Bereiter pCO2 record (BerCO2) spans from -803.0 to -2.0 kyrs BP. The Chalk pCO2 record (ChaCO2) spans from -1240.0 to -1092.0 kyrs BP.

The Lambert dust record (IceDust) spans from -799.0 to -13.0 kyrs B P.

The Martinez-Garcia Fe flux record (MarFe) spans from -3999.0 to -2.0 kyrs BP.

The higher resolution part of the MG record spans from -800.0 to -0.5 kyrs BP.

Plot the time series

(Move to NB1?)

Define the plots of binned resampled time series that we will use in analyse

In [46]:

```
####### Plot the La2004 insolation time series
plot La2004 =
plot(La2004 t fullength, La2004 insol65N fullength,
    color = :orange,
    label = "Ins
    xlabel = "Time [kyrs BP]",
    ylabel = L"Solar \ flux \ [W/m^{2}]",
    legend = :right,
    grid = false,
    size = (1000, 200)
####### plot the BinnedResampled LR04 time series
ts = LR04 binned fullength fullageunc
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the time series median in each bin (0.5 quantile), and the confidence
e interval we want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upperq = quantile.(ts.values, 0.975) .- bin median
bin lowerq = bin median .- quantile.(ts.values, 0.025)
;
plot LR04 =
plot(binmidpoints ts,
    bin median,
    ribbon = (bin lowerg, bin upperg),
    fillalpha = 0.3,
    color = :black,
    label = "LR04
    xlabel = "Time [kyrs BP]",
    ylabel = L"\delta{18}0 \ [\perthousand]",
    grid = false,
    yflip = true,
    size = (1000, 200),
    legend = :right
)
######### plot the Spratt & Lisiecki binned resampled GSL time series
ts = SL binned fullength noageunc
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot SL =
plot(binmidpoints ts, bin median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :darkblue,
    label =
            "SpraSL",
```

```
xlabel = "Time [kyrs BP]",
   ylabel = "GSL [m]",
   grid = false,
   size = (1000, 200),
    legend = :bottomleft
######## plot the Elderfield binned resampled GSL time series
ts = E binned fullength ageunc
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot E =
plot(binmidpoints ts, bin median,
   ribbon = (bin lower, bin upper),
   fillalpha = 0.3,
   color = :royalblue,
   label = "EldSL ",
   xlabel = "Time [kyrs BP]",
   ylabel = "GSL [m]",
   grid = false,
   size = (1000, 200),
   legend = :bottomleft
######## plot the Rohling binned resampled GSL time series
ts = R binned_full
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot R =
plot(binmidpoints ts, bin median,
   ribbon = (bin lower, bin upper),
   fillalpha = 0.3,
   color = :skyblue,
   label = "RohSL ",
   xlabel = "Time [kyrs BP]",
   ylabel = "GSL [m]",
   grid = false,
   size = (1000, 200),
   legend = :right
######## plot the Grant binned resampled GSL time series
ts = G binned fullength
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
```

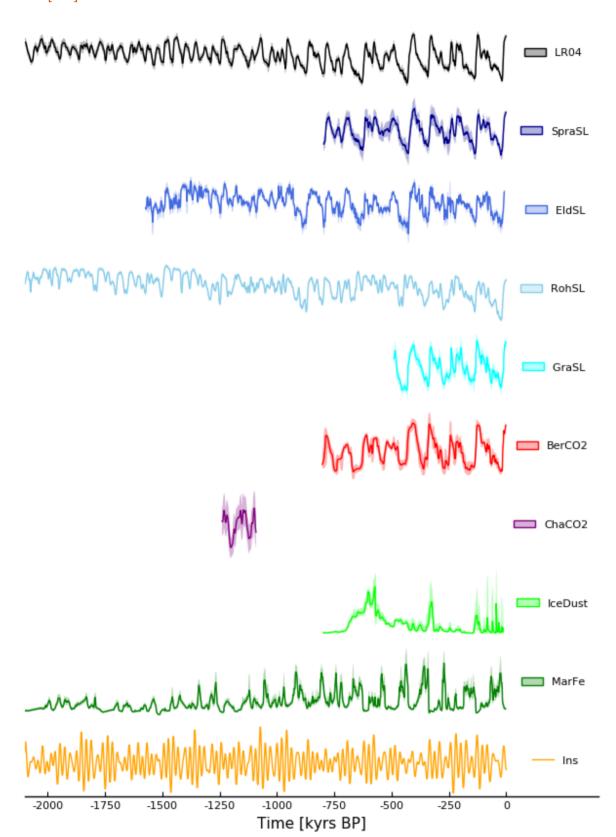
```
bin lower = bin median .- quantile.(ts.values, 0.025)
plot G =
plot(binmidpoints ts, bin median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :cyan,
    label = "GraSL ",
    xlabel = "Time [kyrs BP]",
    ylabel = "GSL [m]",
    grid = false,
    size = (1000, 200),
    legend = :bottomleft
# plot the binned resampled Bereiter CO2 time series
ts = B binned fullength ageunc
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot B =
plot(binmidpoints_ts, bin_median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :red,
    label = "BerCO2",
    xlabel = "Time [kyrs BP]",
    ylabel = L"pCO{2} \setminus [ppmv]",
    grid = false,
    size = (1000, 200),
    legend = :bottomleft
# plot the binned resampled Bereiter CO2 time series - version without age uncer
ts = B binned fullength noageuncEDC
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot B edc =
plot(binmidpoints ts, bin median,
    ribbon = (bin_lower, bin_upper),
    fillalpha = 0.3,
    color = :pink,
    label = "BerCO2 - without age model uncertainty",
    xlabel = "Time [kyrs BP]",
    ylabel = L"pCO{2} \setminus [ppmv]",
    grid = false,
```

```
size = (1000, 200),
    legend = :left
####### plot the binned resampled Chalk CO2 time series
ts = C binned
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot C =
plot(binmidpoints ts, bin median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :purple,
    label = "ChaCO2",
xlabel = "Time [kyrs BP]",
    ylabel = L"pCO {2} \setminus [ppmv]",
    grid = false,
    size = (1000, 200),
    legend = :bottomleft
####### Plot the two pCO2 time series in one figure
plot pCO2 =
plot(xlabel = "Time [kyrs BP]", ylabel = L"pCO{2} \ [ppmv]", grid = false, size
= (1000,200), legend = :bottomleft)
# Bereiter CO2
ts = B binned fullength ageunc
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot!(binmidpoints ts, bin median, ribbon = (bin lower, bin upper), fillalpha =
0.3, color = :red, label = "BerCO2")
# Chalk pCO2
ts = C binned
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot!(binmidpoints ts, bin median, ribbon = (bin lower, bin upper), fillalpha =
0.3, color = :purple, label = "ChaCO2")
####### plot the binned resampled Lambert dust time series
ts = L binned full
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
```

```
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot L =
plot(binmidpoints ts, bin median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :lime,
    label = "IceDust",
    xlabel = "Time [kyrs BP]",
    ylabel = "Dust conc. [\mu q/kq]",
    grid = false,
    size = (1000, 200),
    legend = :topleft
    )
####### plot the binned resampled Lambert dust time series with reduced age un
certainty
ts = L binned full EDC
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot L edc =
plot(binmidpoints ts, bin median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :olive,
    label = "IceDust - age uncertainty between gas and ice",
    xlabel = "Time [kyrs BP]",
    ylabel = "Dust conc. [\mu g/kg]",
    grid = false,
    size = (1000, 200),
    legend = :left
####### plot the binned resampled Martinez-Garcia dust time series
ts = MG binned fullength
binmidpoints ts =[ts.indices[i].value for i in 1:length(ts.indices)]
# computing the median in each bin (0.5 quantile), and the confidence interval w
e want to use (95%)
bin median = quantile.(ts.values, 0.5)
bin upper = quantile.(ts.values, 0.975) .- bin median
bin lower = bin median .- quantile.(ts.values, 0.025)
plot MG =
plot(binmidpoints_ts, bin_median,
    ribbon = (bin lower, bin upper),
    fillalpha = 0.3,
    color = :green,
    label = "MarFe",
```

```
xlabel = "Time [kyrs BP]",
   ylabel = string("Fe flux ", L"[g/m^{2}/yr]"),
   grid = false,
   size = (1000, 200),
   legend = :topleft)
po1 = plot(plot LR04, xlabel = "", xaxis = :off)
po2 = plot(plot_SL, xlabel = "", xaxis = :off)
po3 = plot(plot_E, xlabel = "", xaxis = :off)
po4 = plot(plot_R, xlabel = "", xaxis = :off)
po5 = plot(plot G, xlabel = "", xaxis = :off)
#po6 = plot(plot pCO2, xlabel = "", xaxis = :off)
po6 = plot(plot_B, xlabel = "", xaxis = :off)
po7 = plot(plot_C, xlabel = "", xaxis = :off)
po8 = plot(plot L, xlabel = "", xaxis = :off)
po9 = plot(plot MG, xlabel = "", xaxis = :off)
po10 = plot(plot La2004, xlabel = "Time [kyrs BP]", xaxis = :on)
#p highlight MPT = plot(Shape([-1250, -1250, -700, -700], [-100, 100, 100, -100]), fill
alpha = 0.1, color = :red, label = "MPT", ylims = (-10,10),
        xlabel = "Time [kyrs BP]", xaxis = :on)
po alltimeseries = plot(po1, po2, po3, po4, po5, po6, po7, po8, po9, po10, #p hi
ghlight MPT,
   size = (72*8.27, 72*11.69), # A4 (72 dots/inch)
   layout = grid(10,1),
   xlims = (-2100, 400), xticks = (-5000:250:0),
   ylabel = "", yaxis = :off,
   legend = :right, bg legend = :transparent,
   grid = :off,
   #Shape([-1250,-1250,-700,-700],[-100,100,-100]), fillalpha = 0.1, color
 = :red, label = "MPT",
                            # DOESN'T WORK
#savefiq("../../figurar/BinnedTimeseries/p all ts A4.pdf")
```

Out[46]:



In [1]:

```
# Experiments with layout - highlight
p highlight MPT = plot(Shape([-1250, -1250, -700, -700], [-100, 100, 100, -100]), filla
lpha = 0.1, color = :red, label = "MPT", ylims = (-10,10),
        xlabel = "Time [kyrs BP]", xaxis = :on)
# d180 / GSL
po1 = plot(plot_LR04, xlabel = "", xaxis = :off, grid = true, xtickfont = false)
po2 = plot(plot_SL, xlabel = "", xaxis = :off, grid = true, xtickfont = false)
po3 = plot(plot_E, xlabel = "", xaxis = :off, grid = true, xtickfont = false)
po4 = plot(plot_R, xlabel = "", xaxis = :off, grid = true, xtickfont = false)
po5 = plot(plot G, xlabel = "", xaxis = :off, grid = true, xtickfont = false)
# CO2
#po6 = plot(plot pCO2, xlabel = "", xaxis = :off)
po6 = plot(plot_B, xlabel = "", xaxis = :off, grid = true, xtickfont = false)
po6_edc = plot(plot_B_edc, xlabel = "", xaxis = :off, grid = true, xtickfont = f
po7 = plot(plot C, xlabel = "", xaxis = :off, grid = true, xtickfont = false)
# dust
po8 = plot(plot L, xlabel = "", xaxis = :off, grid = true, xtickfont = false)
po8 edc = plot(plot L edc, xlabel = "", xaxis = :off, grid = true, xtickfont = f
alse)
po9 = plot(plot MG, xlabel = "", xaxis = :off, grid = true, xtickfont = false,)
po10 = plot(plot La2004, xlabel = "", xaxis = :off, grid = true, xtickfont = fal
p xaxis = plot(xlabel = "Time [kyrs BP]", xaxis = :on, xtickfont = 8, yaxis = :o
ff,)
l = @layout [a;b;c;d;e;f;g;h;i;j;k;l; m{0.01h}]
#po alltimeseries =
plot(po1, po2, po3, po4, po5, po7, po6, po6_edc, po8, po8_edc, po9, po10, p_xaxi
s, #p highlight MPT,
   size = (72*8.27, 72*11.69), # A4 (72 dots/inch)
   layout = 1,
   xaxis = :on, xlims = (-5500, 400), xticks = (-5000:250:0),
   yaxis = :off, ylabel = "",
   bg legend = :transparent, #legend = :left,
   grid = :off,
   #Shape([-1250,-1250,-700,-700],[-100,100,-100]), fillalpha = 0.1, color
 = :red, label = "MPT",
                              # DOESN'T WORK
#savefig("../../figurar/BinnedTimeseries/p alltimeseries 2000 Appendix.pdf")
```

UndefVarError: Shape not defined

Stacktrace:

[1] top-level scope at In[1]:1

End of Notebook 1			

Next steps:

Now that the datasets are binned to a regular time grid, we only need to select a common time interval between records to run our analyses. This is done in the NBRs.

Outline for our next notebooks:

NB2) In notebook 2, we thoroughly go through all the steps in our analysis. We use the example of a synthetic autoregressive system where we already know the causal coupling, and how the predictive asymmetry is estimated between timeseries data. This notebook shows all the steps in the analysis with abundant comments and explanatory figures, and aims to give a general understanding of what is done in the following notebooks (NB3 and NBRs).

(After synthetic time series are represented, include note on reflections/challenges for using real-world empirical time series.)

NB3) In notebook 3, is the *toolbox* we prepare to run our analysis. The wrangled data that we have prepared in this notebook (NB1) is read in to easily be included in the next notebooks (NBRs). We also write a function that synthesize the code for computing the predictive asymmetry (detailed in NB2). This function allows us to run the analysis between the many time series more efficiently.

NBRs) The results notebooks (NBR) is where we run the analyses between our time series data and plot our results. By including NB3 in the beginning of the notebook, we include the wrangled time series data (prepared in this notebook NB1) and the functions to run the analyses. We make a separate NBR for every grid (time interval) we will run analysis over. Each time grid is constrained primarily by the resolution and time span of the available records. (ice core records, for example, only cover the last 800 kyrs). We also define time span constraints (*windows*) on time grids that allow us to check if dynamical coupling differs over the hypothesized periods post-MPT (after 700 kyrs BP), syn-MPT (700-1250 kyrs BP) and pre-MPT (before 1.25 Myrs BP). The NBRs are named after the time grid they run analyses over. The grids we will use are:

- post-MPT
 - postMPT 500 [-492:-12] (ca 0-500 kyrs BP, analysis delimited by the Grant record) DONE
 - maybe postMPT 700 (ca. 0-700 kyrs BP)
- svn-MPT
 - synMPT [-1250:-1080] (1.080-1.250 Myrs BP, analysis delimited by the Chalk record) DONE
- pre MPT
 - preMPT 1500 [-1500:1:-1250] (1.250-1.500 Myrs BP, delimited by the Elderfield record) DONE
 - maybe preMPT_4000 [-4000:1:-1250] (1.250-4.000 Myrs BP, Rohling record only available proxy for gsl)
- (full grids)
 - > What would we want to say with these?
 - full 800 (0-800 kyrs BP) almost done
 - all records (except Chalk)

If we were to prioritize between postMPT_700 and full_800 (same notebook), what would be more useful results?

- full_1500 (0-1500 kyrs BP)
 - Elderfield GSL, Rohling GSL, La2004 insolation, Martinez-García Fe dust.

- maybe full_4000 (0-4000 kyrs BP)
 - Rohling GSL, La2004 insolation, Martinez-García Fe dust.

is this interesting? Given the Rohling record is so-so. Maybe if we can cut out and interpolate over the sapropelic intervals, and the method still has statistical power. If not, I think better to leave it be.

In []:	
In []:	